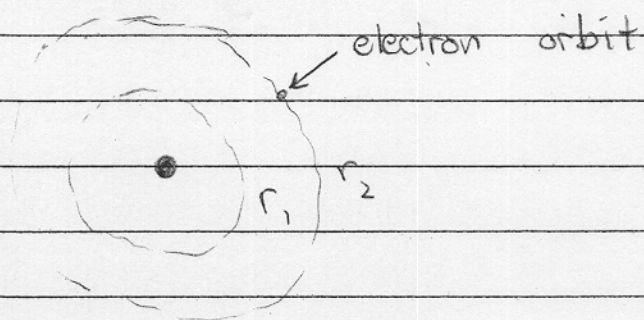


Last Time - Bohr Model



- The electrons are attracted to the proton

$$L = m v_n r_n = n \hbar \quad n = 1, 2, 3, \dots$$

- They can determine the radius of the orbit

$$F_c = m_e \frac{v_n^2}{r}$$

$$\frac{1}{4\pi\epsilon_0} \frac{e^2}{r_n^2} = m_e \frac{v_n^2}{r_n}$$

$$\frac{e^2}{4\pi\epsilon_0} \frac{1}{m v_n^2} = r_n$$

So $L = m v_n r_n = n \hbar$

$$m v_n \cdot \left(\frac{e^2}{4\pi\epsilon_0} \frac{1}{m v_n^2} \right) = n \hbar$$

Simplify

$$V_n = \frac{e^2}{4\pi\epsilon_0 h} \frac{1}{n}$$

$$\frac{V_n}{c} = \left(\frac{e^2}{4\pi\epsilon_0 hc} \right) \cdot \frac{1}{n} = \frac{\alpha}{n} \Rightarrow$$

$$\frac{V_1}{c} = \frac{1}{137} \Leftarrow \text{The electrons are non-relativistic to a good approximation}$$

The

$$r_n = \frac{e^2}{4\pi\epsilon_0} \frac{1}{m_e v_n^2} = \frac{e^2}{4\pi\epsilon_0 hc} \cdot \frac{hc}{m_e c^2 \left(\frac{v_n}{c}\right)^2}$$

$$r = \frac{hc}{m_e c^2 \alpha^2} \frac{1}{n^2}$$

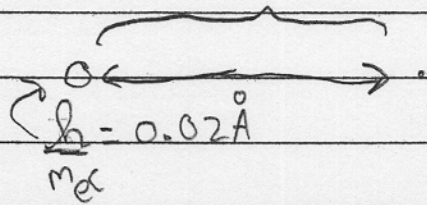
$$r_n = n^2 \left[\frac{hc}{m_e c^2 \alpha} \right] \equiv n^2 a_0 \quad \left[\frac{1}{m} \right]$$

$$a_0 \equiv \frac{hc}{m_e c^2 \alpha} = \frac{137}{2\pi} \left(\frac{hc}{m_e c^2} \right) \quad \text{up to } 2\pi \text{ this is the electron Compton wavelength}$$

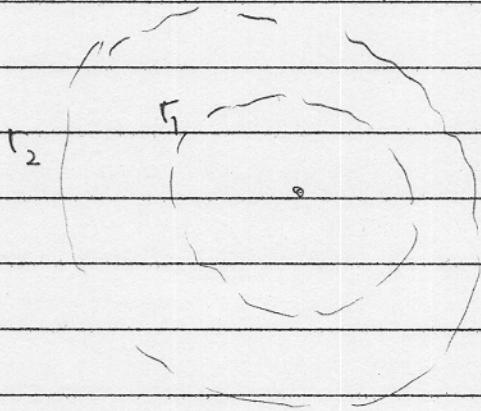
$$a_0 = 137 \cdot \left(\frac{hc}{m_e c^2} \right) = 137 \left(\frac{197 \text{ eV} \cdot \text{nm}}{511000 \text{ eV}} \right) = 0.053 \text{ nm}$$

$$a_0 = 0.53 \text{ \AA}$$

Picture $a_0 = 0.5 \text{ \AA} = \frac{h}{m_e c \alpha}$



Now we want to determine the energies



$$PE_n = -\frac{1}{4\pi\epsilon_0} \frac{e^2}{r_n} = -\left(\frac{e^2}{4\pi\epsilon_0 hc}\right) \frac{hc}{a_0 n^2}$$

$$= -\left(\frac{\alpha hc}{a_0}\right) \frac{1}{n^2}$$

$$= -\left(\frac{1}{137} \cdot \frac{1970 \text{ eV \AA}}{0.5 \text{ \AA}}\right) \frac{1}{n^2} = -\frac{27.2 \text{ eV}}{n^2}$$

For the KE

$$KE = \frac{1}{2} m v^2 = \frac{1}{2} m \left(\frac{\alpha c}{n} \right)^2 = \frac{1}{2} m_e c^2 \frac{\alpha^2}{n^2} = 13.6 \text{ eV} \frac{1}{n^2}$$

Why?

Half of |PE|

$$\frac{e^2}{4\pi\epsilon_0} \frac{1}{r^2} = \frac{m v^2}{r}$$

$$\frac{e^2}{4\pi\epsilon_0} \frac{1}{r} = m v^2$$

$$\underbrace{-PE} = \underbrace{2 KE}$$

$$-\frac{1}{2} PE = KE$$

Energy Summary

$$E_{\text{Tot } n} = PE + KE$$

$$= -\frac{1}{4\pi\epsilon_0} \frac{e^2}{2a_0} \left(\frac{1}{n^2} \right) + \frac{1}{2} \left(\frac{1}{4\pi\epsilon_0} \frac{e^2}{a_0} \right) \frac{1}{n^2}$$

$$= -\frac{1}{4\pi\epsilon_0} \left(\frac{e^2}{2a_0} \right) \left[\frac{1}{n^2} \right] \equiv \boxed{-R_\infty \frac{1}{n^2} = E_n}$$

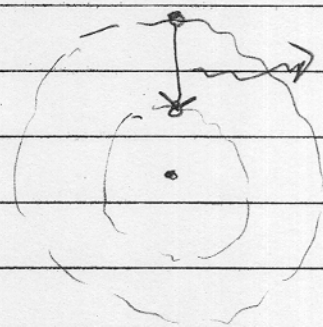
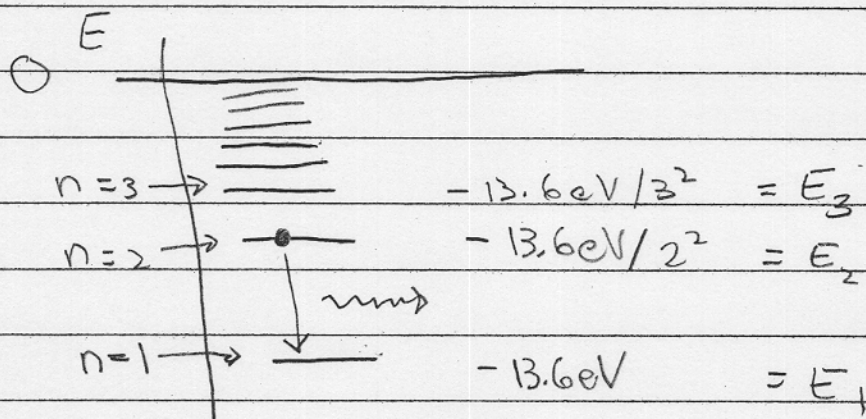
energy of n-th orbit

$$R_\infty = \frac{-e^2}{4\pi\epsilon_0} \frac{e^2}{2a_0} = -13.6 \text{ eV} \leftarrow \text{Energy of lowest orbit}$$

$$E_1 = -13.6 \text{ eV}$$

$$E_1 = \underbrace{-27.2 \text{ eV}}_{PE_1} + \underbrace{13.6 \text{ eV}}_{KE_1}$$

Energy Diagrams



- Consider an atom which transitions from the $n=2$ orbit to the $n=1$ orbit emitting a photon

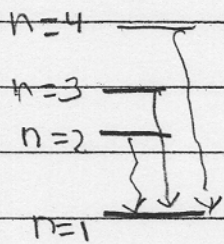
$$E_i = E_f + E_\gamma \quad \leftarrow \text{energy conservation}$$

$$E_\gamma = E_i - E_f$$

$$= \left(-\frac{R_\infty}{2^2} \right) - \left(-\frac{R_\infty}{1^2} \right) = R_\infty \left[\frac{1}{1^2} - \frac{1}{2^2} \right] = R_\infty \cdot \frac{3}{4}$$

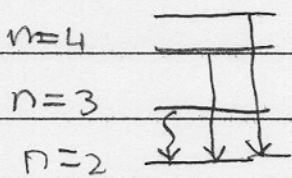
$$E_\gamma = 13.6 \text{ eV} \cdot \frac{3}{4} = 10.2 \text{ eV} \quad \leftarrow \lambda = 124 \text{ nm deep UV}$$

Others:



(Lyman Series)

Still Others



(Balmer Series)

Determine the wave lengths of the Balmer Lines

$$E_{\gamma} = E_i - E_f$$

$$E_{\gamma}^{n \rightarrow 2} = \left(-\frac{R_{\infty}}{n^2} \right) - \left(-\frac{R_{\infty}}{2^2} \right) = R_{\infty} \left[\frac{1}{2^2} - \frac{1}{n^2} \right] = R_{\infty} \left[\frac{n^2 - 2^2}{2n^2} \right]$$

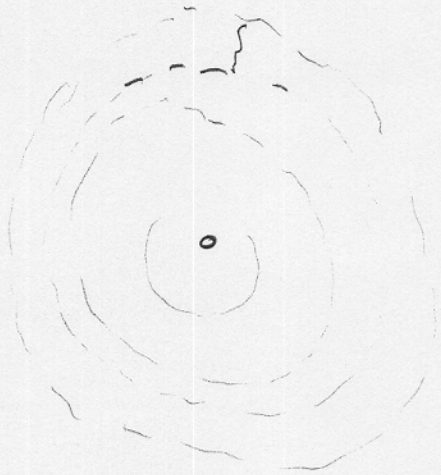
$$E_{\gamma}^{n \rightarrow 2} = \frac{hc}{\lambda_n} \Rightarrow \lambda_n = \frac{hc}{E_{\gamma}^{n \rightarrow 2}} = \frac{hc}{R_{\infty} \left(\frac{n^2 - 2^2}{2n^2} \right)}$$

$$\lambda_n = \frac{2hc}{R_{\infty}} \cdot \left(\frac{n^2}{n^2 - 2^2} \right)$$

Explains
empirical
rules

$$E = \left(\frac{2 \cdot 1240 \text{ eV nm}}{13.6 \text{ eV}} \right) \left(\frac{n^2}{n^2 - 2^2} \right) = (363 \text{ nm}) \left(\frac{n^2}{n^2 - 2^2} \right)$$

The Correspondence Principle



$$r_n = n^2 a_0$$

$$\frac{r_{n+1} - r_n}{r_n} \approx \frac{2na_0}{n^2 a_0} = \frac{2}{n}$$

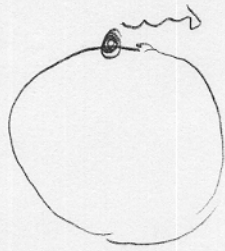
- When n becomes large $r \rightarrow \infty$ and the spacing gets smaller (in a relative sense)
- So we should have

$$\lim_{n \rightarrow \infty} \text{quantum physics} = \text{classical physics}$$

- This is enough to say that if photons come in discrete units
 $E = hf = \hbar\omega$

Then the angular momentum changes in discrete steps

When



Facts

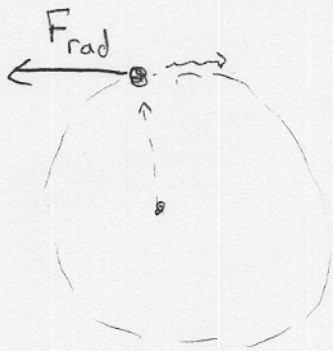
$$\textcircled{1} \quad v = R\omega_e$$

$\textcircled{2}$ For an electron going around in a circle (classically)

Maxwell equations

say that

$$\omega_{\text{photon}} = \omega_e \equiv \omega$$



$$\frac{\Delta E_{\text{lost}}}{\Delta t} = F_{\text{rad}} \cdot v$$

$$\frac{\Delta E_{\text{lost}}}{\Delta t} = F_{\text{rad}} R \omega$$

$$F_{\text{rad}} R = \tau_{\text{or}} = \frac{\Delta L}{\Delta t} \quad \text{from radiation}$$

$$\frac{\Delta E_{\text{lost}}}{\Delta t} = \frac{\Delta L_{\text{lost}}}{\Delta t} \omega$$

$$\Delta E = \Delta L \cdot \omega$$

Now Planck says Energy is emitted in discrete units

$$n h \omega_{\text{photon}} = n h \omega$$

$$n h \omega = \Delta L \omega$$

$$n h = \Delta L$$

change in L during emission of one light packet

Says L comes in discrete + units

De Broglie ~ 1920

- Bohr theory clearly inadequate
 - Want something like: given electron at time one where is electron at time #2
 - Only limited success at describing multi-electron atoms. Just specifies L_{TOT} . not enough to describe each individual electron

Think about light:

$$E = h \frac{c}{\lambda}$$

$$E = cp$$

$$cp = h \frac{c}{\lambda}$$

$$p = \frac{h}{\lambda} \Rightarrow \boxed{\lambda = \frac{h}{p}}$$

Also remember the Bohr model: $\frac{v}{c} = \alpha \quad v = \alpha c$

$$p = m \cdot v = m \alpha c = m \left(\frac{ke^2}{\hbar c} \right) \cdot c = \hbar \underbrace{\left(\frac{mke^2}{\hbar^2} \right)}_{\frac{1}{a_0}}$$

$$p = \frac{\hbar}{a_0}$$

Then

$$a_0 = \frac{h}{p}$$

call it wave length

$$\text{so } \underbrace{\lambda}_{2\pi a_0} = \frac{h}{p}$$

Then Bohr quantization condition

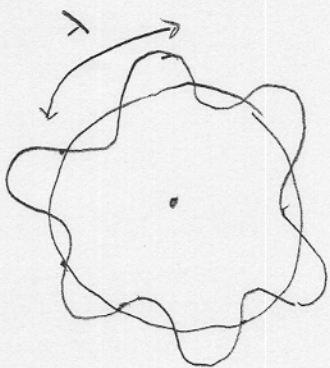
$$\overbrace{m v r} = n \frac{h}{2\pi}$$

$$2\pi \cdot p r = n h$$

$$2\pi \frac{h}{\lambda} r = n h$$

$$\boxed{2\pi r = n \lambda}$$

Picture



have to have complete wave lengths going around the circle

$$\lambda = \frac{h}{p}$$

$$p = \frac{h}{\lambda} = \frac{h}{2\pi} \frac{2\pi}{\lambda} = \boxed{\frac{h}{2\pi} k = p}$$