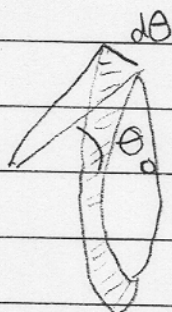


Problem 4.8

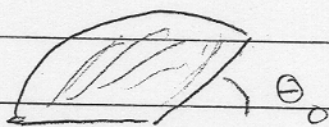
$$dN = [dtL] \frac{d\sigma}{d\Omega} d\Omega$$

$$dN_{\theta} = [N_{\alpha} \rho_{Au} t_{foil}] \cdot \frac{D^2}{16 \sin^4(\theta/2)} \cdot 2\pi \sin\theta d\theta$$



Then dN is the number scattered into a ring of size $d\theta$. To find the total number which scatter with angle greater than θ_0 .

$$N = \sum dN_{\theta} = \int_{\theta_0}^{\pi} [N_{\alpha} \rho_{Au} t_{foil}] \frac{D^2}{16 \sin^4\theta} 2\pi \sin\theta d\theta$$



$$\int \frac{\sin\theta}{\sin^4(\theta/2)} = \int \frac{\sin(2\theta/2)}{\sin^4(\theta/2)} = \int \frac{2 \sin\theta/2 \cos\theta/2}{\sin^4\theta/2} = \frac{-2}{\sin^2\theta/2}$$

S_0

$$N = [N_\alpha P_{Au} t_{foil}] \frac{D^2}{16} \cdot 2\pi \int_{\theta_0}^{\pi} \left[\frac{-2}{\sin^2 \theta/2} \right] \pi$$

$$= [] \frac{D^2}{16} \cdot 4\pi \left[\frac{-1 + 1}{\sin^2 \theta_0/2} \right]$$

$$N = [] \frac{D^2}{4} \pi \left[\frac{\overbrace{1 - \sin^2 \theta_0/2}^{\cos^2 \theta_0/2}}{\sin^2 \theta_0/2} \right] = [] \frac{\pi D^2}{4} \cot^2(\theta_0/2)$$

with $D = \frac{Z Z_\alpha e^2}{4\pi\epsilon_0 (LMV^2)}$

$$D^2 = 4 \left(\frac{Z Z_\alpha e^2}{4\pi\epsilon_0 mV^2} \right)^2 \Rightarrow \frac{D^2}{4} = \left(\frac{Z Z_\alpha e^2}{4\pi\epsilon_0 mV^2} \right)^2$$

S_0

$$N = [N_\alpha P_{Au} t_{foil}] \left(\frac{Z Z_\alpha e^2}{4\pi\epsilon_0 mV^2} \right)^2 \cdot \pi \cos^2(\theta_0/2) \quad \checkmark$$

4.9

$$N_{\theta > 60} = [N_{\alpha} \rho t] \frac{\pi D^2}{4} \cot^2(60/2)$$

$$\cot^2 60/2 = \frac{\cos^2 30}{\sin^2 30} = \frac{(\sqrt{3}/2)^2}{(1/2)^2} = 3$$

$$D = \alpha \frac{Z \cdot z_p \hbar c}{(\frac{1}{2} m v^2)} = \frac{1}{137} \cdot 79 \cdot 1 \cdot 197 \text{ MeV fm} = 18.9 \text{ fm}$$

6 MeV

$$D^2 = 358 \text{ fm}^2 = 3.58 \text{ barns} = 3.58 \times 10^{-24} \text{ cm}^2$$

$$\frac{N_{\theta > 60}}{N_{\alpha}} = 2 \times 10^{-5} = \text{fraction with } \theta > 60^\circ$$

$$t_{\text{foil}} = \frac{(N_{\theta > 60})}{N_{\alpha} \rho t} \frac{1}{\rho_{\text{Au}} \frac{\pi D^2}{4} \cdot 3} \quad (\cot 60/2)^2$$

ρ_{Au} from next problem, $\rho_{\text{Au}} = 0.58 \times 10^{23} \frac{1}{\text{cm}^3}$

$$t_{\text{foil}} = (2 \times 10^{-5}) \frac{1}{0.58 \times 10^{23} \frac{1}{\text{cm}^3} \cdot \frac{(3\pi)}{4} \cdot 3.58 \times 10^{-24} \text{ cm}^2}$$

$$t_{\text{foil}} = 4 \times 10^{-5} \text{ cm} = 0.4 \mu\text{m} \sim \text{"tenths of a micron"}$$

4.10

The number of particles which are scattered into a detector which spans a solid angle $d\Omega$ is

$$dN = [N_{\alpha} \rho_{Au} t_{foil}] \frac{D^2}{16 \sin^4(\theta/2)} d\Omega$$

Units

$$[N_{\alpha} \rho_{Au} t_{foil}] = \frac{\#}{\text{area}}$$

$$\left[\frac{D^2}{16 \sin^4(\theta/2)} \right] = \text{area}$$

$$d\Omega = \text{unit less}$$

A unit of area which is commonly used is

$$\text{barns} = 10^{-24} \text{ cm}^2$$

$$\rho_{Au} = \frac{\text{Number}}{\text{Volume}} = \frac{1}{m_{Au}} \frac{\text{mass}}{\text{Volume}}$$

↑
mass of a gold nucleus

1 Avogadro's # of protons weighs \approx 1 gram

1 Avogadro's of gold \approx 197 g

So

$$m_{Au} = \frac{197 \text{ g}}{6 \times 10^{23}}$$

$$\rho_{Au} = \frac{19.3 \text{ g/cm}^3}{\frac{197 \text{ g}}{6 \times 10^{23}}} = 0.58 \times 10^{23} \frac{1}{\text{cm}^3}$$

$$t_{\text{foil}} = 0.1 \mu\text{m} = 1 \times 10^{-5} \text{ cm}$$

$$N_{\alpha} = \frac{10^4}{\text{Sec}} \times \underbrace{3600 \text{ s}}_{1 \text{ hour}}$$

$$\cancel{N_{\alpha}} [N_{\alpha} \rho_{Au} t_{\text{foil}}] = 0.211 \times 10^{26} \frac{1}{\text{cm}^2} = \boxed{21.1 \frac{1}{\text{barn}}}$$

The distance of closest approach

$$D = \frac{Z_{Au} Z_{\alpha} e^2}{4\pi\epsilon_0 \left(\frac{1}{2}mv^2\right)} = \alpha \frac{Z_{Au} Z_{\alpha} \hbar c}{\frac{1}{2}mv^2}$$

$$D = \frac{1}{137} \frac{79.2 \cdot (197 \text{ MeV fm})}{5.3 \text{ MeV}}$$

$$D = 42.86 \text{ fm} \quad \text{about } 4 \times \text{ larger than the Au nucleus radius}$$

$$D^2 = 1836 \text{ fm}^2 = \boxed{18.3 \text{ barns} = D^2}$$

$$1 \text{ fm}^2 = 0.01 \text{ barn}$$

$$d\Omega = \frac{dA}{r^2} = \frac{1 \text{ cm}^2}{(10 \text{ cm})^2} = \boxed{0.01 = d\Omega}$$

Now:

$$dN = \frac{[N_{\alpha} \rho_{\text{Au}} t_{\text{foil}}]}{16 \sin^4(\theta/2)} \frac{1}{D^2} d\Omega \quad \text{for}$$

$$dN = \left(\frac{21.1}{\text{barn}} \right) (18.3 \text{ barn}) \frac{1}{16 \sin^4(10^\circ/2)} \cdot 0.01$$

$$dN_{10^\circ} = 4180 \quad \text{at } 10^\circ$$

At 45°

$$dN_{45^\circ} = dN_{10^\circ} \cdot \frac{\sin^4(10^\circ/2)}{\sin^4(45^\circ/2)} = 4180 \cdot \frac{\sin^4(10^\circ/2)}{\sin^4(45^\circ/2)} \approx 11$$

Problem 4.11

$$N \propto \rho Z^2$$

where ρ is the number density

$d \equiv$ is the mass density, 19.3 g/cm^3 for Au, and 8.9 g/cm^3 for copper

$\rho \propto \frac{d}{A}$ where A is the atomic number, 197 for Au and 63.6 for Cu. A is proportional to the mass of nucleus $\rho = \frac{d}{\text{mass of nucleus}}$

So $N \propto \frac{d Z^2}{A}$ $\propto d/A$

$$N \propto \frac{d Z^2}{A} \quad \text{constant \#}$$

$$N = \frac{K d Z^2}{A} \quad \text{so} \quad \frac{N_{\text{Cu}}}{N_{\text{Au}}} = \frac{d_{\text{Cu}} (Z_{\text{Cu}}/Z_{\text{Au}})^2}{d_{\text{Au}} (A_{\text{Cu}}/A_{\text{Au}})}$$

$$Z_{\text{Au}} \sqrt{\frac{N_{\text{Cu}}}{N_{\text{Au}}} \frac{d_{\text{Au}}}{d_{\text{Cu}}} \frac{A_{\text{Cu}}}{A_{\text{Au}}}} = Z_{\text{Cu}}$$

So with $N_{\text{Au}} = 4180$ and $N_{\text{Cu}} = 820$ we have:

$$Z_{\text{Cu}} = (79) \left(\frac{820}{4180} \cdot \frac{19.3 \text{ g/cm}^3}{8.9 \text{ g/cm}^3} \cdot \frac{63.6}{197} \right)^{1/2}$$

$$Z_{\text{Cu}} = 29$$

Agrees (w) periodic table

$$\textcircled{1} \quad \Delta t \mathcal{L} = \frac{N_p N_t}{A}$$

$$= N_p \left(\frac{N_t}{A \Delta x} \right) \Delta x$$

$$\Delta t \mathcal{L} = N_p \rho_{Au} t_{\text{foil}} \quad \text{and} \quad \mathcal{L} = \frac{\Delta N_p}{\Delta t} \rho_{Au} t_{\text{foil}}$$

$$\textcircled{2} \quad \sigma_{\text{Backward}} = \frac{N_{\text{backwards}}}{\Delta t \mathcal{L}}$$

$$\sigma_{\text{Backward}} = \frac{[N_p t] \pi \left(\frac{Z z e^2}{m v^2} \right)^2 \cdot \cot(90^\circ/2)}{[N_p t_{\text{foil}}]}$$

$$\sigma_{\text{back}} = \frac{\pi}{4} \left(\frac{Z z e^2}{m v^2} \right)^2 \cdot 1$$

$$\sigma_{\text{back}} = \frac{\pi D^2}{4}$$

$$\textcircled{3} \quad \frac{dN}{dt} = \int \frac{dN}{dt d\Omega} d\Omega = \int_0^\pi N_0 (1 + \cos^2 \theta) 2\pi \sin \theta d\theta$$



$$= N_0 2\pi \int_0^\pi \sin \theta d\theta + N_0 2\pi \int_0^\pi \cos^2 \theta \sin \theta d\theta$$

$$= N_0 [2\pi \cdot 2] + N_0 \frac{2\pi}{3} (-\cos^3 \theta) \Big|_0^\pi$$

$$\frac{dN}{dt} = 4\pi N_0 + \frac{4\pi N_0}{3}$$

$$a) \frac{dN}{dt} = N_0 \cdot \frac{4}{3} \cdot 4\pi = 0.33 \times 10^{21}$$

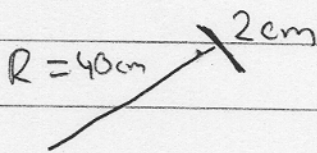
$$b) \frac{dN}{dt d\Omega} = \frac{dN}{dt} \cdot \frac{1}{\Omega_{\text{sphere}}} = \frac{0.33 \times 10^{21}}{4\pi} = 0.266 \times 10^{20} \frac{1}{s}$$

Problem 4

$$a) \quad \Delta\Omega = \frac{A}{r^2} = \frac{4 \text{ cm}^2}{(40 \text{ cm})^2} = 0.0025 \text{ str}$$

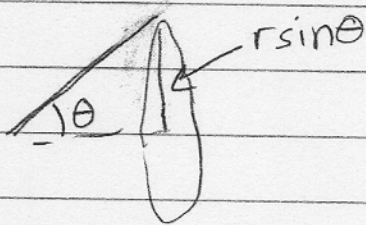
units steradian
↓
but its
↑
unitless

$$b) \quad \Delta\theta = \frac{2 \text{ cm}}{40 \text{ cm}} = \frac{1}{20} \text{ rad} = 2.86^\circ$$

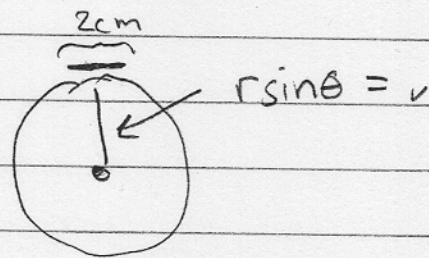


For $\Delta\phi$

side view



top view



$$\Delta\phi = \frac{l}{R \sin \theta} = \frac{2 \text{ cm}}{40 \text{ cm} \sin 30^\circ} = \frac{1}{10} \text{ rad} = 5.72^\circ$$

c) The Power scattered = 3 mW , each photon has

$$E_{\text{photon}} = h\nu = \frac{hc}{\lambda} = \frac{1240 \text{ eV nm}}{422.7 \text{ nm}} = 2.93 \text{ eV}$$

So the Number scattered is

$$\frac{dN}{dt} \text{scatt} = \frac{3 \text{ mW}}{2.93 \text{ eV}}$$
$$= 6.39 \times 10^{15} \text{ 1/s}$$

So since it is spread out uniformly over the whole sphere

$$\Omega = \frac{4\pi r^2}{r^2} = 4\pi$$

$$\frac{dN}{dt d\Omega} = 0.51 \times 10^{15} \text{ 1/s}$$

d) The number into the detector per time

$$\frac{\Delta N}{\Delta t} = \frac{dN}{dt d\Omega} \Delta\Omega$$

$$= 0.51 \times 10^{15} \frac{1}{s} \cdot (0.0025 \text{ str})$$

$$\frac{dN}{dt} = 1.27 \times 10^{12} \frac{1}{s}$$