## See "Technical note on computing radiation fields"

## Problem 1. (Optional) Electric field in the far field

If you get stuck check the notes online. The scalar and vector potential in the far field are

$$
\begin{align*}
\varphi(t, \boldsymbol{r}) & =\frac{1}{4 \pi r} \int d^{3} \boldsymbol{r}_{o} \rho\left(T, \boldsymbol{r}_{o}\right)  \tag{1}\\
\boldsymbol{A}(t, \boldsymbol{r}) & =\frac{1}{4 \pi r} \int d^{3} \boldsymbol{r}_{o} \boldsymbol{J}\left(T, \boldsymbol{r}_{o}\right) / c \tag{2}
\end{align*}
$$

where the retarded time $T=t-\left|\boldsymbol{r}-\boldsymbol{r}_{o}\right| / c$ in the far field is

$$
\begin{equation*}
T=t-r / c+\frac{\boldsymbol{n} \cdot \boldsymbol{r}_{o}}{c} \tag{3}
\end{equation*}
$$

The goal is to compute the electric field

$$
\begin{equation*}
\boldsymbol{E}(t, r)=-\frac{1}{c} \partial_{t} \boldsymbol{A}(t, \boldsymbol{r})-\nabla \varphi(t, \boldsymbol{r}) \tag{4}
\end{equation*}
$$

(a) (Optional) Consider the change of variable $t, \boldsymbol{r}_{o} \rightarrow T, \boldsymbol{r}_{o}$. Show that

$$
\begin{align*}
\frac{\partial}{\partial T} & =\frac{\partial}{\partial t}  \tag{5}\\
\left(\frac{\partial}{\partial \boldsymbol{r}_{o}}\right)_{T} & =\left(\frac{\partial}{\partial \boldsymbol{r}_{o}}\right)_{t}-\frac{\boldsymbol{n}}{c} \frac{\partial}{\partial t} \tag{6}
\end{align*}
$$

(b) (Optional) Compute

$$
\begin{equation*}
\left(\frac{\partial}{\partial t}+c \boldsymbol{n} \cdot \frac{\partial}{\partial \boldsymbol{r}}\right) T=0 \tag{7}
\end{equation*}
$$

You should find a simple result. Interpret the answer using the definition of $T$
$T \equiv$ the time when the light should be emitted from $\boldsymbol{r}_{o}$ to arrive at the observation point $(t, \boldsymbol{r})$.

How do you interpert the derivative:
(c) (Optional) Show that

It is the derivative moving with the light.

$$
\left(\frac{\partial}{\partial t}+c \boldsymbol{n} \cdot \frac{\partial}{\partial \boldsymbol{r}}\right)
$$ If you change the observation point ( $\mathrm{t}, \mathrm{r}$ ) to ( $\mathrm{t}+\mathrm{dt}, \mathrm{r}+\mathrm{c} n \mathrm{dr}$ ) the light (Bie emitted at (T,rO) will reach both of these points without changing T or rO .

$$
\begin{equation*}
\boldsymbol{E}=-\frac{1}{4 \pi r c^{2}} \int_{r_{o}} \frac{\partial \boldsymbol{J}\left(T, r_{o}\right)}{\partial T}+\frac{\boldsymbol{n}}{c} \frac{1}{4 \pi r} \int_{r_{o}} \frac{\partial \rho\left(T, r_{o}\right)}{\partial T} \tag{9}
\end{equation*}
$$

(d) (Optional) Use current conservation to express

$$
\begin{equation*}
\frac{\partial \rho\left(T, \boldsymbol{r}_{o}\right)}{\partial T}=-\left(\nabla_{r_{o}} \cdot \boldsymbol{J}\right)_{T}=-\left(\nabla_{r_{o}} \cdot \boldsymbol{J}\right)_{t}+\frac{\boldsymbol{n}}{c} \cdot \frac{\partial \boldsymbol{J}}{\partial T} \tag{10}
\end{equation*}
$$

where $\left(\nabla_{r_{o}} \cdot \boldsymbol{J}\right)_{t}$ denotes the divergence at fixed observation time
(e) (Optional) Conclude that only the transverse piece of the current to $\boldsymbol{n}$ contributes to the radiation field

$$
\begin{align*}
\boldsymbol{E} & =-\frac{1}{4 \pi r} \frac{1}{c^{2}} \int_{\boldsymbol{r}_{o}} \underbrace{\left[\partial_{t} \boldsymbol{J}-\boldsymbol{n}\left(\boldsymbol{n} \cdot \partial_{t} \boldsymbol{J}\right)\right]}_{\text {the part of } \partial_{t} J \text { transverse to } \boldsymbol{n}}  \tag{11}\\
& =\boldsymbol{n} \times\left[\frac{\boldsymbol{n}}{c} \times \frac{1}{4 \pi r} \int_{\boldsymbol{r}_{o}} \frac{1}{c} \frac{\partial \boldsymbol{J}\left(T, r_{o}\right)}{\partial T}\right] \tag{12}
\end{align*}
$$

## Problem 2. Dipole Fields

Consider a small ectric dipole with harmonic time dependence, $\boldsymbol{p}(t)=\boldsymbol{p}_{o} e^{-i \omega t}$. Recall that in homework 6 we determined the electric field through order $\omega^{2}$ in frequency using a quasistatic near field expansion

$$
\begin{equation*}
\boldsymbol{E}(t)=\frac{3 \boldsymbol{n}(\boldsymbol{n} \cdot \boldsymbol{p}(t))-\boldsymbol{p}(t)}{4 \pi r^{2}}-\frac{\boldsymbol{n}(\boldsymbol{n} \cdot \ddot{\boldsymbol{p}})+\ddot{\boldsymbol{p}}}{8 \pi r c^{2}} \tag{14}
\end{equation*}
$$

The purpose of the problem is to examine the transition to the far field, by computing the exact electric field as a function of radius.
(a) Define near and far field. Express your results in terms of the wave number $k=\omega / c$.
(b) Start from the exact expressions

$$
\begin{align*}
\varphi(t, \boldsymbol{r}) & =\int d^{3} \boldsymbol{r}_{o} \frac{\rho\left(T, \boldsymbol{r}_{o}\right)}{4 \pi\left|\boldsymbol{r}-\boldsymbol{r}_{o}\right|}  \tag{15}\\
\boldsymbol{A}(t, \boldsymbol{r}) & =\int d^{3} \boldsymbol{r}_{o} \frac{\boldsymbol{J}\left(T, \boldsymbol{r}_{o}\right) / c}{4 \pi\left|\boldsymbol{r}-\boldsymbol{r}_{o}\right|} \tag{16}
\end{align*}
$$

where

$$
\begin{equation*}
T=t-\frac{\left|\boldsymbol{r}-\boldsymbol{r}_{o}\right|}{c} \simeq t-\frac{r}{c}+\frac{\boldsymbol{n} \cdot \boldsymbol{r}_{o}}{c}, \tag{17}
\end{equation*}
$$

and assume harmonic time dependence of $\rho\left(t_{o}, \boldsymbol{r}_{o}\right)=\rho\left(\boldsymbol{r}_{o}\right) e^{-i \omega t_{o}}$ and $\boldsymbol{J}\left(t_{o}, r_{o}\right)=$ $\boldsymbol{J}\left(r_{o}\right) e^{-i \omega t}$, without making a far field expansion show that

$$
\begin{align*}
& \varphi(t, \boldsymbol{r})=\frac{e^{-i \omega t+i k r}}{4 \pi r^{2}} \boldsymbol{n} \cdot \boldsymbol{p}_{o}(1-i k r)  \tag{18}\\
& \boldsymbol{A}(t, \boldsymbol{r})=-i k \frac{e^{-\omega t+i k r}}{4 \pi r} \boldsymbol{p}_{o} \tag{19}
\end{align*}
$$

(c) Show by direct differentiation of the potentials $\boldsymbol{A}$ and $\varphi$ that in the far field you recover the result given in class

$$
\begin{equation*}
\boldsymbol{E}=\frac{k^{2} e^{-i \omega t+i k r}}{4 \pi r}\left[-\boldsymbol{n} \times \boldsymbol{n} \times \boldsymbol{p}_{o}\right] \tag{20}
\end{equation*}
$$

Comment on all qualitative features.
When differentiating, note carefully the contribution from $\boldsymbol{A}$ and $\varphi$, and how they conspire to make a field $\boldsymbol{E}$ which is transverse to $\boldsymbol{n}$.
(d) Show that in general

$$
\begin{equation*}
\boldsymbol{E}(t, r)=e^{-i \omega t+i k r}\left[\frac{3 \boldsymbol{n}\left(\boldsymbol{n} \cdot \boldsymbol{p}_{o}\right)-\boldsymbol{p}_{o}}{4 \pi r^{3}}-i k \frac{3 \boldsymbol{n}\left(\boldsymbol{n} \cdot \boldsymbol{p}_{o}\right)-\boldsymbol{p}_{o}}{4 \pi r^{2}}+k^{2} \frac{-\boldsymbol{n} \times \boldsymbol{n} \times \boldsymbol{p}_{o}}{4 \pi r}\right] \tag{21}
\end{equation*}
$$

Comment on all qualitative features. Write Eq. (21) in coordinate space expressed in terms of $\boldsymbol{p}\left(t_{e}\right), \dot{\boldsymbol{p}}\left(t_{e}\right)$, and $\ddot{\boldsymbol{p}}\left(t_{e}\right)$ with $t_{e}=t-r / c$ (compare to the typed course notes sec. 11.3), and show consistency of this result with the near field result derived with quasi-statics.

Dipole Fields
(a) We mean $r \ll \omega / c$ for near field and $r \gg w / c$. for far field.
b) Then

$$
\begin{aligned}
& \varphi(t, r)=\int d^{3} r_{0} \frac{\rho\left(T, r_{0}\right)}{4 \pi\left|\vec{r}-\vec{r}_{0}\right|} \text { where } \\
& T=t-\frac{\left|\vec{r}-\vec{r}_{0}\right|}{c}
\end{aligned}
$$

- Imagine a dipole
- $\vec{r}$
$(-) \vec{r}_{0} \vec{r}_{0}$ integrates over this
Then

$$
\begin{aligned}
& \text { - }\left|\vec{r}-\vec{r}_{0}\right| \simeq r-\vec{n} \cdot \vec{r}_{0} \\
& 8 \quad \frac{1}{\left|r-\vec{r}_{0}\right|} \simeq \frac{1}{r}\left(1+\frac{\vec{n} \cdot \stackrel{\rightharpoonup}{r}_{r}^{r}}{r}\right) \\
& \text { - } T=t-\frac{\left|\stackrel{\rightharpoonup}{r}-\vec{r}_{0}\right|}{c} \simeq t-\frac{r}{c}+\frac{\vec{n}^{\prime}}{\bar{c}_{c}}
\end{aligned}
$$

For this problem $\rho\left(T, \vec{r}_{0}\right)=\rho\left(r_{0}\right) e^{-i \omega T}$
So

$$
\begin{gathered}
\varphi(t, r)=\int \frac{d^{3} r}{4 \pi r} \vec{\rho}\left(\vec{r}_{0}\right) e^{-i \omega\left(t-r / c+\vec{n}^{\circ} \cdot \vec{r}_{0}\right)}\left(1+\frac{\vec{n} \cdot \vec{r}_{0}}{r_{r}}\right) \\
\varphi(t, r) \simeq \frac{1}{r} \int \frac{d^{3} r_{0}}{4 \pi} \rho\left(r_{0}\right) \quad \frac{\text { This is zero. There is }}{\text { no net charge. }} \\
\frac{e^{-i \omega(t-r / /)}}{4 \pi r} \int d^{3} r_{0} \rho\left(r_{0}\right)\left(\vec{n} \cdot \vec{r}_{0}-i k \vec{n} \cdot r_{0}\right)
\end{gathered}
$$

- Since

$$
\int d^{3} r_{0} p\left(r_{0}\right) \stackrel{\rightharpoonup}{r}_{0}=\vec{p}_{0}
$$

We find

$$
\varphi(t, r)=\frac{e^{-i \omega(t-r / c)}}{4 \pi r}\left[\frac{\stackrel{\rightharpoonup}{n} \cdot \vec{p}_{0}-i k \stackrel{\rightharpoonup}{n} \cdot \vec{p}_{0}}{r}\right]
$$

Similarly

$$
\vec{A}(t, r)=\int d^{3} \vec{r}_{0}<\frac{\vec{J}\left(T, \vec{r}_{0}\right) / c}{\left.4 \pi r \vec{r}_{0}-\vec{r}_{j}\right)}
$$

- Let us work to lowest order. Previously we couldn't do this because the zeroth order term vanished on the previous page

$$
\vec{A}(t, r) \simeq \frac{1}{4 \pi r} \int d^{3} r_{0} \frac{\vec{J}\left(T, r_{0}\right)}{\bar{c}}
$$

Recall that

$$
\int d^{3} r_{0} \vec{J}\left(T, r_{0}\right)=\frac{\partial}{\partial T} \vec{p}\left(T, r_{0}\right)
$$

ie. in a dipole approx, with the dipole at the origin

$$
\vec{J}\left(t, r_{0}\right)=\partial_{t} \stackrel{\rightharpoonup}{p}(t) \delta^{3}\left(\vec{r}_{0}\right)
$$

So

$$
\stackrel{\rightharpoonup}{A}=\frac{1}{4 \pi r c} \cdot \frac{\partial}{\partial t} \stackrel{\rightharpoonup}{p}(T)
$$

Now

$$
\begin{gathered}
\vec{p}(T)=\vec{p}_{0} e^{-i \omega T} \\
\frac{\partial \vec{p}}{\partial T}=-i \omega \vec{p}_{0} e^{-i \omega T} \\
\vec{A} \simeq-\frac{-i \omega P_{0} e^{-i \omega T}}{4 \pi r}
\end{gathered}
$$

So to zeroth order

$$
T=t-\frac{r}{c}+\frac{n \cdot r_{0}}{c} \text { at zeroth order }
$$

and

$$
\bar{A}=\frac{-i \omega \vec{p}_{0} e^{-i \omega(t-r / c)}}{4 \pi r c}
$$

The answer follows using $k=w / c$
(c) In the far field $r k \gg 1$

$$
\text { e } \begin{aligned}
\varphi & =\frac{e^{-i \omega t+i k r}\left(-i k \vec{n} \cdot \vec{p}_{0}\right)}{4 \pi r} \\
\cdot \vec{A} & =\frac{e^{-i \omega t+i k r}}{4 \pi r}\left(-i k \vec{p}_{0}\right)
\end{aligned}
$$

- Then

$$
\vec{E}=-\frac{1}{c} \partial_{t} \vec{A}-\nabla \vec{\varphi}
$$

Where in the far field the derivatives act on the phase. If they acted on $\frac{1}{r}$
you would get something which is not far field:
$\nabla\left(\frac{1}{r}\right) \times \frac{1}{r^{2}} \quad$ falls of too fast as

- So

$$
\begin{aligned}
\vec{E}= & \frac{e^{-i \omega t+i k r}}{4 \pi r}\left(+i \frac{\omega}{c}\left(-i k \vec{p}_{0}\right)\right) \\
& +\frac{e^{-i \omega k+i k r}}{4 \pi r}\left((-i k \vec{r})\left(-i k \stackrel{\rightharpoonup}{n} \cdot \vec{p}_{0}\right)\right)
\end{aligned}
$$

Thus

$$
\downarrow^{\text {from } \frac{-1}{c} \partial_{t} \stackrel{\rightharpoonup}{A}} \text { from }-\nabla \varphi
$$

$$
=-k^{2} \frac{e^{-i \omega t}+i k r}{4 \pi r}\left(-\vec{p}_{0}+\vec{n}\left(n \cdot p_{0}\right)\right)
$$

$$
\vec{E}=-k^{2} \frac{e^{-i \omega t+i k r}}{4 \pi r}\left(\vec{n} \times \vec{n} \times \vec{p}_{0}\right)
$$

Watch me!!

- We notice that it is the $-\nabla \varphi$ term which cancels the piece of $-\frac{12}{c}+\vec{A}$ which is parallel to $\vec{n}$. The $\vec{E}$-field must be transverse to $\vec{n}$ in the far field. Current conservation relates $\varphi$ and $\vec{A}$ and guarantees that this will always happen.
d) Algebra:

$$
\begin{array}{ll}
\vec{E}=-\frac{1}{c} \partial_{+} \vec{A}-\vec{\nabla} \varphi \quad & \text { General } \\
\vec{\varphi}=\frac{e^{-i \omega t+i k r}\left(\vec{n} \cdot \vec{P}_{0}\right)}{4 \pi r}\left(\frac{1}{r}-i k\right) \\
\vec{A}=\frac{e^{-i \omega t+i k r}}{4 \pi r}\left(-i k \vec{p}_{0}\right) & \begin{array}{l}
\text { From } \\
\text { previous } \\
\text { item }
\end{array}
\end{array}
$$

$$
-\frac{1}{c} 2 t \vec{A}=\frac{e^{-i \omega t+i k r}}{4 \pi r}\left(+k^{2} \vec{p}_{0}\right)
$$

derivative of phase

$$
\begin{aligned}
-\nabla \varphi= & \frac{e^{-i \omega t+i k r}}{4 \pi}\left(n \cdot p_{0}\right)\left(\frac{1}{r^{2}}-\frac{-k}{r}\right)(-i k \vec{h}) \\
& +\frac{e^{-i \omega t+i k r}}{4 \pi}\left(\vec{n}-p_{0}\right)\left(\frac{2}{r^{3}} \vec{n}-i \frac{k}{r^{2}} \vec{n}\right)
\end{aligned}
$$

${ }^{\text {derivative of }} 1 / r=$

$$
+\frac{e^{-i \omega t+i k r}}{4 \pi}\left(\frac{1}{r^{2}} \frac{i k}{r}\right)\left(\frac{-\vec{p}+\vec{n} k n \cdot \vec{p})}{r}\right)
$$

$$
\begin{equation*}
\frac{\partial}{\partial r^{i}}\left(\frac{r_{j}}{r}\right)=\delta_{i j}-n_{i} n_{j} \tag{n}
\end{equation*}
$$

Collecting terms in powers of Yr

$$
\begin{aligned}
\vec{E}=\frac{e^{-i \omega t+i k r}}{4 \pi}\left[3 \vec { n } \left(\frac{\left(\vec{n} \cdot \vec{p}_{0}\right)}{r^{3}}-\vec{p}_{0}\right.\right. & -i k \frac{3 \vec{n}\left(\vec{n} \cdot \vec{p}_{0}\right)-\vec{p}_{0}}{r^{2}} \\
& +k^{2} \frac{\left(\vec{p}_{0}-\vec{n}\left(n \cdot p_{0}\right)\right)}{r}
\end{aligned}
$$

So in time

$$
\begin{aligned}
& \vec{E}=\frac{1}{4 \pi}\left[\left(1+\frac{r}{c} \frac{\partial}{\partial t}\right)\left(\frac{3 \vec{n}\left(\vec{n} \cdot \vec{p}\left(t_{e}\right)\right)-\vec{p}\left(t_{e}\right)}{r^{3}}\right)\right. \\
&\left.+\frac{\vec{n} \times \vec{n} \times \ddot{p}\left(t_{e}\right)}{4 \pi r c^{2}}\right]
\end{aligned}
$$

Note the $\left(1+\frac{r}{c} \frac{\partial}{\partial t}\right)$ structure is guaranteed by the fact that in quasi-statics the electric field corrects start like $\frac{1}{c^{2}}$

Expanding $t_{e}=t-r / c$ in powers of $r / c$

$$
\vec{E}=\frac{3 \vec{n}(\vec{n} \cdot \vec{p}(t))-\vec{p}(t)}{4 \pi r^{3}}+\underset{\sim}{\uparrow}+O\left(\frac{1}{c^{2}}\right)
$$

cancellation at order $1 / c$ !

The $1 / c^{2}$ term is:

$$
\begin{aligned}
& -\frac{1}{2}\left(\frac{r}{c}\right)^{2}\left(\frac{3 \vec{n}(\vec{n} \cdot \ddot{p}(t))-\ddot{p}}{4 \pi r^{3}}\right)+\frac{-\ddot{p}+\vec{n}(\vec{n} \cdot \ddot{p})}{4 \pi r c^{2}} \\
& =-\left(\frac{\ddot{p}+\vec{n}(\vec{n} \cdot \ddot{p})}{4 \pi r c^{2}}\right)
\end{aligned}
$$

in agreement with our quasistatic result.

## Problem 3. Radiation in the lowest Bohr Orbit

In the Bohr model, a classical non-relativistic electron circles a proton in a circular orbit with angular momentum $L=\hbar$, due to the Coulomb attraction between the electron and the proton.
(a) Recall that that the electron kinetic energy is half of minus its potential energy (for a coulomb orbit). Recall also that the lowest bohr orbit has velocity, $\beta=\alpha$ where $\beta=v_{e} / c$, and $\alpha=e^{2} /(4 \pi \hbar c)=1 / 137$. Prove these statements.
(b) Write down the (total=kinetic + potential) energy and radius of the lowest Bohr orbit in terms of the electron mass, $m_{e}, \hbar, c$ and $\alpha$. What is the size of the Bohr radius $a_{o}$ compared to the electron compton wavelength, i.e. $a_{o} /\left(\hbar /\left(m_{e} c\right)\right)$ ?
(c) One of the difficulties with the Bohr model, is that classically the electron would radiate. Determine the energy lost to radiation per unit time, for an electron in the lowest orbit.
(d) Determine the energy radiated per revolution in the Bohr model, $\Delta E$, and compare $\Delta E$ to the (kinetic+potential) energy of the orbit, i.e. compute $\Delta E / E_{\text {orbit }}$. Express $\Delta E / E_{\text {orbit }}$ in terms of the fine structure constant, and estimate its value.
(e) If the electron moves in the $x, y$ plane determine the time averaged power radiated per solid angle, $\overline{d P} / d \Omega$. Use a complex notation $\boldsymbol{r}(t)=a_{0}(\hat{\mathbf{x}}+i \hat{\mathbf{y}}) e^{-i \omega_{0} t}$.

You should find

$$
\begin{equation*}
\frac{\overline{d P}}{d \Omega}=\frac{e^{2}}{16 \pi^{2} c^{3}} \frac{1}{2}\left(1+\cos ^{2} \theta\right)\left(\omega_{o}^{2} a_{o}\right)^{2} \tag{22}
\end{equation*}
$$

where $\omega_{o}$ is the angular velocity of the electron
(f) Check your result of part (e) by integrating over solid angle and comparing with part (c).
(g) Now we will study the polarization of the light. (These questions do not require calculation).
(i) If the emitted light is observed along $x$ axis, what is the polarization of the radiated light? Explain physically.
(ii) If the emitted light is observed along the $y$ axis, what is the polarization of the radiated light? Explain physically.
(iii) If the emitted light is observed along the $z$ axis, what is the polarization of the light? Explain physically.
(h) The power radiated along the $z$-axis is twice as large as the power radiated along the $x$-axis. Explain this result physically.
a) Optional

For the Bohr ot bit,

$$
\begin{aligned}
& F=\frac{m v^{2}}{r} \\
& \frac{e^{2}}{4 \pi r^{2}}=\frac{m v^{2}}{r} \\
& \frac{e^{2}}{4 \pi r}=m v^{2} \text { so } \rightarrow(-P E)=2 \mathrm{KE}
\end{aligned}
$$

So

$$
\begin{aligned}
& \frac{+e^{2}}{4 \pi a_{0}}=2 \cdot\left(\frac{\hbar^{2}}{2 m a_{0}^{2}}\right) \propto \text { the KE is } \frac{t}{2 m a_{0}^{2}} \\
& a_{0}=\frac{\hbar^{2}}{m}\left(\frac{e^{2}}{4 \pi}\right)=\hbar m c\left(\frac{e^{2}}{4 \pi \hbar c}\right) \\
& a_{0}=\frac{\hbar}{m c \alpha} \\
& L_{0}=m v r \\
& \frac{\hbar}{m a_{0}}=V \Rightarrow V=\alpha c
\end{aligned}
$$

Then
b) So

$$
E=K E+P E=-\frac{1}{2} P E+P E=\frac{P E}{2}
$$

Then,

$$
\begin{aligned}
& E=\frac{-e^{2}}{2(4 \pi a)}=\frac{1}{2} m c^{2} \alpha^{2} \\
& a_{0}=\frac{\hbar}{m c \alpha}
\end{aligned}
$$

$$
\left(\frac{a_{0}}{t / m c}\right)=\frac{1}{\alpha}=137
$$

Bohr
a) $a_{0}=\frac{\hbar}{m c \alpha}$

$$
E=-\frac{1}{2} \frac{e^{2}}{4 \pi a_{0}}
$$

b) Using the Larmour formula

$$
\begin{aligned}
a & =\omega_{0}^{2} a_{0} \\
P & =\frac{e^{2}}{4 \pi} \frac{2}{3} \frac{a^{2}}{c^{3}} \\
P & =\frac{e^{2}}{4 \pi} \frac{2}{3} \frac{\omega_{0}^{4} a_{0}^{2}}{c^{3}} \quad \omega_{0}=\frac{\alpha^{c}}{a_{0}} \\
P & =\frac{e^{2}}{4 \pi} \frac{2}{3} \frac{\alpha^{4} c}{a_{0}^{2}} \\
P & =\frac{e^{2}}{4 \pi a} \frac{2}{3} \alpha^{4} \frac{c}{a_{0}}
\end{aligned}
$$

c) So for part.

$$
\Delta E=\frac{2 \pi a_{0}}{\alpha c} P
$$

So $\quad \Delta E=\frac{e^{2}}{4 \pi a_{0}} \frac{2}{3} \alpha^{4} \frac{c}{a_{0}} \frac{2 \pi a_{0}}{c \alpha}$

$$
\begin{aligned}
\Delta E & =\frac{e^{2}}{4 \pi a_{0}} \frac{4 \pi}{3} \alpha^{3} \\
\text { And } \quad \frac{\Delta E}{E} & =\frac{8 \pi}{3} \alpha^{3} \sim 10^{-6}
\end{aligned}
$$

d) Using the Larmour result

$$
E_{r a d}=\frac{e}{4 \pi r c^{2}} n \times n \times a\left(t_{e}\right)
$$

Using

$$
\begin{aligned}
& \vec{r}=\left(\cos \omega_{0} t_{e}, \sin \omega_{0} t_{e}, 0\right) a_{0} \\
& a=-\omega_{d}^{2} a_{0}\left(\cos \omega_{0} t, \sin \omega_{0} t, 0\right)
\end{aligned}
$$

$\hat{\omega}_{\text {will take }}^{\Rightarrow} \vec{a}=-\omega_{0}^{2} a_{0}\left(\hat{x} e^{-i \omega_{0} t}+i \hat{y} e^{-i \omega_{0} t}\right)$ The $\vec{a}=-\omega_{0}^{2} a_{\Delta} e^{-i \omega t}(\hat{x}+i \hat{y})$
part of this

Bohr pg. 3
So then

$$
\frac{\overline{d p}}{d \Omega}=\frac{e^{2}}{16 \pi r c^{3}}\left|\frac{\vec{n} \times \vec{n} \times \vec{a}}{2}\right|^{2}
$$

$$
n \times n \times \vec{a}=-\vec{a}+\vec{n}(n \cdot \vec{a})
$$

So taking $\vec{n}$ in the $x, z$ plane:

$$
\begin{gathered}
n \times n \times \vec{a} \propto-(\hat{x}+i \hat{y})+\vec{n}(\vec{n} \cdot(\hat{x}+i \hat{y})) \\
\propto-(\hat{x}+i \hat{y})+\vec{n}(n \cdot \hat{x}) \\
\ln \times n \times\left.\vec{a}\right|^{2} \propto\left(\hat{x}^{2}+\hat{y}^{2}\right)-(n \cdot x)^{2} \\
\propto 2-\sin ^{2} \theta
\end{gathered}
$$

Then

$$
\frac{d \bar{P}}{d \Omega}=\frac{e^{2}}{16 \pi c^{3}} \frac{1}{2}\left(2-\sin ^{2} \theta\right)\left(\omega_{0}^{2} a_{0}\right)^{2}
$$

## Problem 4. Radiation from a Phased Array

A current distribution consists of $N$ identical souces. The $k$-th source is identical to the first source except for a rigid translation by an amount $\boldsymbol{R}_{k}(k=1,2, \ldots, N)$. The sources oscillate at the same frequency but have different phases $\delta_{k}$. That is

$$
\begin{equation*}
\boldsymbol{j}_{k} \propto \exp \left(-i\left(\omega t+\delta_{k}\right)\right) \tag{23}
\end{equation*}
$$

(a) Show that the angular distribution of radiated power can be written as a product of two factors: one is the angular distribution for $N=1$; the other depends on $\boldsymbol{R}_{k}$ and $\delta_{k}$, but not on the structure of the sources.
(b) The planes of two square loops (each with sided length $a$ ) are centered on (and lie perpendicular to) the $z$-axis at $z= \pm a / 2$. The loop edges are parrallel to the $x$ and $y$ coordinate axes. Find the angular distribution of power in the $x-z$ plane if the current at all points in both loops is $I \cos (\omega t)$. Make a polar plot of the angular distribution of power for $\omega c / a=2 \pi$ and $\omega c / a \ll 1$. Identify the multipole character of the radiation in the limit $\omega a / c \ll 1$.
You should find

$$
\begin{equation*}
\frac{d P}{d \Omega}=\frac{I_{o}^{2} a^{2} \omega^{2}}{32 \pi^{2} c^{3}}(2 \sin (\sin \theta k a / 2))^{2}(2 \cos (\cos \theta k a / 2))^{2} \tag{24}
\end{equation*}
$$

(c) The limit $\omega a / c \ll 1$ has a simple physical interpretation. Describe this interpretation and show that it reproduces all aspects of the power distribution (including normalization factors) in the limit $\omega a / c \ll 1$.
(d) Repeat part (b) when the current in the upper loop is $I \cos \omega t$ and the current in the lower loop is $-I \cos \omega t$.

Radiation from a phased Array
a) $r \vec{A}_{\mathrm{rad}}=\frac{1}{4 \pi} \int_{r_{0}} \frac{\vec{J}}{c}\left(T, \vec{r}_{0}\right)$

The integration over the current is as follows:


So

$$
\begin{aligned}
\vec{r} \vec{A}_{r a d} & =\frac{e^{-i \omega(t-r / c)}}{4 \pi} \sum_{a} \int_{\Delta r} e^{-i \omega n \cdot \Delta \vec{r} / c} g\left(\overrightarrow{\Delta r}_{r}\right) e^{-i \frac{\omega}{c} n \cdot R_{a}} e^{-i \delta_{a}} \\
& =\left[\frac{e^{-i \omega(t-r / c)}}{4 \pi} \int_{\Delta \vec{r}}^{e^{-i \omega \vec{n} \cdot \Delta \vec{r} / c} \vec{\gamma}(\Delta r)}\right]\left[\sum_{a} e^{-i \omega \frac{n}{c} \cdot R_{a}} e^{-i \delta_{a}}\right]
\end{aligned}
$$

$$
=
$$

Phased Array Pg. 2
So taking the Poynting flux we have


We have since $e^{-i \omega(t-r / c)}$ is an irrelevant overall phase

$$
\begin{aligned}
& \frac{d P}{d \Omega}=\frac{1}{c^{3}}\left|\frac{-i \omega}{4 \pi} \int_{\Delta r_{0}} \vec{n} \times \stackrel{\rightharpoonup}{n} \times g(\Delta r) e^{-i \omega \vec{n} \cdot \Delta r / c}\right|^{2} \\
&=\left|\begin{array}{ll} 
& \sum_{a} e^{-i \delta_{a}} e^{-i \omega n} \cdot \vec{R}_{a} / c
\end{array}\right|^{2}
\end{aligned}
$$

Phased Array 3
b) To. work out the power from the array we must work out the integral

$$
r \vec{A}_{1} \equiv \frac{1}{4 \pi} \int d^{3} \Delta r e^{-i \omega n \cdot \Delta r / c} \hat{f}(\overrightarrow{\Delta r})
$$

from a single square loop. Taking $\vec{n}$ in the $x, z$ plane $\vec{n}=(\sin \theta, 0, \cos \theta)$ as required by the problem statement
$3 D$ view


Top view


Lets work out the phase for each of the four legs
(1) $e^{-i \omega \vec{n} \cdot \Delta \vec{r} / c}$ is with $\quad \Delta \vec{r}=\left(x,-\frac{a}{2}, 0\right)$

$$
e^{-i \omega n \cdot \overrightarrow{\Delta r} / c}=e^{-i \omega \sin \theta x}
$$

(3) Then the phase in leg 3 is also

$$
e^{-i \omega \vec{n} \cdot \Delta \vec{r} / c}=e^{-i \omega \sin \theta x}
$$

Phased Array Pg. 4
Since the phase is the same but the current is opposite

$$
\int_{(1)+(3)} d^{3} \Delta r e^{-i \omega \vec{n} \cdot \Delta \vec{r} / c} \vec{f}(\Delta r)=0
$$

For legs (2) and (4) we have with $\overrightarrow{\Delta r}_{(2)}=(a, y, 0)$
(2) $e^{-i \omega \vec{n} \cdot \overrightarrow{\Delta r} / c}=e^{-i \omega} e^{-i n} \sin \theta a / 2$

And for (4) $\Delta \vec{r}_{\Theta}=(-a / 2, i y, 0)$
(4) $e^{-i \omega n} \cdot \Delta \pi / c=e^{t i \omega}=\frac{\sin \theta o / 2}{}$

So

$$
\begin{aligned}
\int_{(2)} d^{3} \Delta r e^{-i \underline{c} \hat{c} \cdot \Delta \vec{r}} \vec{n} & =\int d y\left(e^{-i \omega \sin \theta a / 2}-e^{+i \omega \sin \theta a / 2}\right) I_{0} \hat{y} \\
& =I_{0} a \sum_{i} \sin (\omega \sin \theta a) \hat{2} \frac{y}{2}
\end{aligned}
$$

And thus

$$
\vec{r} \vec{A}_{1} \equiv \frac{i}{2 \pi} I_{0} a \sin \left(\omega \sin \theta_{\frac{a}{2}}^{2}\right) \hat{y}
$$

we need

$$
\begin{aligned}
\vec{n} \times \vec{n} \times\left(r \vec{A}_{1}\right) & =-r \vec{A}_{1}+\vec{n}\left(n \cdot r \vec{A}_{1}\right) \\
& =-r \vec{A}_{1}
\end{aligned}
$$

Phased Array Pg. 5
To compute the power we have

$$
\frac{d P}{d \Omega}=\frac{1}{c^{3}} \omega^{2}\left(\frac{I_{0} a \sin (\omega \sin \theta a)}{2 \pi}\right)^{2}\left|\frac{\Sigma}{a} e^{-i \delta_{a}} e^{-i \frac{\omega}{e} \vec{n} \cdot \overrightarrow{R_{a}}}\right|^{2}
$$

structure factor
So we have to work but the structure factor:

$$
\begin{array}{rlr}
\text { S.F. } & =\left|e^{-i \underline{\omega} \cos \theta \frac{a}{2}}+e^{-i \frac{\omega}{c} \cos \theta \frac{a}{2}}\right|^{2}, & \vec{R}_{1}=\left(0,0, \frac{a}{2}\right) \leftarrow \begin{array}{c}
\text { position } \\
\text { of }
\end{array} \\
& =\left|2 \cos \left(\cos \theta \frac{\mathrm{ka}}{2}\right)\right|^{2} & \\
& \vec{R}_{2}=(0,0,-a / 2) \in \text { center }
\end{array}
$$

So the full result is

$$
\frac{d P}{d \Omega}=\frac{I_{0}^{2} a^{2} \omega^{2}(2 \sin (\sin \theta k a / 2))^{2}(2 \cos (\cos \theta k a / 2))^{2}}{16 \pi^{2} c^{3}}
$$

For $k a \ll 1$ the second factor is 4, but the first factor is expanded

$$
\begin{aligned}
& \frac{d P}{d \Omega}=\frac{I_{0}^{2} a^{2}}{16 \pi^{2} c^{3}} \omega^{2} a^{2} 4 \sin ^{2} \theta\left(\frac{w a}{c}\right)^{2} \cdot 4 \\
& \frac{d P}{d \Omega}=\left(2 I_{0} / c a^{2}\right)^{2} \frac{1}{16 \pi^{2} c^{3}} \sin ^{2} \theta \\
& \omega^{4}
\end{aligned}
$$

$\zeta$ this is the formula for magnetic dipole radiation with $\vec{m}=2 I{ }_{c}^{I} a^{2}$ with the 2 , because there

Phased Array Pg. 6
c) If the sign of the lower current is flipped then the structure factor changes

$$
\begin{aligned}
\text { S.F. } & =\left|e^{-i \frac{\omega}{c} \cos \theta \theta / 2}-e^{-i \frac{\omega}{c} \cos \theta a / 2}\right|^{2} \\
& =\left(2 \sin \left(\cos \theta \frac{\omega a}{2 c}\right)\right)^{2}
\end{aligned}
$$

And

$$
\frac{d P}{d \Omega}=\frac{I_{0}^{2} a^{2} \omega^{2}(2 \sin (\sin \theta k a / 2))^{2}(2 \cdot \sin (\cos \theta \mathrm{ka} / 2))^{2}}{16 \pi^{2} c^{3}}
$$

For small frequency we find magnetic: quadrupole radiation:

$$
\frac{d P}{d \Omega}=\frac{\epsilon}{16 \pi^{2}}\left(\frac{T_{0}}{c}\right)^{2} \cdot\left(\frac{\omega a}{c}\right)^{6} \sin ^{2} \theta \cos ^{2} \theta
$$

Or

e) To check this result we integrate over $\Omega$

$$
\bar{P}=\frac{e^{2}}{16 \pi c^{3}}\left(\frac{\left.\omega_{0}^{2} a_{0}\right)^{2}}{2} \int d \Omega\left(2-\sin ^{2} \theta\right)\right.
$$

Evaluating I:

$$
\begin{aligned}
I & =2 \pi \int_{-1}^{1} d x\left(2-\left(1-x^{2}\right)\right) \quad x \equiv \cos \theta \\
& =2 \pi \int_{-1}^{1} d x\left(1+x^{2}\right) \\
I & =2 \pi\left(2+\frac{2}{3}\right)=\frac{16 \pi}{3}
\end{aligned}
$$

So $\bar{p}=\frac{e^{2}}{16 \pi c^{3}}\left(\frac{\omega_{0}^{2} a_{0}}{2}\right)^{2} \cdot \frac{16 \pi}{3}$
$\bar{p}=\frac{e^{2}}{6 \pi c^{3}}\left(w_{0}^{2} a_{0}\right)^{2} \quad$ which agrees with

Then the polarization is recorded by: $\vec{a}_{\perp}$

$$
\vec{E}_{r a d}=\frac{e}{4 \pi r c^{2}} \vec{n} \times \vec{n} \times a\left(t_{e}\right)=\frac{e}{4 \pi r c^{2}}\left[-\vec{a}_{1}(t)\right]
$$

Where

$$
\vec{a}=-\omega^{2} a_{0} e^{-i \omega t_{e}}(\hat{x}+i \hat{y})
$$



- On the $x$ axis, $\vec{a}_{T}=-\omega^{2} a_{0} e^{-i \omega t_{e}}(i \hat{y})$, so $\vec{F}$ points on the $y$-axis
- On the $y$-axis, $\vec{a}_{T}=-\omega_{0}^{2} a_{0} e^{-i \omega t}(\hat{x})$ so the polarization points on the $\hat{x}$ axis
- On the $z$-axis $\vec{a}_{\Gamma} \propto e^{-i \omega t_{e}}(\hat{x}+i \hat{y})$
which is circular

$$
\operatorname{Re} \vec{a}_{T}=(\cos \omega t) \hat{x}+(\sin \omega t) \hat{y}=\overbrace{\text { right circular }}
$$

f) One can understand circular motion as a super-position of an $x$-oriented dipole and $y$-oriented dipole, $90^{\circ}$ out of phase. On the $x$-axis only the radiation from the $y$-oriented dipole contributes as the $x$-oriented dipole is parallel to the observation direction only transverse currents contribute to the radiation). On the $z$-axis both the $x$-oriented dipole + the $y$-oriented dipole contribui to the radiation field.

The two dipoles add incoherently

## Problem 5. A Charged Rotor: Zangwill

Two identical point charges of charge $q$ are fixed to the ends of a rod of length $2 \ell$ which rotates with constant angular velocity of $\frac{1}{2} \omega$ in the $x-y$ plane about an axis perpendicular to the rod and through its center
(a) Calculate the electrid dipole moment $\boldsymbol{p}(t)$. Is there electric dipole radiation?
(b) Calculate the magnetic dipole moment $\boldsymbol{m}(t)$. Is there magnetic dipole radiation?
(c) Show that the electric quadruple moment is

$$
Q(t)=3 q \ell^{2}\left(\begin{array}{ccc}
\frac{1}{3}+\cos \omega t & \sin \omega t & 0  \tag{25}\\
\sin \omega t & \frac{1}{3}-\cos \omega t & 0 \\
0 & 0 & -\frac{2}{3}
\end{array}\right)
$$

(d) Show that the time averaged angular distribution of radiated power is

$$
\begin{equation*}
\frac{\overline{d P}}{d \Omega}=\frac{1}{128 \pi^{2} c^{5}} \omega^{6} q^{2} \ell^{4}\left(1-\cos ^{4} \theta\right) \tag{26}
\end{equation*}
$$

(1) Charged Rotor

Then consider

a) There is no electric dipole moment
b) There is a magnetic dipole moment

$$
\begin{aligned}
\vec{m} & =\frac{1}{2 c} \int d^{3} x \vec{x} \times \vec{J} \\
& =\frac{1}{2 c}\left[\omega q l^{2} \hat{z}+\omega q l^{2} \hat{z}\right]=\omega q \frac{l^{2}}{c} \hat{z}
\end{aligned}
$$

But the magnetic moment is constant and does not contribute to the radiation
(2) Charged Rotor
c) To find the quadrupole moment we note the definition

$$
Q^{i y}=\int d^{3} \vec{r} \rho(\vec{r})\left(3 r^{i} r^{y}-r^{2} \delta^{\prime} y\right)
$$

Then lets find the $Q$ at this time $(t=0)$ and then rotate


Qi is clearly zero for the off diagonal components. But the diagonal components are

$$
\begin{aligned}
& Q^{x x}=\left[\left(3 l^{2}-l^{2}\right)+3(-l)^{2}-\left(-l^{2}\right)\right] q \\
& Q^{x x}=4 q l^{2} \\
& Q^{y y}=-2 q l^{2} \\
& Q^{z z}=-2 q l^{2}
\end{aligned}
$$

Now we know the components in the rotated coordinate system, We want the components in the unrotated system.
(3) Charged Rotor

The appropriate rotation matrix is found from the picture:


$$
R_{j}^{i}=(R)_{j}^{i}=\left(\begin{array}{cc|c}
\cos \theta & -\sin \theta & 0 \\
\sin \theta & \cos \theta & 0 \\
\hline 0 & 0 & 1
\end{array}\right)
$$

So

$$
Q^{i y}=(R)_{l}^{i}(R)_{m}^{j} Q^{l m}
$$

Or in matrices

$$
\left.\begin{array}{l}
\underline{Q}=R Q R^{\top} \\
\underline{Q}=q l^{2}\left(\begin{array}{ll}
\cos \theta & -\sin \theta \\
\sin \theta & \cos \theta \\
\hline
\end{array}\right)(1
\end{array}\right)\left(\begin{array}{lll}
4 & \\
& -2 & -2
\end{array}\right) \times \quad\left(\left.\begin{array}{ll}
\cos \theta & \sin \theta \\
-\sin \theta & \cos \theta
\end{array} \right\rvert\,\right.
$$

(4) Charged Rotor

Multiplying out, \& using

$$
1+3 \cos 2 \theta=4 \cos ^{2} \theta-2 \sin ^{2} \theta
$$

gives with $2 \theta=\omega t$

$$
\underline{Q}=q l^{2}\left(\begin{array}{cc|c}
1+3 \cos \omega t & 3 \sin \omega t & 0 \\
3 \sin \omega t & 1-3 \cos \omega t & 0 \\
0 & 0 & -2
\end{array}\right)
$$

Or,

$$
\underline{Q}=3 q l^{2}\left(\begin{array}{cc|c}
\frac{1}{3}+\cos \omega t & \sin \omega t & 0 \\
\sin \omega t & \frac{1}{3}-\cos \omega t & 0 \\
0 & 0 & -2 / 3
\end{array}\right)
$$

In terms of matrices:

$$
\underline{Q}=3 q l^{2}\left[e^{-i \omega t} \sigma_{z}+i e^{-i \omega t} \sigma_{x}\right]+\text { cons }
$$

where const $=\left(\begin{array}{rr}1 / 3 & \\ & 1 / 3 \\ & -2 / 3\end{array}\right) \mathrm{gl}^{2}$, and
$\sigma_{x}$ and $\sigma_{z}$ are pauli matrices. It is understood that we take the real pant of Eq*

$$
\sigma_{x}=\left(\begin{array}{ll|l}
0 & 1 & 0 \\
1 & 0 & 0 \\
\hline 0 & 0 & 0
\end{array}\right) \quad v_{z} \equiv\left(\begin{array}{cc|c}
1 & 0 & 0 \\
0 & -1 & 0 \\
\hline 0 & 0 & 1
\end{array}\right)
$$

(5) Charged Rotor
c) To find the power, then we use

$$
\vec{E}(t, \vec{r})=-\frac{1}{24 \pi r c^{3}}\left[Q^{\infty} \cdot \vec{n}-\vec{n}\left(\vec{n}^{T} \hat{Q} \stackrel{\rightharpoonup}{n}\right)\right] .
$$

Looking at the form of $\vec{Q}$ :

$$
Q=3 q l^{2}\left[e^{-i \omega t} \sigma_{z}+i \sigma_{x}\right]+\text { cons }
$$

we see that

$$
\vec{E}(\vec{t}, \vec{r})=-i \vec{E}(t, r)+\vec{E}_{2}(t, r)
$$

where $\vec{E}_{1}$ and $\vec{E}_{2}$ are real:

$$
\begin{aligned}
& \\
& \hline E_{q} \\
& \left(\vec{E}_{1}(t, r) \equiv+\frac{3 g l^{2} \omega^{3}}{24 \pi r c^{3}}\left[\left(\sigma_{z} \cdot \vec{n}\right)-\vec{n}\left(\vec{n}^{\top} \sigma_{z} n\right)\right]\right. \\
& \vec{E}_{2}(t, r) \equiv+\frac{3 q l^{2} \omega^{3}}{24 \pi r c^{3}}\left[\left(\vec{\sigma}_{x} \cdot \vec{n}\right)-\vec{n}\left(\vec{n}^{\top} \delta_{x} \vec{n}\right)\right]
\end{aligned}
$$

So

$$
\begin{aligned}
& \frac{d P}{d \Omega}=\frac{c}{2} \operatorname{Re}\left[\stackrel{\rightharpoonup}{E} \cdot \vec{E}^{*}\right] r^{2} \\
& \frac{d P}{d \Omega}=\frac{c}{2}\left(\left|E_{1}\right|^{2}+\left|E_{2}\right|^{2}\right) r^{2}
\end{aligned}
$$

(6) Charged Rotor

So we evaluate the square of the bracketed terms []$^{2}$ in Eq (*):

$$
\begin{aligned}
& {\left[\sigma \cdot \vec{n}-\vec{n}\left(n^{\top} \sigma n\right)\right]^{\top}\left[(\sigma \cdot n)-\vec{n}\left(n^{\top} \sigma n\right)\right] } \\
= & \left(n^{\top} \sigma-\left(n^{\top} \sigma n\right) n^{\top}\right)\left(\sigma \cdot n-n\left(n^{\top} \sigma n\right)\right) \\
= & n^{\top} \sigma^{2} n-2\left(n^{\top} \sigma n\right)^{2}+\left(n^{\top} \sigma n\right)^{2} n^{\top} n \\
= & n^{\top} \sigma^{2} n-\left(n^{\top} \sigma n\right)^{2}
\end{aligned}
$$

We used that $\sigma^{2}=\left(\frac{1}{1} f_{0}\right)$ for any pauli matrix and $n^{T} n=1$ for a unit vector.

Then we find
(*) $\frac{\overline{d P}}{d \Omega}=\frac{c}{2}\left(\frac{3 g l^{2} \omega^{3}}{24 \pi c^{3}}\right)^{2}\left[2\left(n^{\top} \sigma^{2} n\right)-\left(n^{\top} \sigma_{x} n\right)^{2}-\left(n^{\top} \sigma_{z} n\right)^{2}\right]$
Taking $\vec{n}$ in the $x-z$ plane:

$$
\begin{aligned}
\vec{n} & =(\sin \theta, 0, \cos \theta) \\
n^{\top} \sigma_{x} n & =\left(\begin{array}{ll}
\sin \theta & 0
\end{array}\right)\left(\begin{array}{ll}
0 & 1 \\
1 & 0
\end{array}\right)\binom{\sin \theta}{0}=0 \\
n^{\top} \sigma_{z} n & =\left(\begin{array}{ll}
\sin \theta & 0
\end{array}\right)\left(\begin{array}{cc}
1 & 0 \\
0 & -1
\end{array}\right)\binom{\sin \theta}{0}=\sin ^{2} \theta
\end{aligned}
$$

(7) Charged Rotor
and

$$
n^{\top} \sigma^{2} n=\left(\begin{array}{lll}
\sin \theta & 0 & \cos \theta
\end{array}\right)\left(\begin{array}{ll}
1 & \\
& 1
\end{array}\right)\left(\begin{array}{c}
\sin \theta \\
0 \\
\cos \theta
\end{array}\right)
$$

$$
=\sin ^{2} \theta
$$

So the term in square brackets in $E q \not \$$ on previous page is

$$
\begin{aligned}
{[] } & =\left[2 \sin ^{2} \theta-\left(\sin ^{2} \theta\right)^{2}\right] \\
& =[1-\cos 4 \theta]
\end{aligned}
$$

And restoring constants we find from Eq \$

$$
\frac{d \bar{P}}{d \Omega}=\frac{q^{2} l^{4} \omega^{6}}{128 \pi^{2} c^{5}}\left(1-\cos ^{4} \theta\right)
$$

## Problem 6. Basics of Relativity

(a) (Optional) The space time event at $X^{\mu}=\left(X^{0}, X^{i}\right)=(c t, \boldsymbol{x})$ happens at $\underline{X}^{\mu}=$ $\left(\underline{X}^{0}, \underline{X}^{i}\right)=(c \underline{t}, \underline{\boldsymbol{x}})$ according to an observer moving to the right along the $x$ axis with velocity $v$. Define the "light-cone" coordinates $x^{+} \equiv X^{0}+X^{1}$ and $x^{-} \equiv X^{0}-X^{1}$. Show that under this boost that the $x^{+}$coordinates are contracted, while the $x^{-}$coordinates are elongated

$$
\begin{align*}
& \underline{x}^{+}=e^{-y} x^{+}=\sqrt{\frac{1-\beta}{1+\beta}} x^{+}  \tag{27}\\
& \underline{x}^{-}=e^{y} x^{-}=\sqrt{\frac{1+\beta}{1-\beta}} x^{-} \tag{28}
\end{align*}
$$

Here

$$
\begin{equation*}
y=\tanh ^{-1} \beta=\frac{1}{2} \log \left(\frac{1+\beta}{1-\beta}\right) \tag{29}
\end{equation*}
$$

is the so-called "rapidity" of the boost. What is $\underline{x}^{+} \underline{x}^{-}$and why is it unchanged under boost?
(b) (Optional) A Lorentz tensor transforms as

$$
\begin{equation*}
\underline{T}^{\mu \nu}=L_{\rho}^{\mu} L_{\sigma}^{\nu} T^{\rho \sigma} \tag{30}
\end{equation*}
$$

Show that the transformation rule can be alternatively written

$$
\begin{equation*}
\underline{T}^{\mu}{ }_{\nu}=(\mathcal{L})^{\mu}{ }_{\rho} T^{\rho}{ }_{\sigma}\left(\mathcal{L}^{-1}\right)^{\sigma}{ }_{\nu} \tag{31}
\end{equation*}
$$

or equivalently

$$
\begin{equation*}
\underline{T}_{\nu}^{\mu}=L^{\mu}{ }_{\rho} L_{\nu}{ }^{\sigma} T^{\rho}{ }_{\sigma} \tag{32}
\end{equation*}
$$

(c) (Optional) The frequency and wave number of a plane wave of light, $e^{-i \omega t+i \boldsymbol{k} \cdot \boldsymbol{x}}=e^{i K \cdot X}$, form a lightlike four vector

$$
\begin{equation*}
K^{\mu}=\left(\frac{\omega}{c}, \boldsymbol{k}\right) \tag{33}
\end{equation*}
$$

(i) Show that $K \cdot K=K_{\mu} K^{\mu}=0$ (this is the statement that $K$ is lightlike.)
(ii) If a photon has frequency $\omega_{o}$ and is propagating along the $z$-axis, show (using the 4 -vector properties of $K^{\mu}$ ) that according to an observer propagating in the negative $z$ direction with speeed $\beta$

$$
\begin{equation*}
\omega=\sqrt{\frac{1+\beta}{1-\beta}} \omega_{o} \tag{34}
\end{equation*}
$$

(d) (Optional) Show that the four velocity $U^{\mu}=d x^{\mu} / d \tau$ satisfies $U_{\mu} U^{\mu}=-c^{2}$.
(e) (Optional) For a particle with four momentum $P^{\mu}=\left(\frac{E}{c}, \boldsymbol{p}\right)=m U^{\mu}$ show that $P_{\mu} P^{\mu}=$ $-\left(m c^{2}\right)^{2} / c^{2}$. This determines $E(\boldsymbol{p})$ the relation between energy and momentum:

$$
\begin{equation*}
\frac{E(\boldsymbol{p})}{c}=\sqrt{\boldsymbol{p}^{2}+(m c)^{2}} \tag{35}
\end{equation*}
$$

(i) Show the velocity of the particle (i.e. the group velocity) is

$$
\begin{equation*}
\boldsymbol{v}_{\boldsymbol{p}} \equiv \frac{\partial E(\boldsymbol{p})}{\partial \boldsymbol{p}}=\frac{c^{2} \boldsymbol{p}}{E} \tag{36}
\end{equation*}
$$

(f) (Do me! Not optional) A particle with velocity $v_{p}$ in the x direction. Using the 4vector transformation properties of $U^{\mu}$, show that according to an observer moving to the right with velocity $v$, the particle moves with velocity

$$
\begin{equation*}
\underline{v}_{p}=\frac{v_{p}-v}{1-v_{p} v / c^{2}} \tag{37}
\end{equation*}
$$

Basics The lorentz transformation is
a) $\quad\binom{\underline{x}^{0}}{\underline{x}^{1}}=\left(\begin{array}{cc}\gamma & -\gamma \beta \\ -\gamma \beta & \gamma\end{array}\right)\binom{x^{0}}{x^{1}}$

So

$$
\begin{aligned}
& x^{+}=x^{0}+x^{1} \\
& x^{-}=x^{0}-x^{1}
\end{aligned}
$$

And

$$
\begin{aligned}
\underline{x^{+}} & =x^{0}+\underline{x^{1}} \\
& =\gamma x^{0}-\gamma \beta x^{1}+-\gamma \beta x^{0}+\gamma x^{\prime} \\
& =\gamma\left(x^{0}+x^{1}\right)-\gamma \beta\left(x^{0}+x^{1}\right) \\
& =\gamma(1-\beta) x^{+} \\
\underline{x}^{+} & =\sqrt{\frac{1-\beta}{1+\beta}} x^{+}
\end{aligned}
$$

Note

$$
\operatorname{tanhy}=\beta \text { or } y=\frac{1}{2} \log \left(\frac{1+\beta}{1-\beta}\right)
$$

Basics pg. 2
So

$$
\underline{x}^{+}=e^{-y} x^{+}
$$

Similarly

$$
\begin{aligned}
\underline{x}^{-} & =\underline{x}^{0}-\underline{x}^{1} \\
& =\gamma(1+\beta) x^{-} \\
\underline{x^{-}} & =\sqrt{\frac{1+\beta}{1-\beta} x^{-}} \\
\underline{x}^{-} & =e^{+y} x^{-}
\end{aligned}
$$

b) $\quad T^{\mu \nu}=L^{\mu} L^{\nu}{ }_{\delta} T^{\rho \sigma}$

Then raising/lowering $\sigma$ and lowering $v$

$$
\begin{aligned}
I^{\mu} \nu & =L^{\mu} \rho L_{\nu}{ }^{\sigma} T^{\rho} \sigma \\
& =(\mathcal{L})^{\mu} \rho\left(\mathcal{L}^{-1 T}\right)_{\nu}^{\sigma} T^{\rho} \sigma \\
I^{I^{\mu}} & =(\alpha)^{\mu} \rho T^{\rho} \sigma\left(\mathcal{L}^{-1}\right)^{\sigma}{ }^{\rho}
\end{aligned}
$$

Basics pg. 3
c) A plane wave of light is.

$$
\begin{equation*}
e^{-i \frac{w}{c} x^{0}+i k x^{\prime}}=e^{i \phi} \tag{A}
\end{equation*}
$$

Under boost the speed of light is constaint. The only way to geaarranty this is if the phase is invariant under Lorentz transformations

$$
K^{\mu}=\left(\frac{w}{c}, k\right)
$$

Then

$$
k_{\mu} x^{\mu}=-\frac{\omega}{c} x^{0}+k x^{1}
$$

So
i) $\quad k_{\mu} k^{\mu}=-\left(\frac{w}{c}\right)^{2}+k^{2}=0$
since $w=c k$ is required in $E_{q} \ngtr$
ii) Under boost in negative $z$-direction

$$
\begin{aligned}
& \binom{\underline{\omega} / c}{\underline{k}}=\left(\begin{array}{cc}
\gamma & +\gamma \beta \\
+\gamma \beta & \gamma
\end{array}\right)\binom{\omega / c}{k} \\
& \underline{\omega} / c=\gamma \omega / c+\gamma \beta k
\end{aligned}
$$

Basics pg. 4
But $k=w / c$ so

$$
\begin{aligned}
\underline{\omega} / c & =\gamma(1+\beta) \omega / c \\
\underline{\omega} / c & =\sqrt{\frac{1+\beta}{1-\beta}} \omega / c
\end{aligned}
$$

d)

$$
\begin{aligned}
u^{\mu}=\frac{d x^{\mu}}{d \tau} & \text { so } \\
u_{\mu} u^{\mu}=\frac{d x}{d \tau} \frac{d x^{\mu}}{d \tau} & =\frac{\left(d x_{\mu} d x^{\mu}\right.}{(d \tau)^{2}} \\
& =-\frac{c^{2}(d \tau)^{2}}{(d \tau)^{2}}
\end{aligned}
$$

We used the definition of proper time $c^{2} d \tau^{2}=-d x_{\mu} d x^{\mu}=d s^{2}$

You can also use

$$
u^{\mu}=\left(\gamma_{c}, \gamma v\right)=c\left(\gamma, \gamma_{\beta}\right)
$$

So

$$
u_{\gamma} u^{\alpha}=c^{2}\left[-\gamma^{2}+\gamma^{2} \beta^{2}\right]=-c^{2}
$$

## Problem 7. One liners

(a) Starting from the Maxwell equations for $F^{\mu \nu}$ and the definition of $F^{\mu \nu}$, derive the wave equation $-\square A^{\mu}=J^{\mu} / c$.
(b) Starting from the maxwell equations for $F^{\mu \nu}$ in covariant form, show that we must have $\partial_{\mu} J^{\mu}=0$ for consistency.
(c) (This is two lines) Show that the energy conserivation and force laws

$$
\begin{align*}
\frac{d E_{\boldsymbol{p}}}{d t} & =q \boldsymbol{E} \cdot \boldsymbol{v}_{p}  \tag{38}\\
\frac{d \boldsymbol{p}}{d t} & =q\left(\boldsymbol{E}+\frac{\boldsymbol{v}_{\boldsymbol{p}}}{c} \times \boldsymbol{B}\right) \tag{39}
\end{align*}
$$

can be written covariantly

$$
\begin{equation*}
\frac{d P^{\mu}}{d \tau}=F^{\mu \nu} u_{\nu} / c \tag{40}
\end{equation*}
$$

Note that $E_{\boldsymbol{p}}$ (the energy of the particle) is different from $\boldsymbol{E}$ the electric field.
(d) From Eq. (40) show that $P_{\mu} P^{\mu}$ is constant in time.
(e) Show that $F_{\mu \nu}=\partial_{\mu} A_{\nu}-\partial_{\nu} A_{\mu}$ is invariant under the gauge transform

$$
\begin{equation*}
A_{\mu} \rightarrow A_{\mu}+\partial_{\mu} \Lambda(X) \tag{41}
\end{equation*}
$$

where $\Lambda$ is an arbitrary function of $X=(t, \boldsymbol{r})$.
(f) Given $F^{\mu \nu}$ the only two Lorentz invariant quantities are $F_{\mu \nu} F^{\mu \nu}$ and $F_{\mu \nu} \tilde{F}^{\mu \nu}$. Evaluate these two invariants in terms of $\boldsymbol{E}$ and $\boldsymbol{B}^{1}$

[^0]One liners
a)

$$
\begin{aligned}
-\partial_{\mu} F^{\mu v} & =J^{v} / c \\
-\partial_{\mu} & \left(\partial^{\mu} A^{v}-\partial^{v} A^{\mu}\right)=J^{v} / c \\
& -\left(\partial_{\mu} \partial^{\mu}\right) A^{v}+\partial^{v}\left(\partial A^{\mu}\right)=J^{v} / c
\end{aligned}
$$

with $\partial_{\mu} A^{\mu}=0$ and $\square=\partial_{\mu} \partial^{\mu}$ find

$$
-\square A^{v}=J^{v} / c
$$

b)

$$
\begin{aligned}
\partial_{\mu} F^{\mu \nu} & =J^{\nu} / c \\
\partial_{\nu}^{\partial_{\mu} F^{\mu \nu}} & =\partial_{\nu} J^{\nu} / c
\end{aligned}
$$

Fur is anti-symmetric, while $\partial_{\mu} \partial_{v}$ is symmetric, so this is zero

$$
\partial_{v} J^{v} / c=0
$$

c)

$$
\begin{aligned}
& \frac{d E_{p}}{d_{t}}=q \vec{E} \cdot \vec{v}_{p} \\
& \frac{d \vec{p}}{d t}=q\left(\stackrel{\rightharpoonup}{E}+\frac{\vec{u}_{p}}{c} \times \vec{B}\right)
\end{aligned}
$$

One liners pg. 2
Then multiply both sides by $\gamma$. Use

$$
\gamma \frac{d}{d t}=\frac{d}{d \tau} \quad \frac{d t}{\gamma_{p}}=d \tau
$$

So that with $U^{\mu}=(\gamma c, \gamma v)$ and $E^{i}=F^{0}$,

$$
\begin{aligned}
\frac{d\left(E_{p} / c\right)}{d \tau} & =q F^{0 i} \frac{u_{i}}{c} \quad \quad F_{i j}=\varepsilon_{i j k} B^{k} \\
\frac{d p^{i}}{d \tau} & =q\left(E^{i} \gamma+\varepsilon_{j k}^{i} \frac{\gamma}{c}_{c_{p}^{j}}^{c} B^{k}\right) \\
& =q\left(F_{0}^{i} \frac{u^{j}}{c}+F_{j}^{i} \frac{v^{j}}{c}\right)
\end{aligned}
$$

So

$$
\frac{d P^{\mu}}{d \tau}=q F^{\mu} v \frac{u^{\nu}}{c}
$$

d) $\quad P^{\mu} \frac{d P}{a \tau} \mu=q F_{\mu v} \frac{P^{\mu} u^{\nu}}{c}=q_{c}^{m} F_{\mu v} u^{\mu} u^{\prime}$

$$
=0
$$

since $F_{\mu \nu}$ is antisymm while $u^{\mu u} u^{v}$ is symmetric.

One liners pg. 3
e) $\quad F_{\mu \nu}=\partial_{\mu} A_{\nu}-\partial_{\nu} A_{\mu}$

So $\quad A_{\mu} \rightarrow A_{-\mu}=A_{\mu}-\partial_{\mu} \Lambda$

$$
\begin{aligned}
& =0 \\
& F_{\mu \nu}=\partial_{\mu} A_{\nu}-\partial_{\nu} A_{\mu}-\partial_{\mu} \partial_{\nu} \Lambda+\partial_{\nu} \partial_{\mu} \Lambda \\
& \text { f) } \quad F_{\mu \nu} F_{v}^{\mu \nu}=-2 E^{2}+2 B^{2}=-2\left(E^{2}-B^{2}\right) \\
& F^{\mu \nu}=\left(\begin{array}{c|ccc}
0 & \vec{E} \\
\hline-\vec{E} & 0 & B_{z} & -B \\
& -\vec{B}^{\prime} & 0 & B^{x} \\
- & & 0
\end{array}\right) \\
& F_{\mu v}=\left(\begin{array}{c|cc}
0 & -\vec{E}- \\
\hline E & 0 & B^{z} \\
-B^{y} \\
1 & 0 & B^{x}
\end{array}\right) \\
& \widetilde{F}_{\mu j}=\left(\begin{array}{c|cc}
0 & \vec{B}- \\
\hline-\vec{B} & 0 & E^{z}-E^{y} \\
& -E^{z} & 0 \\
E^{x} \\
E^{y} & -E^{x} & 0
\end{array}\right) \\
& F_{\mu \nu} F^{\mu \nu}=-4 \vec{E} \cdot \vec{B}
\end{aligned}
$$


[^0]:    $1_{\text {answers: }} 2\left(B^{2}-E^{2}\right)$ and $-4 E \cdot B$

