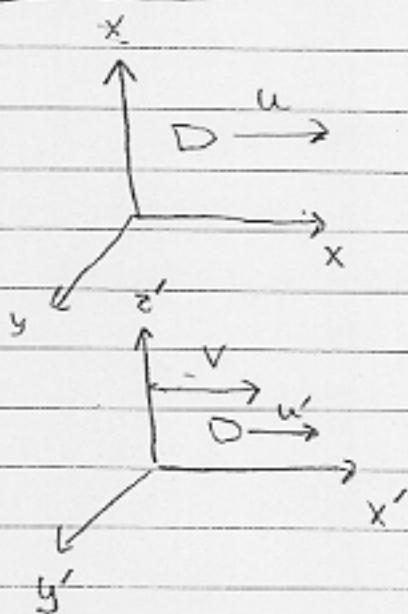


Last time



Earth Observer makes measurements of the rocket

$$t, x, y, z, u$$

For instance the position of the rocket vs. time

$$x = ut$$

u = The velocity of the rocket

The observer on Jupiter which is moving with respect to Earth in the x direction measures different coordinates and velocities

$$t' = t$$

$$x' = x - vt$$

$$y' = y$$

$$z' = z$$

$$u' = u - v$$

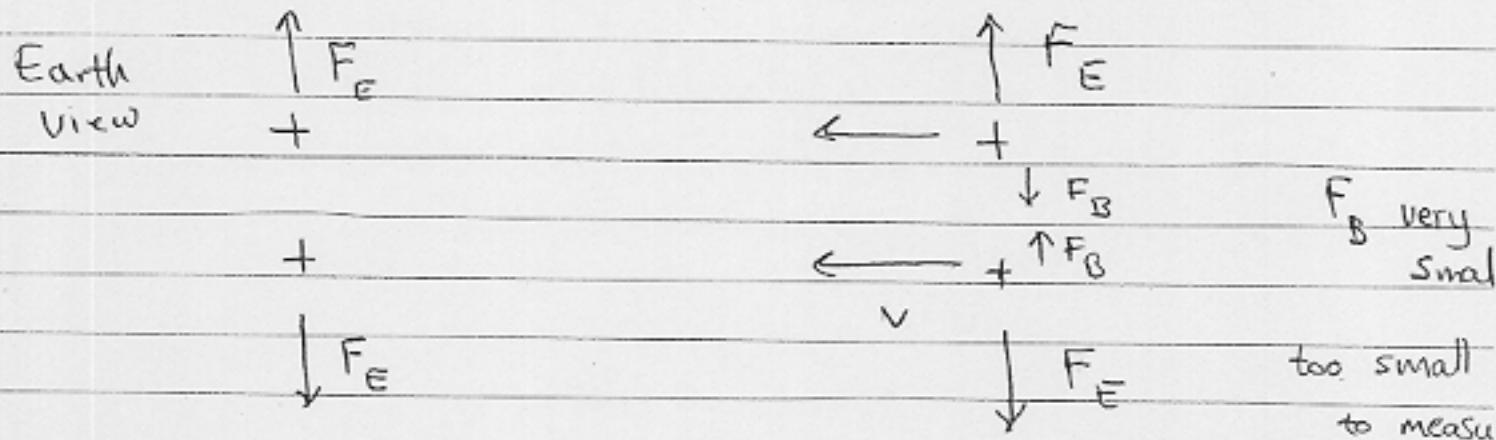
The Rules of Physics are supposed to be the same on Jupiter and Earth

For example $F = ma$ $F' = m'a'$

$$u = \frac{\Delta x}{\Delta t}$$

$$u' = \frac{\Delta x'}{\Delta t'}$$

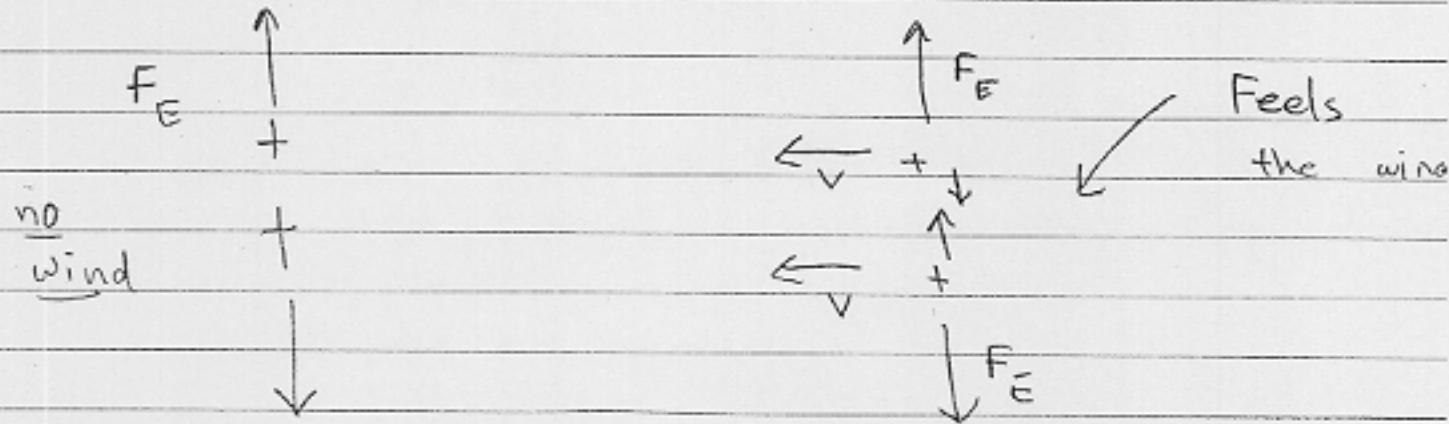
But: $E + M$ Doesn't seem to work



Maxwell's answer:

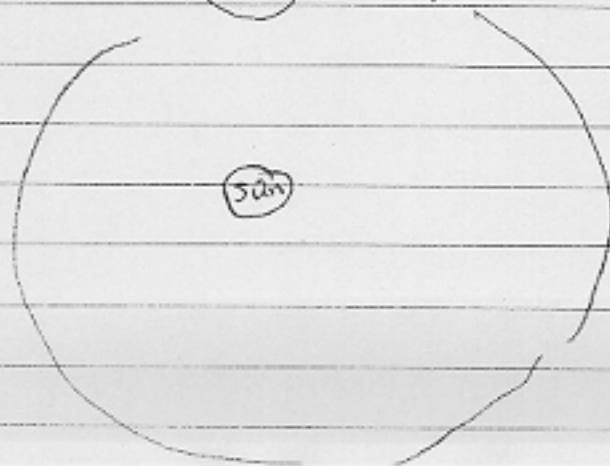
- Based on the thought that light is a disturbance of some medium aether
- Kind of like sound is a disturbance of air. Then there is a preferred frame. Namely the frame where the aether is at rest.

- For instance if you throw a ball straight up when you're standing still you get a different answer than if you throw a ball straight up in a convertible car, because of wind



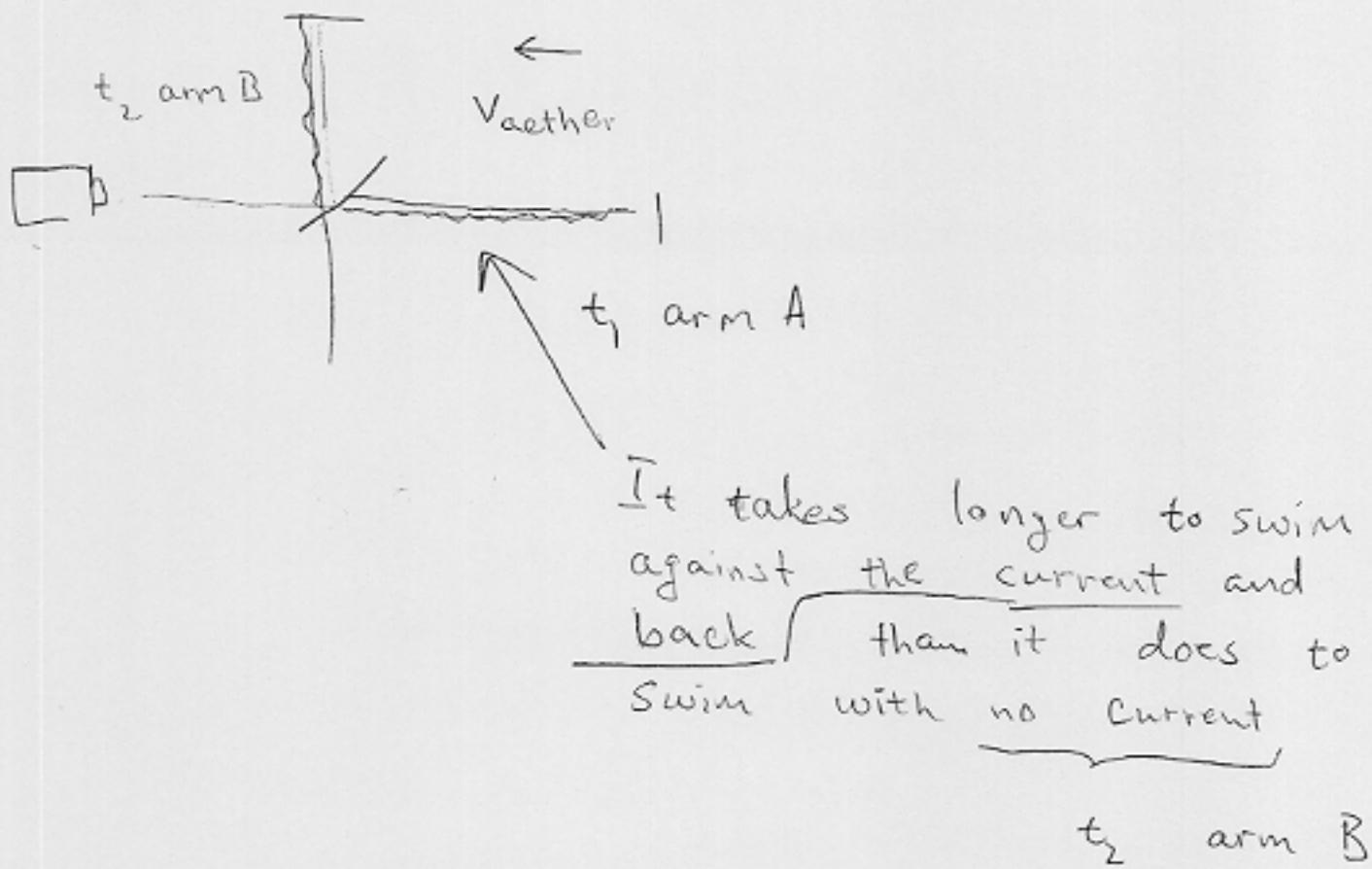
- Need to move near the speed of light to feel the wind
- aether
- There should be a wind all the time

$$\textcircled{(\text{earth})} \rightarrow V_{\text{earth}} = \frac{2\pi R}{1 \text{ year}} = 30 \text{ km/s}$$

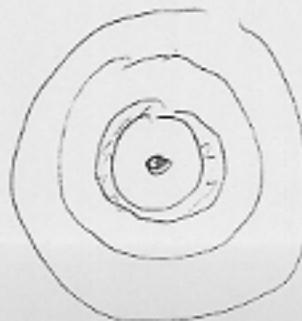


$$V_C \approx 1 \times 10^{-4}$$

The michelson - Morley experiment



- Then the time difference makes the light interfere
- By rotating the table can change the time difference between the arms



- expect the interference pattern to change based on the time difference

- Michelson saw no change in fringes

Einstein and Relativity

- The speed of light is constant in all inertial frames (i.e. observers moving with constant velocity v)
- All Laws of Physics have the same form in all reference frames

Einstein Reasserts the principle of relativity that is ~~to~~ change all unprimed equations to primed ones.

$$v = \frac{\Delta x}{\Delta t} \quad v' = \frac{\Delta x'}{\Delta t'}$$

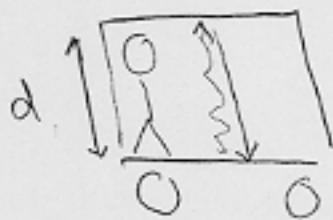
But changes the relation between u and u'

$$u' \neq u - v$$

Time Dilation

- The speed of light is constant in all frames

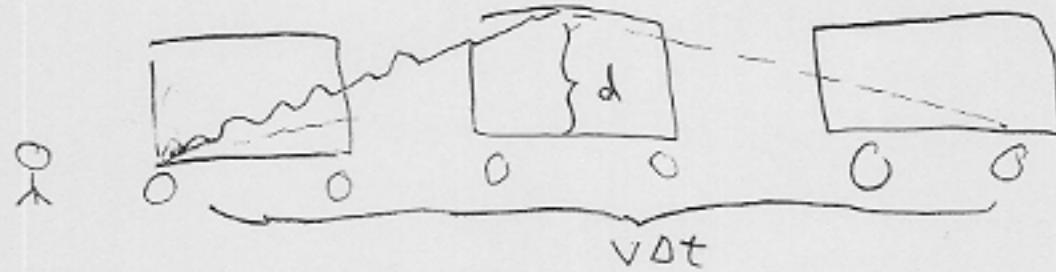
Time it take for person sitting still to throw and catch the light



$$\Delta t = \frac{2d}{c}$$

proper time: time measured at the same position at spacetime

For an observer moving to the left with speed v



Total Distance = speed of Light = c
Total time

$$\frac{2\sqrt{d^2 + (v\Delta t/2)^2}}{\Delta t} = c$$

Solver For Δt :

$$\sqrt{(2d)^2 + (v\Delta t)^2} = c\Delta t$$

$$(2d)^2 + (v\Delta t)^2 = c^2 \Delta t^2$$

$$\left(\frac{2d}{c}\right)^2 + \left(\frac{v}{c}\right)^2 \Delta t^2 = \Delta t'^2$$

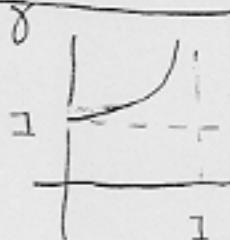
$$\left(\frac{2d}{c}\right)^2 = \left[1 - \left(\frac{v}{c}\right)^2\right] (\Delta t)^2$$

$$\frac{\left(\frac{2d}{c}\right)^2}{\left(1 - \left(\frac{v}{c}\right)^2\right)} = (\Delta t)^2 \quad \Delta t$$

$$\Delta t = \left(\frac{2d}{c}\right) \sqrt{\frac{1}{1 - (v/c)^2}} \Rightarrow \Delta t = \frac{\tau}{\sqrt{1 - (v/c)^2}}$$

Define

$$\gamma = \frac{1}{\sqrt{1 - (v/c)^2}}$$



Moving clocks run slow

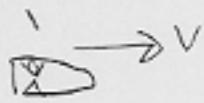
$$\Delta t' = \gamma \Delta t \quad \gamma > 1$$

time for someone on ground

time interval for someone on train

Length Contraction

Consider a Ruler Stick:



ground obs. A - B

L_p



proper length = length as measured by someone at rest w.r.t. the ruler stick

Consider the ground observer he sees that the space-ship takes a time to complete his journey

$$\Delta t = \frac{L_p}{v}$$

$$\Delta T_s = \frac{L_s}{v}$$

Now

$$\gamma \Delta t = \frac{L_p}{v}$$

$$\Delta T = \frac{(L_p/\gamma)}{v} = \frac{\text{length of ruler}}{\text{as seen by space g.}} \frac{v}{v}$$

$$L = \frac{L_p}{\gamma}$$

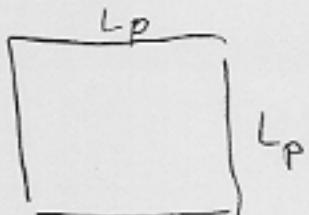
proper length

Moving ruler sticks are length contracted

Remark:

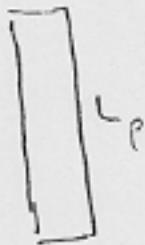
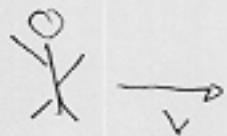
- Only those directions in the direction of motion are length contracted.

Fixed Observer:



- Transverse Directions not contracted.

Moving Observer:



Muon and mountain : Earth Observer



- The μ decays in $2.2 \mu\text{s}$ in its own frame (a proper time)
- To an observer on earth the muon decays in $\Delta t = \gamma \Delta \tau$ $\gamma = \frac{1}{\sqrt{1 - (v/c)^2}} = 7.1$
 $\Delta t = (7.1)(2.2 \mu\text{s}) \approx 16 \mu\text{s}$
- The distance travelled is $d = v \Delta t \approx c \cdot 16 \mu\text{s} = 470$ the muon reaches the bottom!

Muon and mountain: Muon Observer



$$L = L_0 / \gamma = \frac{4700 \text{ m}}{7.1} \approx 650 \text{ m}$$

The muon says $x = \Delta \tau v$ amount of now passes him
 $x = 2.2 \mu\text{s} (0.99c) \approx 650 \text{ m}$

The muon agrees he reaches the
bottom

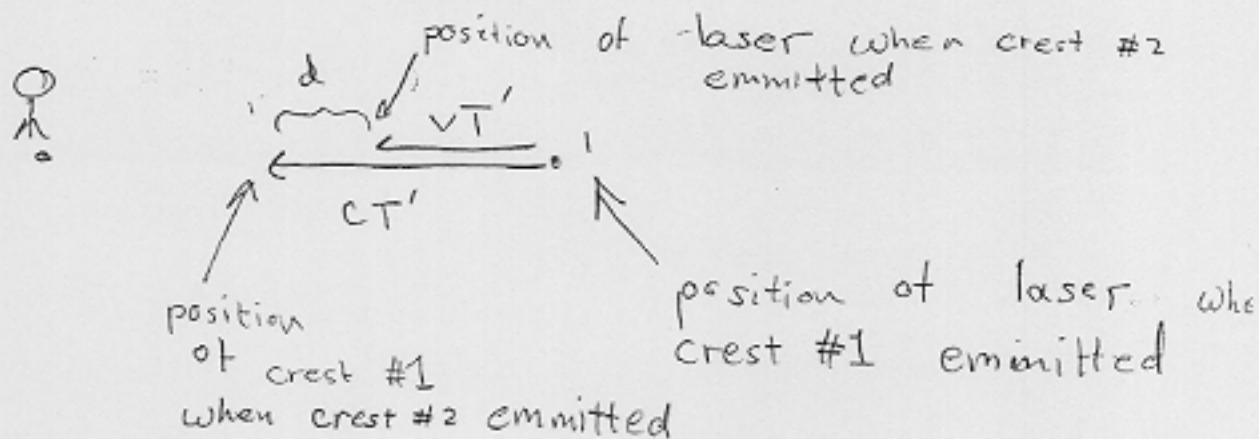
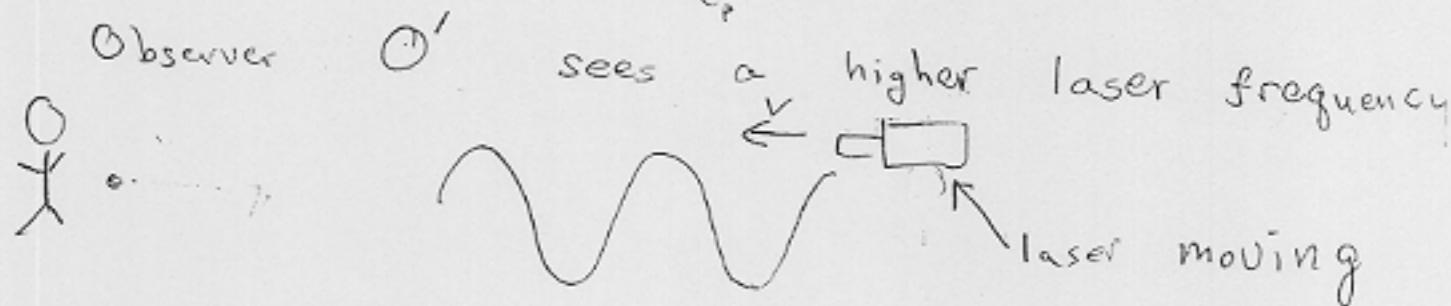
Doppler Shift



laser fixed

- For someone ^{on the laser} The crests of the waves emitted with period T_p :

$$c = \lambda f \quad f = \frac{1}{T_p}$$



distance between crests $= \lambda' = cT' - vT'$

T' = time between crests as seen by O'

$$\tau' = \gamma \tau_p$$

time between crests for a laser fixed
in space

$$\lambda' = (c-v) \tau' = (c-v) \gamma \tau_p$$

Now

$$c = \lambda' f' \Rightarrow f' = c/\lambda'$$

$$f' = \frac{c}{\lambda'} = \frac{c}{(c-v)\gamma \tau_p}$$

$$f' = \frac{1}{(1-v/c)\gamma \tau_p} = \frac{\sqrt{1-(v/c)^2}}{1-v/c} f_0$$

$$\sqrt{1-(v/c)^2} = \sqrt{(1-v/c)(1+v/c)} \quad 80$$

$$f' = f_0 \sqrt{\frac{1+v/c}{1-v/c}}$$