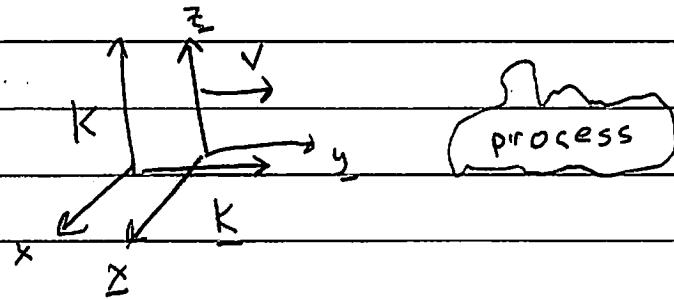


## Last Time

### ① Introduced Relativity



- All laws of physics are the same

$$\frac{dp}{dt} = q \frac{\vec{E}}{c} + \frac{\vec{v}}{c} \times \vec{B}$$

Laws of physics  
in K-frame

$$\frac{dp}{dt} = q \frac{\vec{E}}{c} + \frac{\vec{v}}{c} \times \vec{B}$$

Laws of Phys  
in K frame

- Speed of Light is const

② To relate  $t, \underline{x}$  to  $t, x$  we constructed a change of coords which leaves light speed the same

$$\underline{x}^m = L_{\nu}^m(v) \cdot x^\nu$$

$$L_{\nu}^m = \begin{pmatrix} \gamma & -\gamma\beta \\ -\gamma\beta & \gamma \end{pmatrix}$$

$$\underline{x}_\mu = x_\nu (L^{-1})^\nu_\mu$$

So

$$x_\mu x^m = -(ct)^2 + \vec{x}^2 \text{ is invariant}$$

Last time pg. 1

- ④ Also introduced the concept of a four vector.  $A^{\mu} = (a^0, \vec{a})$ . A four vector is a set of physical quantities which transform under change of frame (observer)

$$A^{\mu} = L^{\mu}_{\nu} (v) A^{\nu}$$

The dot product

$$\begin{aligned} A \cdot B &= A_{\mu} B^{\mu} = a_0 b^0 + a_i b^i \\ &= -a^0 b^0 + \vec{a} \cdot \vec{b} \end{aligned}$$

- ⑤ Formalize the transition between upper and lower indices

$$A_{\mu} = g_{\mu\nu} A^{\nu} \quad g_{\mu\nu} \approx \begin{pmatrix} -1 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix}$$

metric tensor

$$g_{00} = -1 \quad g_{11} = g_{22} = g_{33} = 1 \quad \text{all others zero}$$

Then

$$A \cdot B = A^{\mu} g_{\mu\nu} B^{\nu} \quad \text{is lorentz invariant}$$

## Formalize Raising and Lowering pg. 2

Indices are raised with

$$A^{\mu} = g^{\mu\nu} A_{\nu} \quad \text{so} \quad g^{\mu\nu} = \begin{pmatrix} -1 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix}$$

$$A^{\mu} = \underbrace{g^{\mu\rho} g_{\rho\sigma}}_1 A^{\sigma}$$

$$\delta^{\mu}_{\sigma} = g^{\mu\nu} \delta_{\nu\sigma} = \text{identity matrix} = \begin{pmatrix} 1 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix}$$

Then  $\underline{g}_{\mu\nu}$  same as  $\underline{g}_{\mu\nu} \rightarrow$  that's relativity

$$\underline{A^{\mu}} \stackrel{\checkmark}{\underline{g_{\mu\nu}}} \underline{B^{\nu}} = \underline{A^{\mu}} \underline{g_{\mu\nu}} \underline{B^{\nu}}$$

$$\text{Matrices } \underline{A^T} \underline{g} \underline{B} = \underline{A^T} \underline{g} \underline{B} \quad \underline{A} = L A$$

$$\underline{A^T} \underline{L^T} \underline{g} \underline{L} \underline{B} = \underline{A^T} \underline{g} \underline{B} \quad \underline{A^T} = \underline{A^T} \underline{L^T}$$

So require

$$\underline{L^T} \underline{g} \underline{L} = \underline{g} \quad \text{or} \quad \underline{g} = (\underline{L^{-1}})^T \underline{g} \underline{L^{-1}} = \underline{g}$$

Or since,  $(\underline{L^T})^{-1} = (\underline{L^{-1}})^T$  and  $\underline{g^{-1}} = \underline{g}$

$$\underline{g} \underline{L} \underline{g} = (\underline{L^{-1}})^T \quad \text{or} \quad \boxed{\underline{g} \underline{L^T} \underline{g} = \underline{L^{-1}}}$$

$$\boxed{\begin{aligned} \underline{g_{\rho\tau}} \underline{L^{\mu}} \underline{g^{\nu\sigma}} &= (\underline{L^{-1}})^{\mu}_{\rho} \delta^{\nu\sigma} \\ &\equiv \underline{L^{\mu}}_{\rho}^{\sigma} \end{aligned}}$$

Formalize Raising and Lowering Pg. 3

with this cryptic but standard notation

$$\underline{A}_{\nu} = A_{\mu} (L^{-1})^{\mu}_{\nu}$$

$$= (L^{-1})_{\nu}^{\mu} A_{\mu}$$

$$\underline{A}_{\nu} = L_{\nu}^{\mu} A_{\mu}$$

## Exercise:

- A Tensor transforms as

$$\star \underline{T}^{\mu\nu} = L^{\mu}_p L^{\nu}_o T^{po}$$

- Writing  $\underline{T}^{\mu\nu}$  as a matrix  $(T^{\mu\nu})$   
Show starting from ~~the~~ that

$$(T^{\mu\nu}) = (L)^{\mu}_p (T^{po}) (L^{-1})^o_{\nu}$$

Or

$$\underline{T}^{\mu\nu} = L^{\mu}_p L^{\nu}_o T^{po}$$

Solution: Lowering  $\nu$  by multiplying  $\circledcirc g_{\nu\nu}$ :

$$\underline{T}^{\mu\nu} = L^{\mu}_p L^{\nu}_o T^{po} = L^{\mu}_p L^{\nu}{}^o T^{po}$$

$$= L^{\mu}_p (L^{-1T})^o_{\nu} T^{po}$$

$$\underline{T}^{\mu\nu} = L^{\mu}_p T^{po} (L^{-1})^o_{\nu}$$

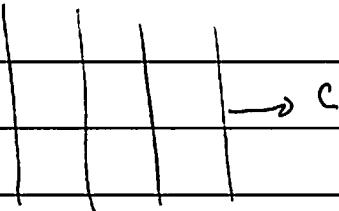
Reverse order of  
 $(\sigma, \nu)$  and  
get rid of transpose

Can note:  $A_{\mu} B^{\mu} = A_{\mu} \delta^{\mu}_{\sigma} B^{\sigma}$

$$= A_{\mu} g^{\mu\nu} g_{\nu o} B^o = A^{\nu} B_{\nu}$$

$$= A^{\mu} B_{\mu}$$

## Example - Relativistic Doppler Shift



$A e^{i\omega t + \vec{k} \cdot \vec{x}}$  : a wave with  $\omega = ck$

Since the speed of Light is constant

$$\phi = -\omega t + \vec{k} \cdot \vec{x} = -\underline{\omega t} + \underline{\vec{k} \cdot \vec{x}}$$

must be constant. This shows that  $\frac{(\omega, \vec{k})}{c}$  form a four vector

$$K^m = \left( \frac{\omega}{c}, \vec{k} \right)$$

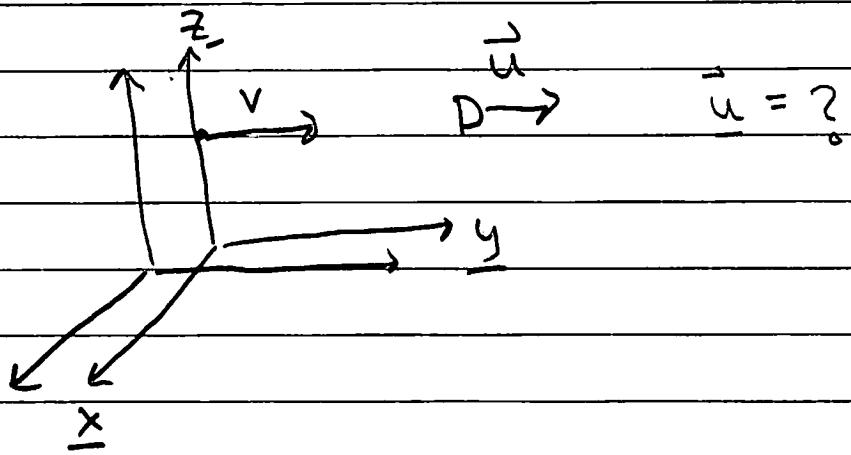
$$\text{So } K_m X^m = \underline{K_m} \underline{X^m} = -\omega t + \vec{k} \cdot \vec{x} \text{ is const.}$$

• So if you know the frequency in one frame can find the frequency in another

• Then  $K^2$  is lightlike (means  $K^2 = 0$ )

$$K^2 = -\left(\frac{\omega}{c}\right)^2 + \vec{k}^2 = 0 \quad \text{in all frames}$$

## Velocity Transformation Rule



$$\vec{u} = \frac{d\vec{x}}{dt} \quad \text{and} \quad \underline{\vec{u}} = \frac{d\underline{\vec{x}}}{dt}$$

Use  $\underline{x}^m = L^m \downarrow x^v$  Find

$$\underline{\vec{u}''} = \frac{\vec{u}'' - \vec{v}}{(1 - \vec{u} \cdot \vec{v}/c^2)} \quad \underline{u}_\perp = \frac{u_\perp}{\gamma(1 - \vec{u} \cdot \vec{v}/c^2)}$$

This is complicated because both the numerator and denominator transform

$$\frac{dx}{dt} \leftarrow \begin{array}{l} \text{numerator} \\ \text{denominator} \end{array}$$

## Proper Time

- Consider a particle moving with velocity  $\vec{u}$

$$dx^m = (dx^0, \vec{dx}) = dx^0 (1, \beta_m)$$

$$\beta_m = \frac{\vec{u}}{c} = \frac{d\vec{x}}{dx^0} \quad \leftarrow \text{put } \beta_u = \frac{u}{c} \text{ and } \gamma_u = \frac{1}{\sqrt{1-(u/c)^2}}$$

- Then:

to distinguish from

$$\beta = v/c \text{ and } \gamma = \frac{1}{\sqrt{1-(v/c)^2}}$$

$$ds^2 \equiv dx^m dx^m = -c^2 dt^2 (1 - \beta_m)^2$$

is invariant. In the rest frame of particle have

$$ds^2 = -c^2 d\tau^2 \leftarrow \text{proper time}$$

So

$$d\tau = dt \sqrt{1 - \beta_m^2} = \boxed{\frac{dt}{\gamma_m} = d\tau}$$

- Define the four velocity, as the distance per proper time

$$u^m = \frac{dx^m}{d\tau} \leftarrow \text{this is a four vector}$$

$$u^m = \gamma_m \frac{dx^m}{dt} = (\gamma_c, \gamma \vec{u}) \quad \frac{d\vec{x}}{dt} = \vec{u}$$

## A preview of 4-momentum

We will see:

$$\boxed{P^{\mu} = m u^{\mu}}$$

↑  
four momentum

$$\boxed{\left(\frac{E}{c}, \vec{p}\right) \equiv P^{\mu}}$$

↑ four velocity

$$P^{\mu} P_{\mu} = m^2 u^{\mu} u_{\mu}$$

$$u^{\mu} u_{\mu} = -c^2$$

$$= -m^2 c^2$$

$$- \gamma^2 c^2 + \gamma^2 u^2$$

Then:

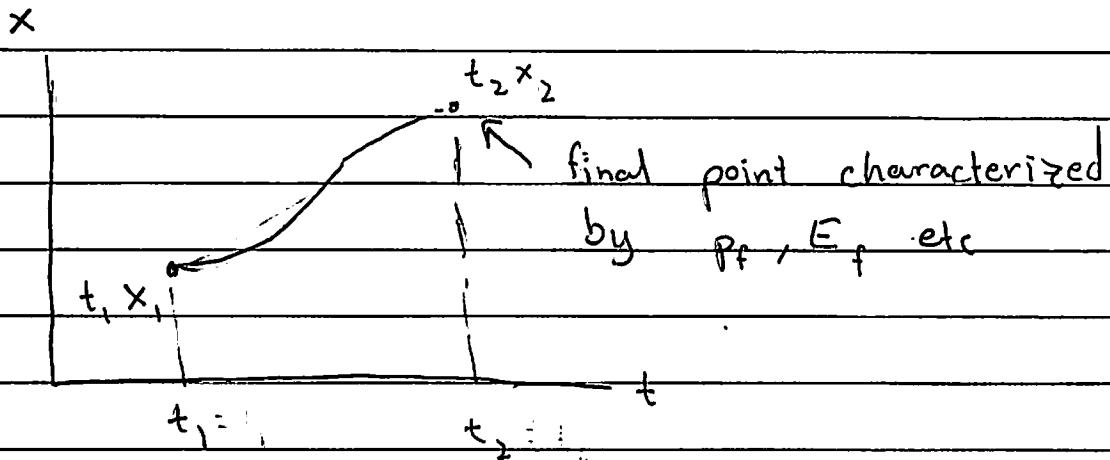
$$P^{\mu} P_{\mu} = -\left(\frac{E}{c}\right)^2 + \vec{p}^2 = -m^2 c^2$$

$$\Rightarrow E = \sqrt{(mc^2)^2 + (\gamma p)^2}$$

and

$$\boxed{\vec{u} = \frac{\vec{p}}{(E/c^2)} = \frac{\gamma m \vec{u}}{(\gamma m c^2)/c^2} = \vec{u}}$$

## Energy and Momentum



Classical Dynamics minimizes the action

$$I = \int dt L[x, \dot{x}]$$

Look at the "on-shell action" or Hamilton-Jacobi function. Means you find the solution  $x(t) = x_s(t)$  and evaluate I.

$$S(t_1, x_1; t_2, x_2) = I \quad | \quad = \int dt L(x_s, \dot{x}_s(t))$$
$$x(t) = x_s(t)$$

How does S change as we change  $t_2, x_2$

Answer :

$$\delta S = -E_f \delta t_2 + p_f \delta x_2$$

## Energy + Momentum pg. 2

Prf

$$S = \int L dt = \int (p dq - H) dt$$

$$S = \int p dq - H dt$$

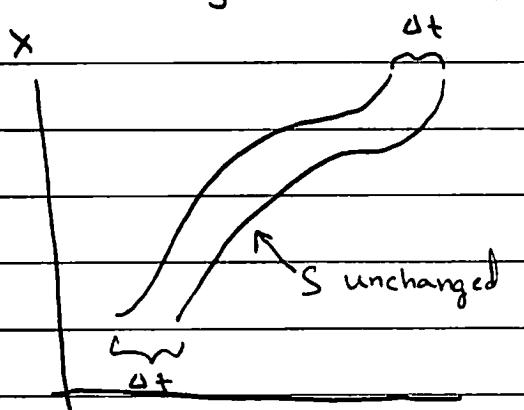
So if you change  $\delta q$  and  $\delta t$

$$\delta S_2 = -E_f \delta t_2 + p_f \delta x_2$$

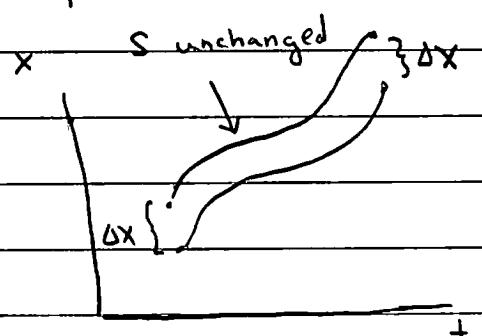
Similarly if you change  $t_i$  and  $x_i$

$$-\delta S_i = -E_i \delta t_i + p_i \delta x_i$$

Now for a system which has no preferred time origin (i.e. the resulting motion is independent of when you start), then the action integral



is unchanged by changing the starting and final times by  $\Delta t$ ,  $\Delta S = 0$



Similarly if the system has no spatial origin, then the action is unchanged by shifting the initial and final points by  $\Delta x$   $\Delta S = 0$

## Energy and Momentum pg. 3

So

$$\Delta S = \Theta = - (E_f - E_i) \Delta t + (p_f - p_i) \Delta x$$

And then energy and momentum  
are conserved

$$E_f = E_i$$

$$p_f = p_i$$

So energy and momentum is conservation  
is a consequence of the homogeneity  
of space-time. It should be true for all  
observers.

$$\delta S = - E \delta t + p \delta x = - \underline{E} \underline{\delta t} + \underline{p} \underline{\delta x}$$

Thus if  $(c \delta t, \delta x)$  is a four vector

- then,  $\underline{P}^m = (\frac{E}{c}, \vec{p})$ , should also be

a four vector, so we have the invariant,

$$P_m \delta x^m = - E \delta t + p \delta x$$

This guarantees the energy and momentum  
conservation in all frames.

## Energy and Momentum pg. 4

Using for small velocity  $p = mv$

$$\begin{pmatrix} E/c \\ p \end{pmatrix} = \begin{pmatrix} 1 & \beta \\ \beta & 1 \end{pmatrix} \begin{pmatrix} E/c_{\text{rest}} \\ 0 \end{pmatrix}$$

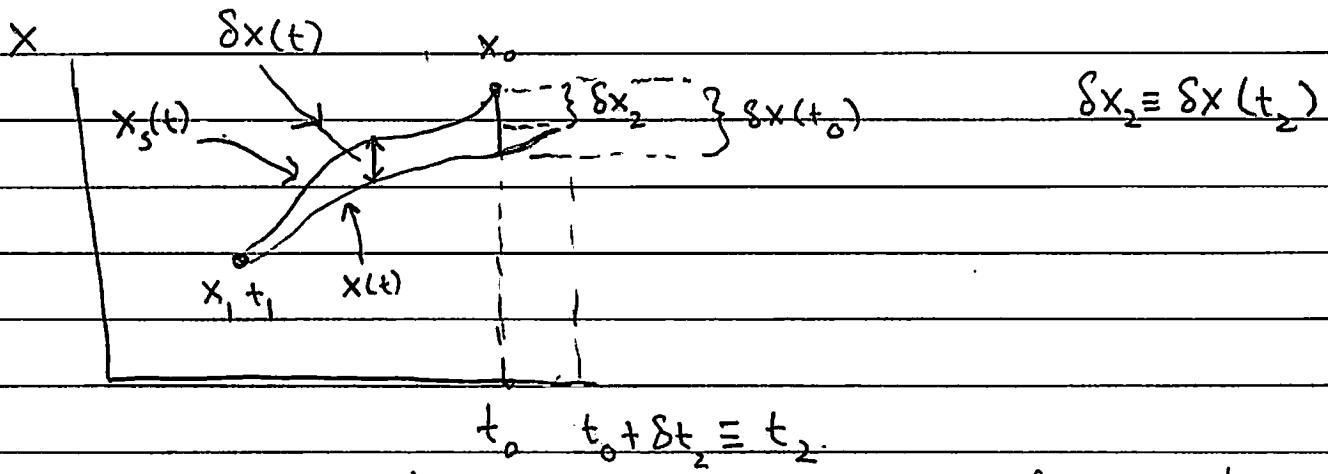
$$\begin{pmatrix} E/c \\ p \end{pmatrix} = \begin{pmatrix} 1 & \beta \\ \beta & 1 \end{pmatrix} \begin{pmatrix} E_{\text{rest}}/c \\ 0 \end{pmatrix}$$

$$p = \frac{E_{\text{rest}}}{c} \beta \approx \frac{E_{\text{rest}}}{c^2} v$$

So must identify  $E_{\text{rest}} = mc^2$ , and

$$P^m P_\mu = -m^2 c^2$$

## Exercise



- Work entirely within the Lagrangian framework:

$$S = \int_{t_1}^{t_0 + \delta t} L [x_s(t) + \delta x(t)] dt$$

Show that

$$\delta S = -E \delta t_2 + p \delta x_2 = \left( \frac{\partial S}{\partial t_2} \right)_{x_2} \delta t_2 + \left( \frac{\partial S}{\partial x} \right)_{t_2} \delta x_2$$

Note:  $x(t) = x_s(t) + \delta x(t)$

$$x(t_0) = x_s(t_0) + \delta x(t_0)$$

$$\approx x(t_0 + \delta t) - \dot{x}(t_0 + \delta t) \delta t = x_s(t_0) + \delta x(t_0)$$

$$[x(t_0 + \delta t) - x_s(t_0)] - \dot{x}(t_0) \delta t = \delta x(t_0)$$

$$\delta x_2 - \dot{x}(t_2) \delta t = \delta x(t_0)$$

## Solution to Exercise

$$SS = \mathcal{L} \cdot \delta t + \int_{t_1}^{t_0} \mathcal{L} (x + \delta x)$$

$$= \mathcal{L} \cdot \delta t + \int_{t_1}^{t_0} \left[ \frac{\partial \mathcal{L}}{\partial \dot{x}} \dot{x} + \delta x + \frac{\partial \mathcal{L}}{\partial x} \delta x \right]$$

integrate

by parts

$$= \mathcal{L} \delta t + \left. \frac{\partial \mathcal{L}}{\partial \dot{x}} \delta x \right|_{t_1}^{t_0} + \int_{t_1}^{t_0} dt \left[ -\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{x}} + \frac{\partial \mathcal{L}}{\partial x} \right] \delta x$$

= 0 by EOM

$$= \mathcal{L} \delta t + \left. \frac{\partial \mathcal{L}}{\partial \dot{x}} \delta x(t_0) \right.$$

Now

$$\delta x_2 = \delta x(t_0) + \dot{x}(t_0) \delta t_2 \quad (\text{see picture})$$

So

$$SS = \left( \mathcal{L} - \frac{\partial \mathcal{L}}{\partial \dot{x}} \dot{x} \right) \delta t_2 + \frac{\partial \mathcal{L}}{\partial x} \delta x_2$$

$$SS = -E \delta t + p \delta x$$

manifest

## Covariant Formulation

$$\textcircled{1} \quad \partial_\mu \equiv \frac{\partial}{\partial x^\mu} = \left( \frac{1}{c} \frac{\partial}{\partial t}, \vec{\nabla} \right) \quad \leftarrow \text{is a covariant vector}$$

Also  $\partial^\mu \equiv \frac{\partial}{\partial x_\mu} = \left( -\frac{1}{c} \frac{\partial}{\partial t}, \vec{\nabla} \right)$

$$\boxed{\partial_\mu \partial^\mu = -\frac{1}{c^2} \frac{\partial^2}{\partial t^2} + \nabla^2 \equiv \square} \quad \text{is invariant}$$

\textcircled{2} Then the continuity eqn:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \vec{J} = 0 \Rightarrow \frac{1}{c} \frac{\partial(\rho c)}{\partial t} + \nabla \cdot \vec{J} = 0$$

So take  $J^\mu = (c\rho, \vec{J})$  to be a four vector

$$\boxed{\partial_\mu J^\mu = 0}$$

\textcircled{3} Then the equations for the gauge potential  
 $-\square \Psi = J^0/c$  read

$$-\square \vec{A} = \vec{J}/c$$

Together

$$\frac{1}{c} \frac{\partial \Psi}{\partial t} + \nabla \cdot \vec{A} = 0$$

## Covariance pg. 2

So in order to have a Lorentz invariant theory take  $(\varphi, \vec{A})$  to be a four vector

$$A^{\mu} = (\varphi, \vec{A})$$

Then the wave eq becomes

$$-\square A^{\mu} = J^{\mu}/c$$

$$\partial_{\mu} A^{\mu} = 0$$

Lorentz gauge condition

④ Now the fields

$$E^i = -\frac{\partial A^i}{c \partial t} - \frac{\partial \varphi}{\partial x^i} = \underbrace{\partial^0 A^i}_{= F^{0i}} - \underbrace{\partial^i A^0}$$

$$= F^{0i}$$

$$B_k = (\nabla \times A)_k$$

$$\epsilon^{ijk} B_k = \epsilon^{ijk} \underbrace{\epsilon_{klm}}_{(S^{il} S^{jm} - S^{il} S^{km})} \partial_l A_m$$

$$\epsilon^{ijk} B_k = \underbrace{\partial^i A^j - \partial^j A^i}_{= F^{ij}}$$

## Covariance pg. 3

So we can see

$$F^{\alpha\beta} = \partial^\alpha A^\beta - \partial^\beta A^\alpha$$

$\beta \longrightarrow$

$$F^{\alpha\beta} = \begin{vmatrix} 0 & E^x & E^y & E^z \\ -E^x & 0 & B^z & -B^y \\ -E^y & B^z & 0 & B^x \\ -E^z & -B^y & -B^x & 0 \end{vmatrix}$$

Note  $F^{\alpha\beta}$  is a second rank anti-symmetric tensor and transforms in the following way

$$F^{\mu\nu} = L^\mu_\alpha L^\nu_\beta F^{\alpha\beta}$$

(5) Now the EOM (Part I)

$$\nabla \cdot E = \rho \Rightarrow -\partial_i F^{i0} = \rho$$

$$-\frac{1}{c} \partial_t E + \nabla \times B = J/c \Rightarrow -\left(\frac{\partial F^{oi}}{\partial x^0} + \frac{\partial F^{li}}{\partial x^i}\right) = J/c$$

Or

$$-\partial_\alpha F^{\alpha\beta} = \frac{J^\beta}{c}$$

## Covariance pg. 4

(6) The remaining two Maxwell eqs. follow from the fact that

$$F^{\alpha\beta} = \partial^\alpha A^\beta - \partial^\beta A^\alpha$$

which we derived from:

$$\begin{aligned} \nabla \cdot \vec{B} &= 0 & \text{or} \\ -\nabla \times \vec{E} &= -\frac{1}{c} \frac{\partial \vec{B}}{\partial t} \quad \Rightarrow \quad -\frac{1}{c} \frac{\partial \vec{B}}{\partial t} - \nabla \times \vec{E} = 0 \end{aligned}$$

Comparison with the other two Maxwell eqs in the absence of currents

$$\left. \begin{aligned} \nabla \cdot \vec{E} &= 0 \\ -\frac{\partial \vec{E}}{\partial t} + \nabla \times \vec{B} &= 0 \end{aligned} \right\} \quad \text{So if } \vec{E}, \vec{B} \text{ are solutions then} \\ \vec{E}' \rightarrow \vec{E}' = -\vec{B} \text{ and } \vec{B}' \rightarrow \vec{B}' = -\vec{E}$$

Shows that we want to replace  $\vec{E} \rightarrow \vec{B}$  and  $\vec{B} \rightarrow -\vec{E}$  in the field strength

$$\tilde{F}^{\mu\nu} = \left( \begin{array}{c|ccc} 0 & B^x & B^y & B^z \\ \hline -\vec{B} & 0 & -E^x & E^y \\ & E^x & 0 & -E^z \\ & E^y & E^z & 0 \end{array} \right)$$

So

$$-\partial_\mu \tilde{F}^{\mu\nu} = 0$$

## Manifest Covariance part (6) continued pg.5

Then  $\tilde{\epsilon}^{\mu\nu\alpha\beta}$  implements  $\vec{B} \rightarrow -\vec{E}$  and  $\vec{E} \rightarrow \vec{B}$

$$\tilde{F}^{\mu\nu} = \frac{1}{2} \epsilon^{\mu\nu\alpha\beta} F_{\alpha\beta}$$

where  $\epsilon^{\mu\nu\alpha\beta} = \begin{cases} +1 & \text{for even perm of } 0,1,2,3 \\ -1 & \text{for odd perm of } 0,1,2,3 \\ 0 & \text{otherwise} \end{cases}$

$$-\partial_\mu \tilde{F}_{\mu\nu} = 0$$

This can also be written in terms of  $F$

$$\boxed{\partial_\mu F_{\mu\nu} + \partial_\nu F_{\nu\rho} + \partial_\rho F_{\rho\mu} = 0}$$

↑  
This is the Bianchi - Identity

## Covariance pg. 6

⑦ Finally the Force Law

$$\frac{dE}{dt} = q \vec{E} \cdot \vec{v} = q F^{oi} v^i$$

$$\frac{dp^i}{dt} = q \vec{E} + \frac{\vec{v}}{c} \times \vec{B} = q \left( F^i_0 + F^i_j \frac{v^j}{c} \right)$$

So

$$\boxed{\frac{dp^u}{dt} = q F^u_v u^v}$$