

Problem 1. Levi-civita practice

- (a) Using the Levi-civita tensor, show that for a constant field magnetic \mathbf{B} field show that the vector potential ($\mathbf{B} = \nabla \times \mathbf{A}$) can be written:

$$\mathbf{A} = \frac{-1}{2} \mathbf{r} \times \mathbf{B} \quad (1)$$

- (b) Show (using the Levi-civita symbol) and $\epsilon_{ijk}\epsilon^{ijk} = 3!$ that

$$\det(\mathbb{A}) = \det(\mathbb{A}^T) \quad (2)$$

and

$$\det(\mathbb{A}\mathbb{B}) = \det(\mathbb{A})\det(\mathbb{B}) \quad (3)$$

- (c) When differentiating $1/r$ we write

$$\frac{1}{r} = \frac{1}{\sqrt{x^i x_i}} \quad (4)$$

with $\mathbf{x} = x^i \mathbf{e}_i$, and use results like

$$\partial_i x^j = \delta_i^j \quad \partial_i x_j = \delta_{ij} \quad \partial_i x^i = \delta_i^i = d = 3 \quad (5)$$

where $d = 3$ is the number of spatial dimensions. (It is usually helps to write this as d rather than 3 to get the algebra right). In this way, one finds that field due to an electric point charge (monopole) is the familiar $\hat{\mathbf{r}}/r^2$. Go through these steps!

$$j\text{-th component of } -\nabla(1/r) = \left(-\nabla \frac{1}{r}\right)_j = -\partial_j \frac{1}{\sqrt{x^i x_i}} = \frac{\frac{1}{2}(x^i \delta_{ji} + x_i \delta_j^i)}{(x^k x_k)^{3/2}} = \frac{x_j}{r^3} = \frac{(\mathbf{n})_j}{r^2} \quad (6)$$

where $\hat{\mathbf{r}} \equiv \mathbf{n} = \mathbf{r}/r$. In general

$$\partial_j r^\alpha = \alpha r^{\alpha-1} n_j. \quad (7)$$

Using tensor notation (*i.e.* indexed notation) show that

$$\nabla \times \frac{\hat{\mathbf{r}}}{r^2} = 0 \quad (8)$$

This verifies that the $\nabla \times \mathbf{E} = 0$ for a point charge.

- (d) The vector potential of a magnetic dipole is

$$\mathbf{A} = \frac{\mathbf{m} \times \mathbf{n}}{4\pi r^2} \quad (9)$$

where \mathbf{m} is a constant vector known as the magnetic dipole moment and $\mathbf{n} = \mathbf{r}/r$. Recall that $\mathbf{B} = \nabla \times \mathbf{A}$. Using the tensor notation (*i.e.* indexed notation) show that

$$\mathbf{B} = \frac{3(\mathbf{n} \cdot \mathbf{m})\mathbf{n} - \mathbf{m}}{4\pi r^3} \quad (10)$$

Problem 2. Easy important application of Helmholtz theorems

- (a) We showed in class that the source free Maxwell equations (*i.e.* those without ρ and \mathbf{j}), are solved by writing \mathbf{E} and \mathbf{B} in terms of a scalar field Φ (the scalar potential) and a vector field \mathbf{A} (the vector potential)

$$\mathbf{B} = \nabla \times \mathbf{A} \quad (11)$$

$$\mathbf{E} = -\frac{1}{c} \partial_t \mathbf{A} - \nabla \Phi \quad (12)$$

Now, using the sourced Maxwell equations (*i.e.* those with ρ and \mathbf{j}), show that current must obey the conservation Law

$$\partial_t \rho + \nabla \cdot \mathbf{j} = 0, \quad (13)$$

to be consistent with the Maxwell equations.

Problem 3. The multipole expansion and a rotating quadrupole

Consider a charge density $\rho(\mathbf{x})$ as shown below. The potential is given by

$$\phi(\mathbf{r}) = \int d^3\mathbf{x} \frac{\rho(\mathbf{x})}{4\pi|\mathbf{r} - \mathbf{x}|} \quad (14)$$

The multipole expansion determines the potential far from the charges, i.e. for $\mathbf{r} \gg \mathbf{x}$.

- (a) Show that for $\mathbf{r} \gg \mathbf{x}$ we have to quadratic order in \mathbf{x} (or more formally x^i/r) the expansion

$$\frac{1}{|\mathbf{r} - \mathbf{x}|} = \frac{1}{\sqrt{r^2 - 2\mathbf{r} \cdot \mathbf{x} + \mathbf{x} \cdot \mathbf{x}}} \simeq \frac{1}{r} + \frac{r_i x^i}{r^3} + \frac{r_i r_j}{2r^4} (3x^i x^j - x^2 \delta^{ij}) + \dots \quad (15)$$

where $x^2 = \mathbf{x} \cdot \mathbf{x} = x_\ell x^\ell$, and we are using some power series expansions which all physics students must know by heart but don't always¹.

Using this expansion, confirm for yourself (but do not bother turning in!) that the potential far from a charge distribution takes the form

$$\phi(\mathbf{r}) = \frac{1}{4\pi} \left[\frac{q_{\text{tot}}}{r} + \frac{p^i n_i}{r^2} + \frac{1}{2} \frac{Q^{ij} n_i n_j}{r^3} + \dots \right] \quad (18)$$

Here $n_i = r_i/r$ is unit vector in the direction of \mathbf{r} , and the monopole, dipole, and quadrupole moments, are respectively

$$q_{\text{tot}} = \int_V d^3\mathbf{x} \rho(\mathbf{x}) \quad (19)$$

$$p^i = \int_V d^3\mathbf{x} \rho(\mathbf{x}) x^i \quad (20)$$

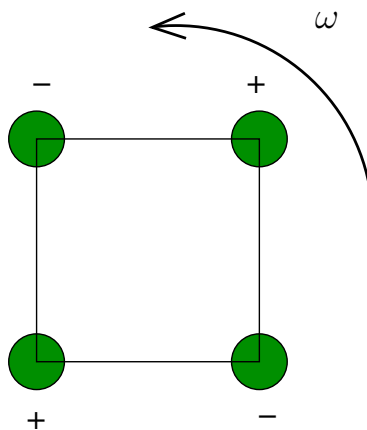
$$Q^{ij} = \int_V d^3\mathbf{x} \rho(\mathbf{x}) (3x^i x^j - x^2 \delta^{ij}) \quad (21)$$

- (b) Consider four point charges of charge $+q, -q, +q, -q$ arranged in a square of side $2a$ lying flat in the x, y plane. The square is rotating with constant angular velocity ω in a counter clockwise fashion, and at time $t = 0$ is in the configuration shown below.

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$$(1+z)^\alpha \simeq 1 + \alpha z + \frac{\alpha(\alpha-1)}{2!} z^2 + \dots \quad (16)$$

$$\log(1+z) \simeq z - \frac{z^2}{2} + \frac{z^3}{3} + \dots \quad (17)$$



- (i) Determine all components of the quadrupole tensor at $t = 0$.
- (ii) In class we said that under rotations the components of the quadrupole tensor transform as

$$\underline{Q}^{ij} = R^i_{\ell} R^j_m \underline{Q}^{\ell m} \quad (22)$$

Use this transformation rule to show that the components of the quadrupole tensor as a function of time are given by

$$(\underline{Q})^{ij} = 12qa^2 \begin{pmatrix} -\sin(2\omega t) & \cos(2\omega t) & 0 \\ \cos(2\omega t) & \sin(2\omega t) & 0 \\ 0 & 0 & 0 \end{pmatrix} \quad (23)$$

Give a one (or at most two) sentence explanation why \underline{Q}^{xx} is negative when $\omega t = \frac{\pi}{4}$.
 Next semester you may be asked to use this result to determine the power radiated by such a rotating array of charges.

Problem 4. Integration by Parts (IBP) like mad!

Answer briefly

- (a) Assume ϕ and $|\mathbf{G}|$ fall faster than $1/r$ as $\mathbf{r} \rightarrow \infty$. Let $\mathbf{F} = \nabla\phi$ and $\nabla \times \mathbf{G} = 0$, use indices and IBP like mad to show that $\int d^3x \mathbf{F} \times \mathbf{G} = 0$
- (b) Consider a current density \mathbf{j} entirely contained within a volume V . The current is steady and therefore satisfies $\nabla \cdot \mathbf{j} = 0$. Show that²

$$\int_V d^3x j^\ell(\mathbf{x}) = 0 \quad (24)$$

- (c) Consider a two dimensional surface S bounded by a loop C . Show that³

$$\int dS_i = \frac{1}{2} \oint_C (\mathbf{r} \times d\boldsymbol{\ell})_i \quad (25)$$

Here the magnitude $d\mathbf{S}$ is the area of the infinitesimal surface element. $d\mathbf{S}$ points normal to the surface.

- (d) Let S be the surface that bounds a volume V . Show that

$$\oint dS_i = 0 \quad (26)$$

and

$$\frac{1}{3} \oint d\mathbf{S} \cdot \mathbf{r} = V \quad (27)$$

²Hint: write $j^i \delta_i^\ell = j^i \partial_i x^\ell$ and IBP like mad!

³Hint: use the results problem 1(a) for the constant vector \mathbf{e}_i .