

Last Times:

For Large nuclei:

$$\rho = \frac{A}{\frac{4}{3}\pi R^3} \rightarrow \text{constant} = \frac{1}{6} \frac{\text{nuc}}{\text{fm}^3}$$

- means that instead of considering finite nuclei, can consider nuclear matter



We looked At The masses

$$M(A)c^2 = Z m_p c^2 + N m_n c^2 - B, \quad \frac{B}{A} \approx 8.5 \frac{\text{MeV}}{\text{nucleon}}$$

Compare (neglect mass diff bet ween protons & neutro

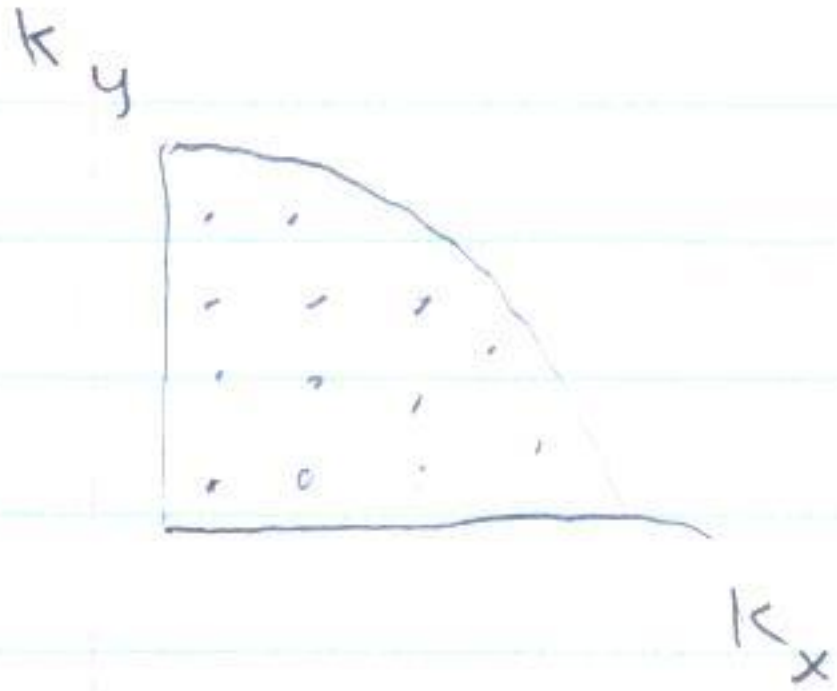
$$\frac{\text{Energy}}{\text{per nucleon}} = \frac{M(A)c^2}{A} = \bar{m}_N - B$$

also

$$\frac{\text{Energy}}{\text{Nucleon}} = \bar{m}_N + \frac{KE}{\text{Nucleon}} + \frac{PE}{\text{Nucleon}}$$

$$-B = \frac{KE}{\text{Nucleon}} + \frac{PE}{\text{Nucleon}}$$

We ~~evaluated~~ ^{estimated} the Kinetic Energy per particle with the Fermi-gas



For fixed Volume and fixed Number we fill all states up to the Fermi momentum

$$\left(\frac{N}{V}\right)^{\frac{1}{3}} \propto \frac{1}{\Delta x} = \frac{1}{\text{distance between particles}}$$

$$\frac{P_F}{\hbar} \sim \frac{1}{\Delta x} \propto \left(\frac{N}{V}\right)^{\frac{1}{3}}$$

Exactly
$$\frac{P_F}{\hbar} = \left(6\pi^2 \frac{N}{V}\right)^{\frac{1}{3}}$$

← density of ^{say} protons with spin up

$$\frac{N}{V} = \frac{1}{4} \cdot \frac{1}{6} \frac{\text{Nucleons}}{\text{fm}^3}$$

$$P_F \approx 200 \text{ MeV}$$

$$E_F = \frac{P_F^2}{2m} = 35 \text{ MeV}$$

Finally used the Fermi Gas model

$$\langle \frac{KE}{A} \rangle = \frac{3}{5} \epsilon_F = 21 \text{ MeV}$$

Then

$$\underbrace{-8 \text{ MeV}} \quad \underbrace{21 \text{ MeV}} \quad \underbrace{-29 \text{ MeV}}$$

$$-B = \frac{KE}{A} + \frac{PE}{A}$$

• Ultra Important:

- Nuclei are the result of a sharp cancellation of Kinetic and Potential Forces

$$\frac{B}{A} \approx \overset{\sim 8.5 \text{ MeV}}{a_v}$$

$$B = a_v A - \underbrace{a_s A^{2/3}}_{\text{Surface}} - \underbrace{a_c \frac{Z(Z-1)}{A^{1/3}}}_{\text{Coulomb}} - a_{\text{sym}} \frac{(Z-2Z)^2}{A}$$

0.72 MeV

$$+ a_{\text{sym}} \frac{(A - 2Z)^2}{A} + \delta$$

↑
tends to make
number of neutrons
= protons

↑
Nuclei like
to have even
of protons
and neutrons
→ well get to it

$$a_v = 15.8 \text{ MeV}$$

$$a_s = 17.8 \text{ MeV}$$

$$a_c = 0.71 \text{ MeV}$$

$$a_{\text{sym}} = 23 \text{ MeV}$$

Pairing

- There are four nuclei with odd N and Z
 ${}^2\text{H}$, ${}^6\text{Li}$, ${}^{10}\text{B}$, ${}^{14}\text{N}$
- 167 species with even N and Z

$$\delta = \begin{cases} +a_p A^{-3/4} & Z \text{ even} - N \text{ even} \\ 0 & \text{even-odd} \\ -a_p A^{-3/4} & \text{odd-odd} \end{cases}$$

$$B = a_v A - a_s A^{2/3} - a_c \frac{Z(Z-1)}{A^{1/3}} - a_{\text{sym}} \frac{(A-2Z)^2}{A} + \delta$$

Then

$$M(Z, A) = Z m({}^1\text{H}) + N m_N - B(Z, A)/c^2$$

$$Z_{\text{min}} = \frac{[m_N - m({}^1\text{H})] + a_c A^{-1/3} + 4a_{\text{sym}}}{2a_c A^{-1/3} + 8a_{\text{sym}}/A}$$

$$= \left. \frac{\partial M}{\partial Z} \right|_{A \text{ fixed}}$$

$$Z_{\min} \approx \frac{A}{2} \frac{1}{\left(1 + \frac{1}{4} A^{2/3} a_c / a_{\text{sym}}\right)}$$

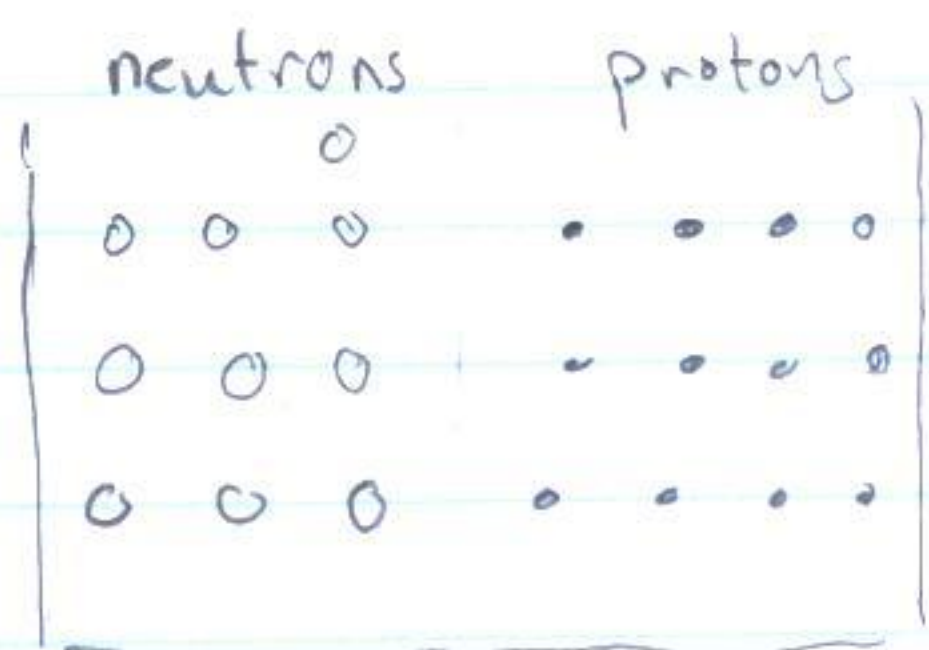
Then this explains the following

- Small nuclei have $Z \approx \frac{A}{2}$
- As the nucleus grows number of neutrons is larger than number of protons to offset the increase of coulomb
- heavy nuclei, approximately correct

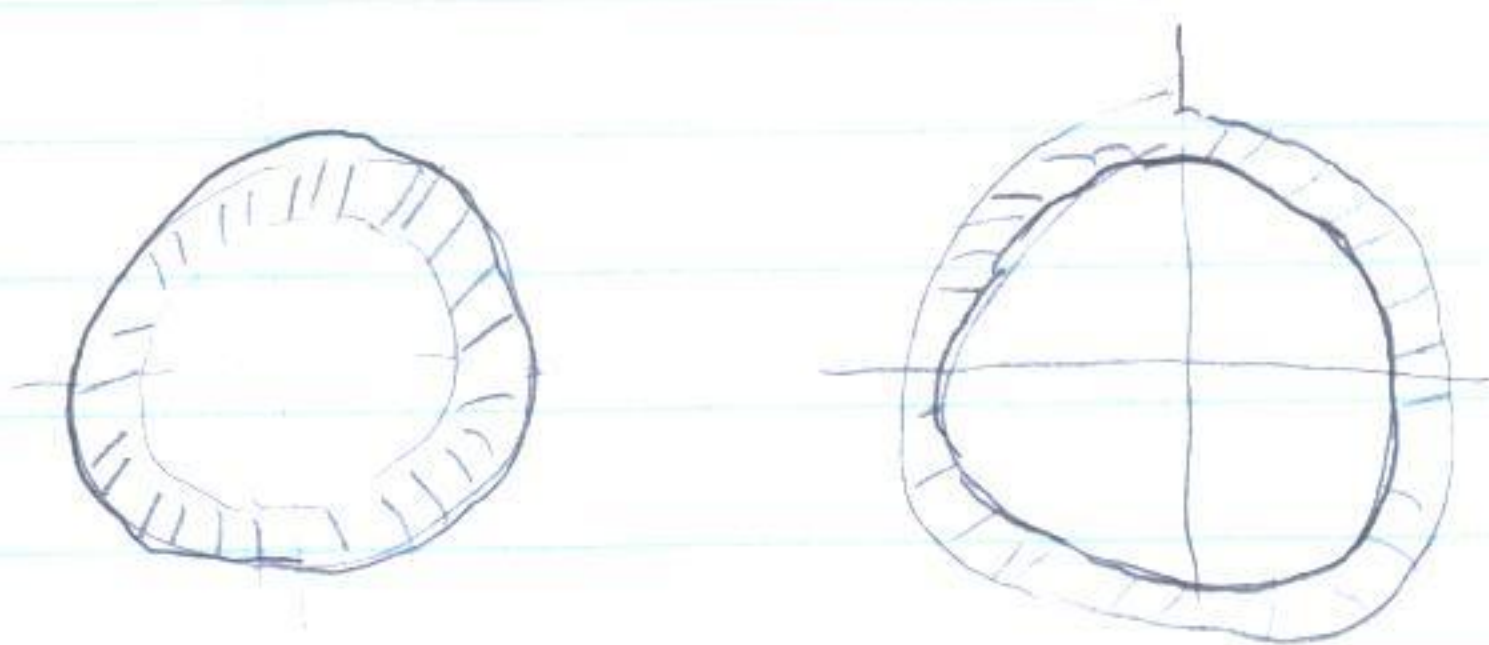
$$\frac{Z}{A} \approx 0.41$$

$$\frac{Z_{\min}}{A/2} \approx \frac{1}{\left(1 + \frac{1}{4} A^{2/3} a_c / a_{\text{sym}}\right)} = 0.41 \quad \Big| \quad A=140$$

Then consider



If we convert a proton to a neutron then we will add to the fermi sea



$$N = \left(\frac{g}{2\pi^2} \right) \frac{1}{3} \left(\frac{p_f}{\hbar} \right)^3 \cdot V \quad E = \frac{g}{2\pi^2} \frac{1}{3} \left(\frac{p_f}{\hbar} \right)^3 \left(\frac{3}{5} \frac{p_f^2}{2m} \right) \cdot V$$

$$\left(\frac{2\pi^2 N}{g V} \right)^{1/3} = \frac{p_f}{\hbar}$$

$$E = V C_E p_f^5$$

$$N = V C_N p_f^3$$

Then

$$\Delta E_{\text{Tot}} = \Delta E_P + \Delta E_N = 2 \left[\frac{1}{2} \frac{\partial^2 E}{\partial p_f^2} (\Delta p_f)^2 + \frac{1}{2} \frac{\partial^2 E}{\partial p_f^2} (\Delta p_f)^2 \right]$$
$$= \frac{\partial^2 E}{\partial p_f^2} (\Delta p_f)^2$$

$$E = V \cdot C_E p_f^5$$

$$C_E = \frac{g}{2\pi^2} \frac{1}{5} \cdot \frac{1}{2m} \frac{1}{h^3}$$

$$\frac{\partial^2 E}{\partial p_f^2} = 5 \cdot 4 \cdot V C_E p_f^3$$

$$\frac{\Delta N}{\Delta p_f} = \frac{\partial N}{\partial p_f} \Delta p_f = 3 C_N \cdot V p_f^2 \Delta p_f$$

$$\Delta E_{\text{Tot}} = 20 C_E \cdot V p_f^3 \cdot \left[\frac{\Delta N}{(3 C_N \cdot V \cdot p_f^2)} \right]^2$$

$$\Delta E_{\text{Tot}} = \frac{20 C_E \cdot V}{9 C_N^2 V^2} \frac{1}{p_f} (\Delta N)^2 \quad p_f =$$

$$\Delta E_{\text{Tot}} = \frac{20 C_E}{9 C_N^2} \frac{1}{(p_f)} \frac{(\Delta N)^2}{V}$$

↘ Const

$$\Delta N = \left(\frac{N-Z}{2} \right)$$

$$V = \frac{4\pi}{3} R^3$$

$$V = \rho_0 \frac{A}{\rho_0}$$

$$\text{So } \Delta E_{\text{TOT}} = \frac{20}{9} \frac{C_E}{C_N^2} \frac{1}{\rho_f} \frac{1}{\rho_0} \frac{1}{4} \cdot \frac{(N-Z)^2}{\frac{A}{\rho_0}}$$

$$C_N \rho_f^3 = \frac{\rho_0}{2} \rightarrow$$

$$\Delta E_{\text{TOT}} = \frac{20}{9} \frac{C_E}{C_N^2} \cdot \frac{1}{\rho_f} \cdot \frac{1}{4} \overbrace{\rho_0}^{2C_N \rho_f^3} \frac{(N-Z)^2}{A}$$

$$\Delta E_{\text{TOT}} = \frac{10}{9} \frac{C_E}{C_N} \rho_f^2 \frac{(N-Z)^2}{A}$$

$$\Delta E_{\text{TOT}} = \frac{10}{9} \underbrace{\frac{\rho_f^2}{2m}} \frac{(N-Z)^2}{A}$$

$\sim 25 \text{ MeV}$

from before

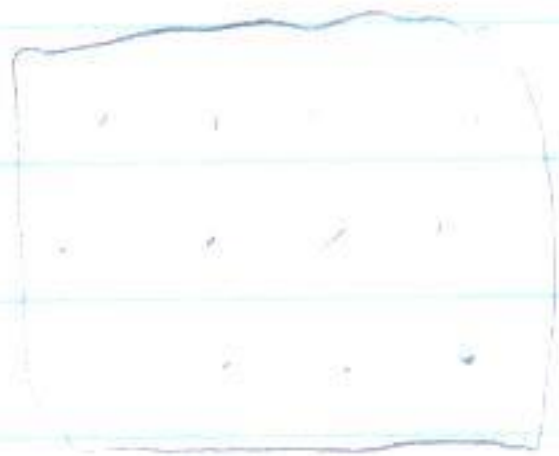
$$\Delta E \sim 27 \text{ MeV} \left(\frac{N-Z}{A} \right)^2$$

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