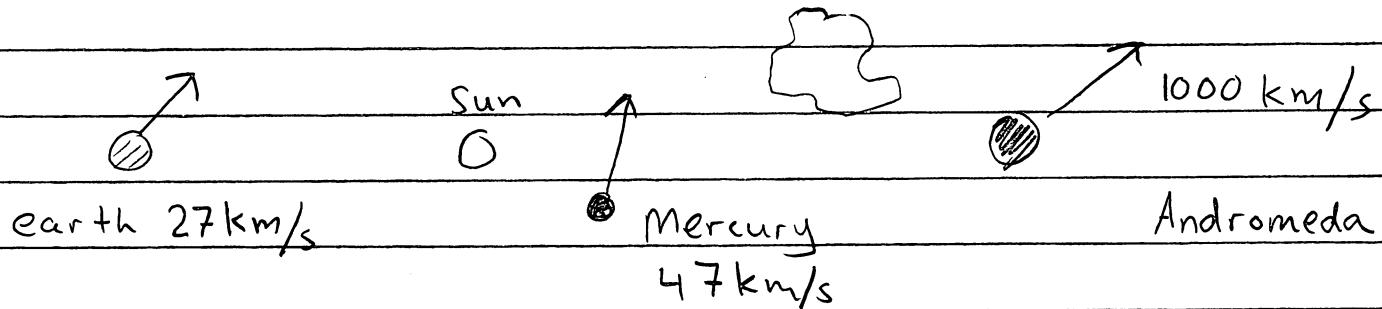


Relativity

Motivation for relativity. The universe is in constant motion

E+M Process



All observers are in constant motion relative to each other. Why is any single observer preferred? Yet they all seem to measure different forces in the E+M processes:

$$\vec{F} = q(\vec{E} + \frac{\vec{v}_p}{c} \times \vec{B})$$

they all measure
a different $\vec{v}_p \equiv$ the
particles velocity

Let's write the equations of E+M relativistically:

$$\text{Earth} \quad \frac{d\vec{p}}{dt} = q(\vec{E} + \frac{\vec{v}_p}{c} \times \vec{B})$$

The only thing
relativity changes
is the relation

$$-\nabla \cdot \vec{E} = \rho$$

between the velocity
and momentum:

$$\nabla \times \vec{B} = \vec{j}/c + \frac{1}{c} \partial_t \vec{E}$$

$$\vec{v}_p = c \frac{\vec{p}}{\sqrt{\vec{p}^2 + (mc)^2}}$$

$$\nabla \cdot \vec{B} = 0$$

$$-\nabla \times \vec{E} = \frac{1}{c} \partial_t \vec{B}$$

Then an observer on Andromeda measures the same rules, with his own quantities

$$\underline{t}, \underline{x}, \vec{\underline{p}}, \underline{E}, \underline{B}, \vec{\underline{J}}, \rho$$

Andromeda

$$\frac{d\vec{\underline{p}}}{dt} = q(\vec{\underline{E}} + \frac{\vec{\underline{v}}_p \times \underline{E}}{c})$$

$$-\nabla \cdot \underline{E} = \rho$$

$$\nabla \times \underline{B} = \underline{J}/c + \frac{1}{c} \partial_t \underline{E}$$

$$\nabla \cdot \underline{B} = 0$$

$$-\nabla \times \underline{E} = \frac{1}{c} \partial_t \underline{B}$$

$$\underline{v}_p = \frac{\underline{P}}{\sqrt{P^2 + (mc)^2}}$$

The speed of Light is constant in all frames / observers.

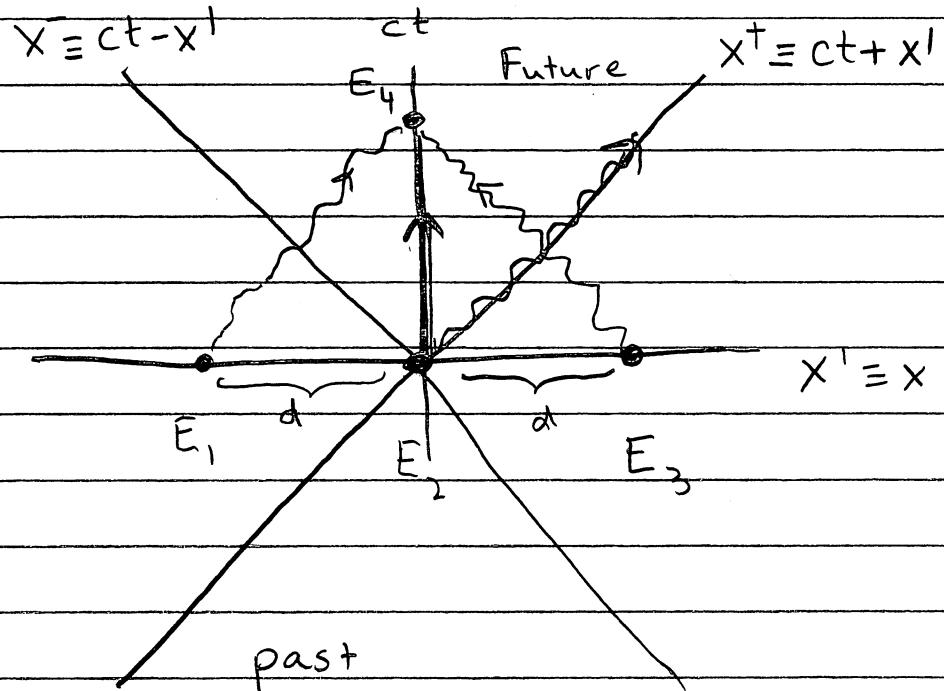
Relativity relates the unbarred quantities (the earth observer) to the barred quantities (the Andromeda observer)

Frames, Events, and Coordinates

Each observer sets up his own coordinate system (K) + labels the events that happen in space-time by these coords

$$(X^\mu) = (x^c, x^1, x^2, x^3) \equiv (ct, \vec{x}) \quad \mu=0,1,2,3$$

Thus $X^0 = ct$ labels when events happen, while x^i labels where events happen. We can record where and when events happen on a space-time diagram. Consider the following:



which puts the time (well ct , actually) of the events on the y -axis and spatial coordinates on the x -axis

Here we show a number of events.

- (1) E_1 = Light released at $x = -d$ travelling to right.
- (2) E_2 = A particle born at rest (thick line), and light born travelling to right along the line,
 $x^- \equiv c^0 t - x = 0$, with increasing, $x^+ \equiv ct + x$.
- (3) E_3 = Light born at $x = d$ moving to left (with increasing $x^- \equiv ct - x$ and constant x^+).

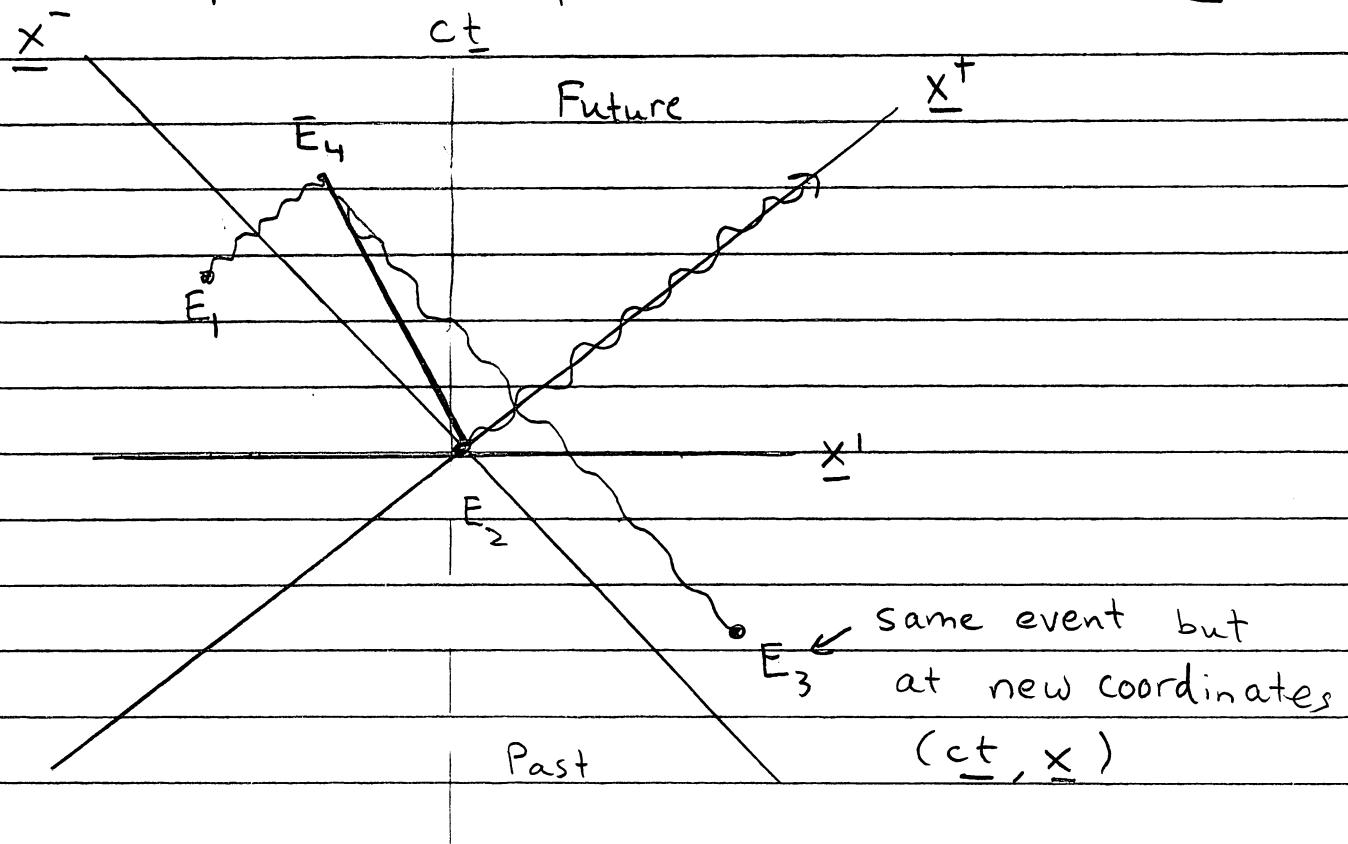
Light rays travel along 45° lines with fixed
 $x^+ \equiv ct + x$ or $x^- \equiv ct - x$. At E_4 the particle & rays meet.

An observer \underline{K} moving to right with velocity v relative to K , measures the particle moving to left with velocity v . All light rays still move with speed of light = 45° lines. The new \underline{K} coordinates are related by Lorentz transformation.

As we will show, the effect of the Lorentz transformation is to (for a boost to the right) contract the x^+ coordinate, and elongate the x^- coordinate for all events. See below:

$$\underline{x}^+ = \sqrt{\frac{1-\beta}{1+\beta}} x^+ \quad \underline{x}^- = \sqrt{\frac{1+\beta}{1-\beta}} x^-$$

So the space-time picture looks like in \underline{K} :



So in this frame E_3 happens first, then E_2 , then E_1 . E_4 is causally connected with E_1, E_2, E_3 and therefore must happen after E_1, E_2, E_3 in all frames. But, E_1, E_2, E_3 are not causally connected and can appear in various orders.

We seek a change of coordinates which leave the trajectory of light fixed, $c = x/t$

$$-(ct)^2 + \underline{x}^2 = -ct^2 + \underline{x}^2$$

i.e. it is the same for both observers

So $x^m \rightarrow \underline{x}^n = L^m(v) \underline{x}^v$, or as matrices

$$\begin{pmatrix} \underline{x}^0 \\ \underline{x}^1 \\ \underline{x}^2 \\ \underline{x}^3 \end{pmatrix} = L(v) \begin{pmatrix} x^0 \\ x^1 \\ x^2 \\ x^3 \end{pmatrix}$$

Properties $L(-\vec{v})L(\vec{v}) = I$

$$L(v_2)L(v_1) = L(v_3)$$

This is known as a group of transformations. The Lorentz Group. With these properties find, for v in \vec{x}^1 direction

$$\begin{pmatrix} x^0 \\ x^1 \\ x^2 \\ x^3 \end{pmatrix} = \underbrace{\begin{pmatrix} \gamma & -\gamma\beta & 0 & 0 \\ -\gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}}_{\text{defines } L^m} \begin{pmatrix} x^0 \\ x^1 \\ x^2 \\ x^3 \end{pmatrix}$$

$\gamma = \frac{1}{\sqrt{1-\beta^2}}$

$\beta = \frac{v}{c}$

(in general use vectors to express boosts in a general direction)

Often use a parameter y (the rapidity) to parametrize the boost matrix instead of γ to parametrize the boost

$$\frac{v}{c} = \tanh y \rightarrow y = \tanh^{-1} v_p = \frac{1}{2} \ln \left(\frac{1+\beta}{1-\beta} \right)$$

Then

$\approx \beta$ for small β

$$\gamma = \cosh y \quad \text{and}$$

$$\gamma\beta = \sinh y$$

The Lorentz boost is a hyperbolic rotation

$$\begin{pmatrix} x^0 \\ x^1 \end{pmatrix} = \begin{pmatrix} \cosh y & -\sinh y \\ -\sinh y & \cosh y \end{pmatrix} \begin{pmatrix} x^0 \\ x^1 \end{pmatrix}$$

Exercises

- ① Show that the Lorentz boost compresses x^+ and expands x^- , by the factors of e^{-y} and e^{+y} .

$$\underline{x}^+ = \sqrt{\frac{1-\beta}{1+\beta}} x^+ \quad \underline{x}^- = \sqrt{\frac{1+\beta}{1-\beta}} x^-$$

$$\underline{x}^+ = e^{-y} x^+ \quad \underline{x}^- = e^{+y} x^-$$

Four Vectors and Tensors (Mechanics of Indices)

① The coordinates obey a transformation rule

$$\underline{x}^m = L^m \downarrow x^\nu$$

We call any set of four quantities which transform in this way a four vector. Upper indices are referred to as contravariant components

$$\underline{A}^m = L^m \downarrow A^\nu$$

$$\underline{B}^m = L^m \downarrow B^\nu$$

Any two four vectors which transform in this way have an invariant product, $A^{\tilde{m}} \equiv (a^0, \vec{a})$ and $B^{\tilde{m}} \equiv (b^0, \vec{b})$

$$A \cdot B = -a^0 b^0 + \vec{a} \cdot \vec{b} = \underline{A} \cdot \underline{B} = -\underline{a}^0 \underline{b}^0 + \underline{\vec{a}} \cdot \underline{\vec{b}}$$

because L was adjusted to preserve this quadratic form

② For any set of contravariant indices, define their covariant counterparts with the metric tensor:

$$X_\mu = (-ct, \vec{x})$$

$$X_\mu = g_{\mu\nu} X^\nu$$

$$g_{\mu\nu} = \begin{pmatrix} -1 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix}$$

Lowering indices just changes sign of 0-th component, Sim, we raise indices with $g^{\mu\nu}$:

$$x^\mu = g^{\mu\nu} x_\nu \quad g^{\mu\nu} = \begin{pmatrix} -1 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix}$$

So clearly

$$x^\mu = g^{\mu\nu} g_{\nu\sigma} x^\sigma \quad g^{\mu\sigma} = \begin{pmatrix} 1 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix}$$

$$\delta^\mu_\sigma = g^{\mu}_\sigma = \text{Identity matrix}$$

We define Covariant indices so the inner product, can now be written as

$$A \cdot B = A_\mu B^\mu = -a^\circ b^\circ + \vec{a} \cdot \vec{b}$$

(3) Note if $A_\mu A^\mu$ is to be invariant under Lorentz transformation, then since $\underline{A}^\mu = L^\mu_\nu A^\nu$, we define the transformation rule for lower indices (covariant components) with L^{-1} and as a row

$$\underline{A}_\nu = A_\mu (L^{-1})^\mu_\nu \quad \text{or} \quad \underline{A}_\nu = (L^{-1T})_\nu^\mu A_\mu$$

i.e. lower indices transform with inverse and as a row

$$(\underline{A}_0, \underline{A}_1, \underline{A}_2, \underline{A}_3) = (A_0, A_1, A_2, A_3) \begin{pmatrix} L^{-1} \end{pmatrix}$$

So that

$$\begin{aligned}\underline{A}_\mu \underline{B}^\mu &= (\underline{A}_0 \underline{A}_1 \underline{A}_2 \underline{A}_3) \begin{pmatrix} \underline{B}^0 \\ \underline{B}^1 \\ \underline{B}^2 \\ \underline{B}^3 \end{pmatrix} \\ &= (\underline{A}_0 \underline{A}_1 \underline{A}_2 \underline{A}_3) (L^{-1}) (L) \begin{pmatrix} \underline{B}^0 \\ \underline{B}^1 \\ \underline{B}^2 \\ \underline{B}^3 \end{pmatrix} \\ &= \underline{A}_\mu \underline{B}^\mu\end{aligned}$$

(4) Note that under Lorentz transform

$$\begin{aligned}\underline{A} \cdot \underline{B} &= \underline{A}^\mu g_{\mu\nu} \underline{B}^\nu = -\underline{a}^0 \underline{b}^0 + \underline{\vec{a}} \cdot \underline{\vec{b}} \\ &= \underline{A}^\mu g_{\mu\nu} \underline{B}^\nu = -\underline{a}^0 \underline{b}^0 + \underline{\vec{a}} \cdot \underline{\vec{b}} = A \cdot B\end{aligned}$$

Without the need of transforming $g_{\mu\nu}$. This says that $g_{\mu\nu}$ is an invariant tensor

$$(L^{-1T})_\mu{}^\rho g_{\rho\sigma} L^{-1\sigma}{}_\nu = g_{\mu\nu}$$

Or in matrices:

$$(L^{-1T}) g L^{-1} = g$$

i.e.

$$L^{-1T} = g L g \quad (\text{note } g^{-1} = g)$$

Restoring indices

$$g_{\mu\nu} (L^\nu)_\rho g^{\rho\sigma} = (L^{-1T})_\mu^\sigma$$

In a fit of notational madness (which is standard) we define

$$\boxed{L_\mu^\sigma = g_{\mu\nu} L^\nu_\rho g^{\rho\sigma} = (L^{-1T})_\mu^\sigma}$$

So perhaps its not so mad ...

$$\underline{A}_\mu = A_\nu (L^{-1})^\nu_\mu$$

$$= (L^{-1T})_\nu^\mu A_\mu$$

$$\boxed{\underline{A}_\mu = L_\nu^\mu A_\mu}$$

Exercise

- A tensor transforms as

$$\underline{T}^{\mu\nu} = L^\mu{}_p L^\nu{}_\sigma T^p{}^\sigma$$

Show that the transformation rule for $\underline{T}^m{}_v$ is

$$\underline{T}^m{}_v = L^m{}_p T^p{}_\sigma (L^{-1})^\sigma{}_v$$

or equivalently

$$\underline{T}^m{}_v = L^m{}_p L^\sigma{}_v T^p{}_\sigma$$

Solution:

- Lower v ,

$$\underline{T}^m{}_v = L^m{}_p L_{v\sigma} T^p{}^\sigma$$

$$\underline{T}^m{}_v = L^m{}_p L^\sigma{}_v T^p{}_\sigma$$

$$= L^m{}_p T^p{}_\sigma (L^{-1}T)^\sigma{}_v$$

$$= L^m{}_p T^p{}_\sigma (L^{-1})^\sigma{}_v$$

Exercise :

- Explain that if a plane wave of light $e^{-i\omega t + \vec{k} \cdot \vec{x}}$ where $k = \frac{\omega}{c}$, is to move at

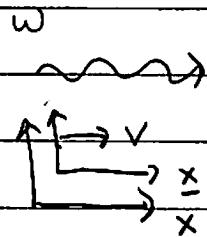
the speed of Light in all frames, then

$$k^{\mu} = \left(\frac{\omega}{c}, \vec{k} \right)$$

must be a four vector. ① Show that

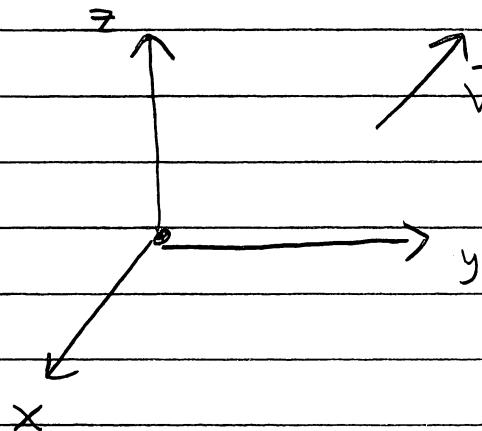
$\vec{k} \cdot \vec{k} = 0$. ②) and show the relativistic doppler shift formula is

$$\underline{\omega} = \sqrt{\frac{1-\beta}{1+\beta}} \omega$$

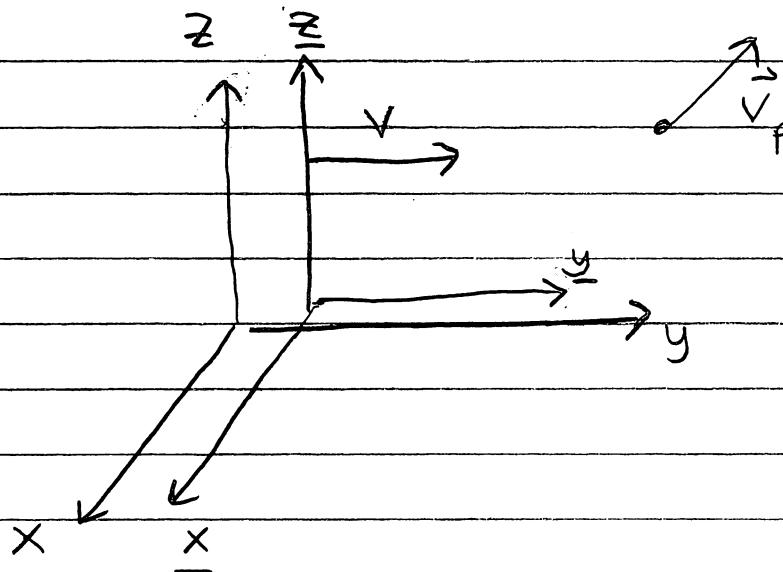


Particles In Special Relativity

Consider a particle with velocity $\vec{\beta}_p = \vec{v}_p/c$. Note we put the "p" sub-label to keep it apart from β and v which will label the velocity of a new frame/observer that we wish to boost to.



We will calculate \vec{v}_p in another frame below. The new frame moves with velocity \vec{v} relative to the first, and has its own coordinate $(ct, \underline{x}, \underline{y}, \underline{z})$ shown below



- The change in coordinates of the particle is

$$d\vec{x}^m = (c dt, d\vec{x}) = \underbrace{dx^0}_{= c dt} (1, \vec{\beta}_p)$$

where

$$\vec{\beta}_p = \frac{d\vec{x}}{dx^0} = \frac{\vec{v}_p}{c}$$

Then we can construct the invariant space-time interval

$$ds^2 = dx_\mu dx^\mu = -c^2 dt^2 (1 - \beta_p^2) \quad (\star)$$

Since it is invariant we can interpret it in any frame. In the rest frame of the particle, the particle is not moving, $\beta_p = 0$, its time increases:

$$dt \text{ in Local Rest Frame} = d\tau = \text{proper time}$$

So

$$ds^2 = -c^2 d\tau^2$$

And

$$d\tau = \sqrt{-ds^2} = \frac{1}{c} \sqrt{-dx_\mu dx^\mu}$$

Then we have from Eq *

$$d\tau = dt \sqrt{1 - \beta_p^2}$$

$$\boxed{d\tau = \frac{dt}{\gamma_p}}$$

$$\gamma_p \equiv \frac{1}{\sqrt{1 - \beta_p^2}}$$

Then we define the four velocity u^m as

$$\boxed{u^m \equiv \frac{dx^m}{d\tau}}$$

Now

$$dx^m = dt (c, \frac{d\vec{x}}{dt}), \quad \text{or}$$

$$dx^m = \gamma_p d\tau (c, \vec{v}_p).$$

So

$$\boxed{u^m = (\gamma_p c, \gamma_p \vec{v}_p)} \quad (\star\star)$$

$$\text{Clearly } u^m u_m = -c^2$$

$$u^m u_m = \frac{dx^m}{d\tau} \frac{dx_m}{d\tau} = \frac{ds^2}{(d\tau)^2} = -c^2$$

You can also verify directly from Eq $(\star\star)$

Energy + Momentum

Consider the action of a particle

$$S = \int_{t_1}^{t_2} L dt = \int_{t_1}^{t_2} (p \dot{q} - H(p, q)) dt$$

So for a free particle $\vec{p} = \text{const}$ $H = E$ and we have ;

$$S = \int p d\vec{x} - E dt$$

So since $(cdt, d\vec{x})$ is a four vector it is theoretically tempting to define a four vector

$$P^\mu = \left(\frac{E}{c}, \vec{p} \right)$$

Then $P \cdot dX = P_\mu dx^\mu = -E dt + \vec{p} \cdot d\vec{x}$ is Lorentz invariant. Requiring at small velocities $\vec{p} = m\vec{v}_p$ leads to

$$P^\mu = mu^\mu = (mc\gamma_p, m\gamma_p v_p) = \left(\frac{E}{c}, \vec{p} \right)$$

In particular we see that the energy of a particle is $E = \gamma_p mc^2$

Energy and Momentum Conservation

Consider the reaction $1 + 2 \leftrightarrow 3 + 4$

Energy and momentum are conserved

$$P_1^m + P_2^m = P_3^m + P_4^m \quad (\star)$$

In addition the energies of each of the particles are related to their masses.

In working with collisions we use a shorthand notation setting $c=1$. With this we have, for example:

$$P^2 = P_m^m = -m^2, \quad E = \sqrt{P^2 + m^2},$$

$$V = \frac{\vec{P}}{E} \quad \text{etc.}$$

One constraint on the reaction kinematics is found by squaring Eq. \star , e.g.

$$(P_1 + P_2)^2 = P_1^2 + P_2^2 + 2P_1 \cdot P_2$$

$$= -m_1^2 - m_2^2 + (-2E_1 E_2 + p_1 p_2 \cos \theta)$$

So

Or squaring both sides

$$m_1^2 + m_2^2 + 2E_1 E_2 - \vec{p}_1 \cdot \vec{p}_2 =$$

$$m_3^2 + m_4^2 + 2E_3 E_4 - \vec{p}_3 \cdot \vec{p}_4$$

This is only one constraint. In general one needs to work fairly hard, writing (Eq \star)

$$E_1 + E_2 = E_3 + E_4$$

$$\vec{p}_1 + \vec{p}_2 = \vec{p}_3 + \vec{p}_4$$

in components and solving the equations to find the relations between energy and momenta between the particles.