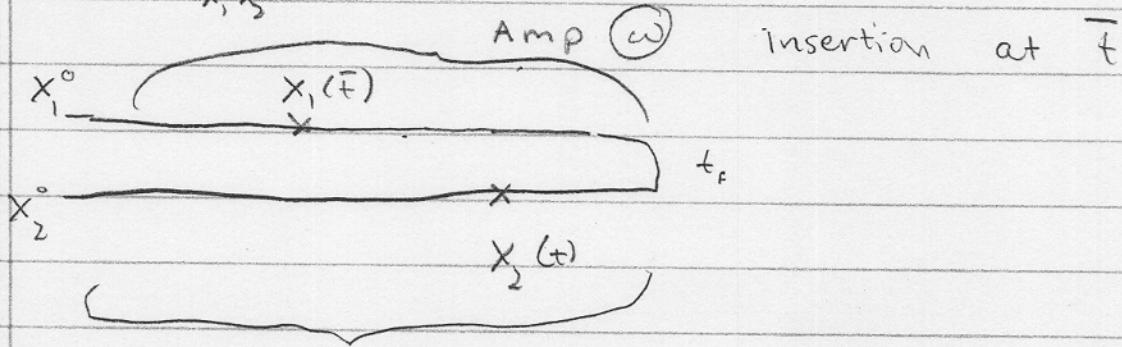


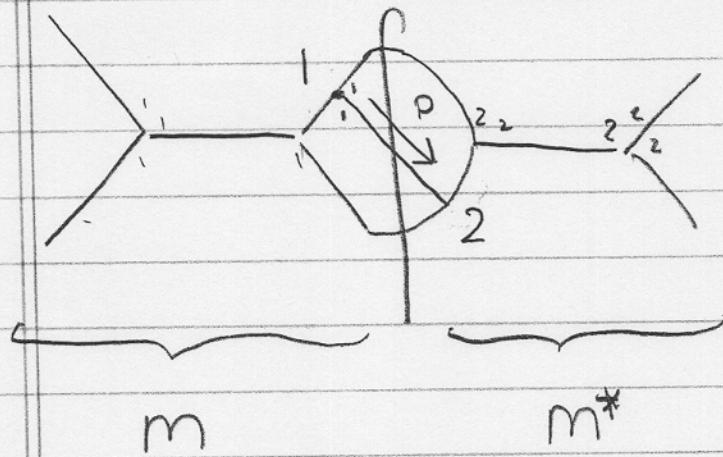
Last Time:

$$\int dx_1^0 dx_2^0 \rho \int_{x_1^0, x_2^0} D\bar{x}_1 D\bar{x}_2 e^{iS_1 - iS_2} x_2(\bar{t}) x_1(\bar{t})$$



So

Example:  $T=0$



Cut Lines:

$$\Theta(p^0) 2\pi \delta(p^2 + m^2) \\ = \frac{2\pi}{2E_p} \delta(p^0 - E_p)$$

$$G^>(p) = G_{21}(p) = \int e^{ip \cdot x} \langle \hat{\phi}(x) \hat{\phi}(0) \rangle$$

Use  $\hat{\phi}(x) = \sum_p a_p \frac{e^{ip \cdot x}}{\sqrt{2E_p}} + a_p^+ \frac{e^{-ip \cdot x}}{\sqrt{2E_p}}$

Find

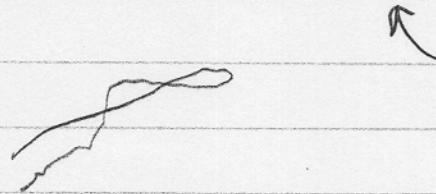
$$G^> = (1 + n_p) \frac{2\pi}{2E_p} \delta(p^0 - E_p) + n_{-p} \frac{2\pi}{2E_p} \delta(p^0 + E_p)$$

$\Rightarrow$   $T=0$   $\frac{2\pi}{2E_p} \delta(p^0 - E_p)$

↑ absorption of particle

Then I introduced:

$$x_r = \frac{x_1 + x_2}{2}, \quad x_a = x_1 - x_2$$



small in classical limit

Then saw good things

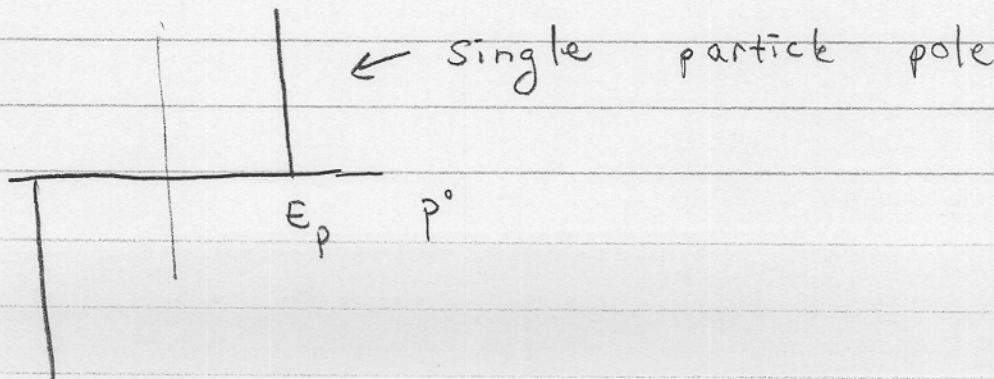
$$\underset{a}{\overbrace{\dots}} \rightarrow iG_{ra} = \langle x_r x_a \rangle = \Theta(t) \langle [\hat{x}(t), \hat{x}(t)] \rangle$$

$$iG_{ra} = \frac{-i}{-\omega^2 + \omega_0^2 - i\epsilon\omega_0} \stackrel{FT \text{ theory}}{\Rightarrow} \frac{-i}{-(p^0)^2 + E_p^2 - i\epsilon p^0} = \frac{-i}{p^2 + m^2 - i\epsilon p^0}$$

Spectral Density:

$$\rho(\omega) = -2 \operatorname{Im} G_R(p) = \frac{2\pi}{2E_p} \delta(p^0 - E_p) - \frac{2\pi}{2E_p} \delta(p^0 + E_p)$$

So

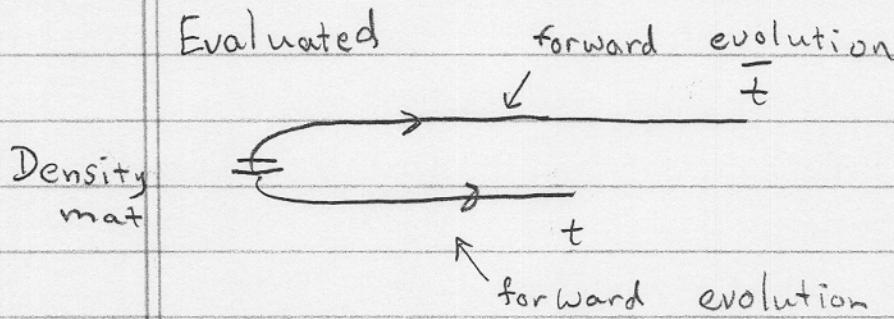


We also evaluated:

$$G_{rr} = \langle x_r(t) x_r(\bar{t}) \rangle = \frac{1}{2} \langle \{x(t), x(\bar{t})\} \rangle$$

Find

$$G_{rr}(P) = \left( \frac{1}{2} + n(P^0) \right) \rho(P^0)$$

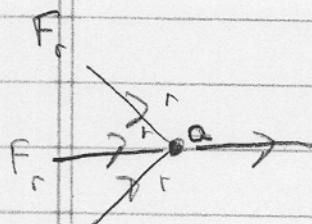


Hard to draw in Feynman Diagrams:

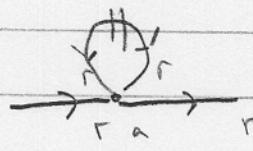
$$\overline{r} \leftarrow \parallel \rightarrow r \equiv G_{rr}(\omega) = \left( \frac{1}{2} + n(P^0) \right) \rho(P)$$

## Interactions:

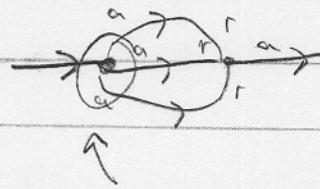
$$\frac{1}{4!} x_1^4 - \frac{1}{4!} x_2^4 = \underbrace{\frac{1}{3!} x_r^3 x_a}_\text{Classical} + \underbrace{\frac{1}{3} x_a^3 x_r}_\text{Quantum}$$



non linear response  
classical



Thermal vacuum  
fluctuation contribution  
to retarded response



Quantum Correction to  $G_R$

# Deriving Boltzmann: Free theory

Yesterday we showed that the symmetrized

$$\left( -\frac{\partial^2}{\partial x^2} + m^2 \right) G_{rr}(x, y) = 0$$

$$G_{rr}(x, y) \left( -\frac{\partial^2}{\partial y^2} + m^2 \right) = 0$$

Now introduce a Wigner transform

$$G(\bar{x}, p) = \int ds e^{-ip \cdot s} G(x + \frac{s}{2}, \bar{x} - \frac{s}{2})$$

So

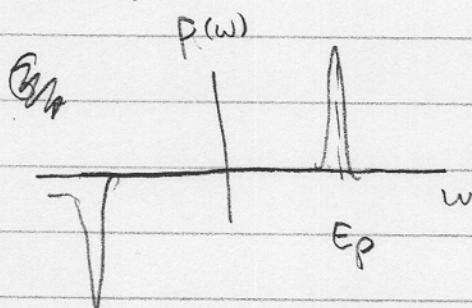
$$\left. \begin{aligned} & \left( \frac{\partial_x^2 + \partial_y^2}{2} + m^2 \right) G(x, y) = 0 \\ & (\partial_x^2 - \partial_y^2) G(x, y) = 0 \end{aligned} \right\} \begin{aligned} & (p^2 + m^2) G(x, p) = 0 \\ & 2p^m \frac{\partial}{\partial x^m} G(x, p) = 0 \end{aligned}$$

Recall free:

$$G_{rr}(p) = \left( \frac{1}{2} + n(p^0) \right) \left\{ \frac{2\pi}{2E_p} \delta(p^0 - E_p) - \frac{2\pi}{2E_p} \delta(p^0 + E_p) \right\}$$

$\underbrace{\qquad\qquad\qquad}_{\rho(\omega)}$

Non-equilibrium



assume

$$G_{rr} \approx \left( \frac{1}{2} + n_p(x) \right) \frac{2\pi \delta(p^o - E_p)}{2E_p} + \left( \frac{1}{2} + \bar{n}_p(x) \right) \frac{2\pi \delta(p^o + E_p)}{2E_p}$$

i.e.

$$\frac{1}{2} + \frac{n_p(x)}{2E_p} \approx \int \frac{dp^o}{2\pi} G_{rr}(x, p)$$

$E_p + \text{bit}$

Free streaming

So

$$\frac{2p^o}{2E_p} \frac{\partial}{\partial x^\mu} n_p(x) = 0 \Rightarrow \left[ \left( \partial_t + v_p \frac{\partial}{\partial x} \right) n_p(x) \right] = 0$$

Then this

$$\left( \partial_t + v_p \frac{\partial}{\partial x} \right) \bar{n}_p = 0$$

wave number

- ① When momentum<sup>↑</sup> is large compared  
to homogeneities wave packets  
move along geodesics

Now Consider the unfree case

$$(-\omega_x^2 + m^2) G_{ss'}(x, y) + \int_{ss'} \Pi_{ss'}(x, z) G_{ss'}(z, y) = \underline{1}_{ss'}$$

Eg.

$$G = \text{---} + \text{---} + \text{---} + \dots$$

$$G = \begin{pmatrix} G_{11} & G_{12} \\ G_{21} & G_{22} \end{pmatrix} \quad \Pi = \begin{pmatrix} \Pi_{11} & \Pi_{12} \\ \Pi_{21} & \Pi_{22} \end{pmatrix}$$

Find

$$G^> \sim (1+n) \frac{2\pi}{2\varepsilon_p} \delta(p^0 - \varepsilon_p)$$

$$2 p^m \frac{\partial}{\partial x^m} G^< = \Pi^> - \underbrace{G^< \Pi^>}_{\sim}$$

$$n_p \frac{2\pi}{2\varepsilon_p} \delta(p^0 - \varepsilon_p)$$

Skip

$$2 p^m \frac{\partial}{\partial x^m} n_p = -n_p \Pi^> + \Pi^< (1+n_p)$$

E. g. Photons

$$iM_{fi}(\vec{k}) = \langle \vec{E}_f | i \int_X J_\mu(x) A^\mu(x) | i \rangle$$

$$A_\mu(x) = \sum_{k\lambda} \frac{e^{ik \cdot x}}{\sqrt{2\varepsilon_k}} \epsilon_\mu^\lambda + e^{-ik \cdot x} \frac{a^\dagger}{\sqrt{2\varepsilon_p}} \epsilon_\mu^{\lambda*}$$

So

$$iM_{fi} = i\sqrt{1+n} \int_X \langle f | J_\mu(x) | i \rangle \frac{e^{-ik \cdot x}}{\sqrt{2\varepsilon_k}} \epsilon_\mu^{\lambda*}$$

$$(2\pi)^3 \frac{dP}{d^3k} = \frac{1}{Z} \sum_{fi} |M_{fi}|^2 e^{-\beta E_i}$$

Using

$$\int_X \int_Y = \int \overline{X} = \frac{x+y}{2} \int_{x-y}$$

We

$$\sum \epsilon_\nu \epsilon_\mu^* \rightarrow g_\mu$$

$$(2\pi)^3 \frac{dP}{d^3k} = \int d^4 \bar{X} \int_{x-y} \bar{e}^{iK(x-y)} \langle J_\nu(Y) J_\mu(X) \rangle \tilde{g}^\mu$$

$x \quad \frac{1+n_{ll}}{2\varepsilon_k}$

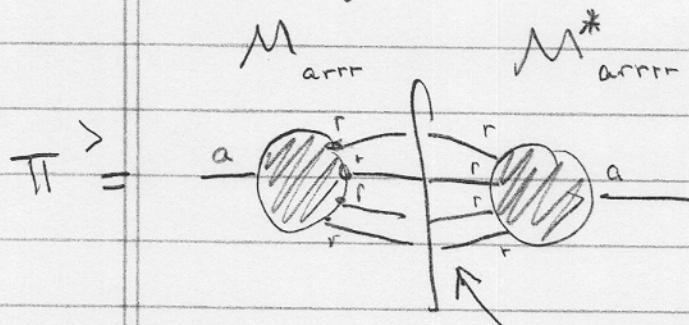
$$(2\pi)^3 \frac{d^3 p}{d^4 x d^3 k} = \frac{1}{2E_k} \int_{x-y} e^{i\vec{k}(x-y)} \langle J_\nu(y) J_\mu(x) \rangle g^{\mu\nu}$$

$\equiv \Pi^<(k)$

$\times (1 + n_k)$

$$(2\pi)^3 \frac{d^3 p}{d^4 x d^3 k} = \frac{\Pi^<(k)}{2E_k} \times (1 + n_k)$$

We have seen that  $\Pi^>$  and  $\Pi^<$  are squared matrix elements; (Simon-Caron-Huot)

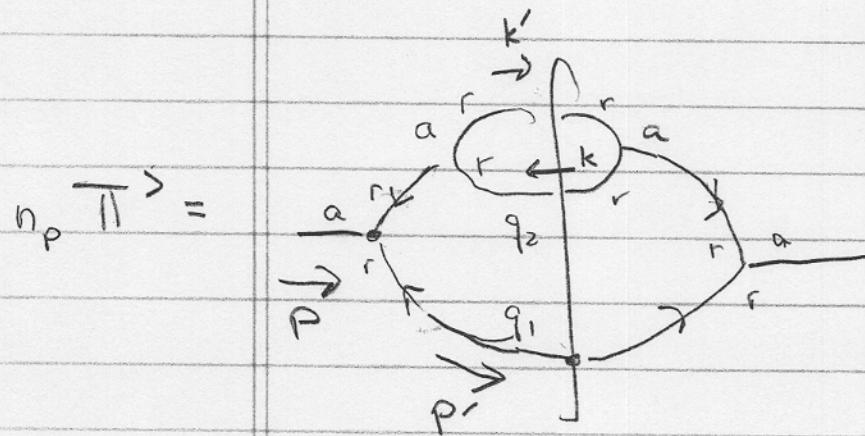


$M_{arrr}^*$

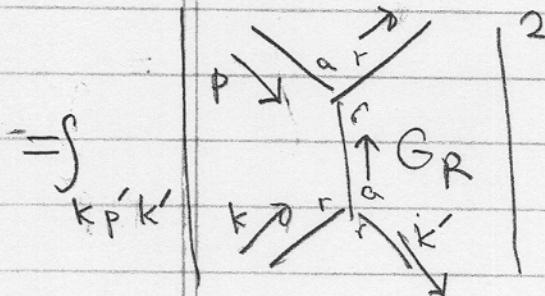
$\Pi^>$

$$\text{cut lines} = G^>(p) = 1 + n_p \frac{2\pi \delta(p^0 - E_p)}{2E}$$

$$+ n_{-p} \frac{2\pi \delta(p^0 + E_p)}{2E_p}$$



$n_p \Pi^>$



$\sim$  Feynman Graph  
But "exchange gluon"  
is retarded

$$\times n_p n_k (1 + n_{p'}) (1 + n_{k'})$$

QCD at last :

Use KT

$$\left[ \partial_t + v_p \frac{\partial}{\partial x} \right] f = C[f]$$

Scales:

Hard Particles:  $p \sim T$  almost onshell - Carry most

$$P^+ \sim T = P^0 + P^z \quad \text{of energy}$$

$$P^- \sim g^2 T = P^0 - P^z$$

$$P_\perp \sim g T \approx$$

$$P^2 = P^+ P^- + P_\perp^2 \sim O(g^2 T^2)$$

Soft Fields (Off shell)

$$P \sim g T$$

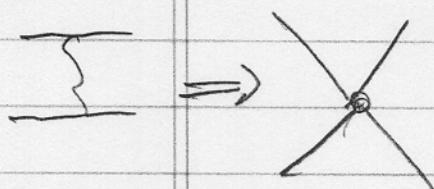
$$P^2 \sim g^2 T^2$$

Magnetic Sector  $\sim$  non perturbative

$$P \sim g^2 T$$

## Interactions:

① Hard Collisions  $\sim T$

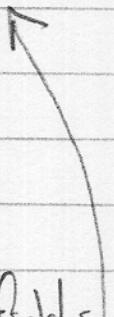


Randomizing Collisions

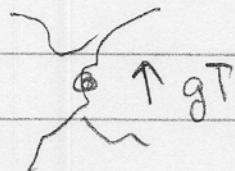
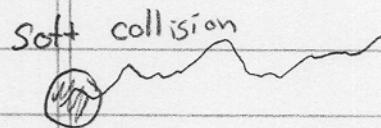
$$(\Delta p)^2 \sim \frac{t}{t_{\text{coll}}} T^2 \sim t T^3 g^4$$

$$t_d \sim \frac{1}{T}$$

$$t_{\text{coll}} \sim \frac{1}{g^4 T}$$



② Random Walk - Interactions ③ soft fields



$$t_d \sim \frac{1}{gT}$$

$$t_{\text{coll}} \sim \frac{1}{g^2 T}$$

Same

$$(\Delta p)^2 = \frac{t}{t_{\text{coll}}} (\delta p)^2 \approx \frac{t}{\frac{1}{g^2 T}} (gT)^2 \sim t g^4 T^3$$

### ③ Collinear Bremsstrahlung

Random Walk causes bremsstrahlung:

$\xrightarrow{1-x}$   
 $\xleftarrow{x}$

$\propto \sim g^{1/2} T$

Rate

$$\Gamma \sim \frac{\alpha}{t_{\text{soft}}} \sim \frac{g^2}{\gamma g^{1/2}} \sim g^4 T$$

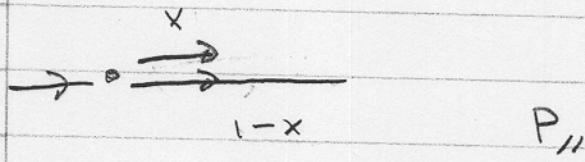
$$t_{\text{coll}} \sim \frac{1}{\Gamma} \sim \frac{1}{\alpha^2 T}$$

also  
same  
 $\downarrow$

So

$$(\Delta p)^2 \sim \frac{1}{t_{\text{coll}}} (\delta p)^2 \sim \frac{1}{\frac{1}{\alpha^2 T}} (T^2) \sim T^3 g^4 t$$

# Collinear Bremm : Formation Times



$$E_{\text{initial}} = \omega, k_{\perp} \quad \left. \begin{array}{l} \theta \sim gt \\ E' = E(1-x)P_{\parallel} \end{array} \right\} \text{take } x \text{ small} \quad \omega \approx xP$$

$$\Delta t \sim \frac{\hbar}{\Delta E} \sim \frac{\hbar}{E - (E' + \omega)}$$

$$E' = \sqrt{(1-x)^2 P_{\parallel}^2 + k_{\perp}^2} \approx (1-x)P_{\parallel} + \frac{k_{\perp}^2}{2P_{\parallel}(1-x)}$$

$$\omega = \sqrt{xP_{\parallel}^2 + k_{\perp}^2}$$

$$\approx xP_{\parallel} + \frac{k_{\perp}^2}{2xP_{\parallel}}$$

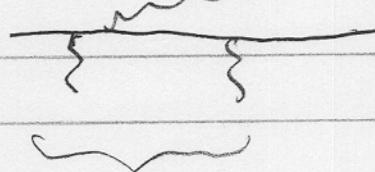
So

$$\Delta t_{\text{form}} \sim \frac{\hbar}{P_{\parallel} - \left[ (1-x)P_{\parallel} + \frac{k_{\perp}^2}{2P_{\parallel}(1-x)} + xP_{\parallel} + \frac{k_{\perp}^2}{2P_{\parallel}(x)} \right]}$$

$$\Delta t_{\text{form}} \approx \frac{2x(1-x)P_{\parallel}}{k_{\perp}^2} \underset{x \text{ small}}{\approx} \frac{\omega}{k_{\perp}^2}$$

$$\Delta t_{\text{form}} \sim \frac{I}{(gT)^2} \sim \frac{1}{g^2 T}$$

$$t_{\text{form}} \sim \frac{1}{g^2 T}$$

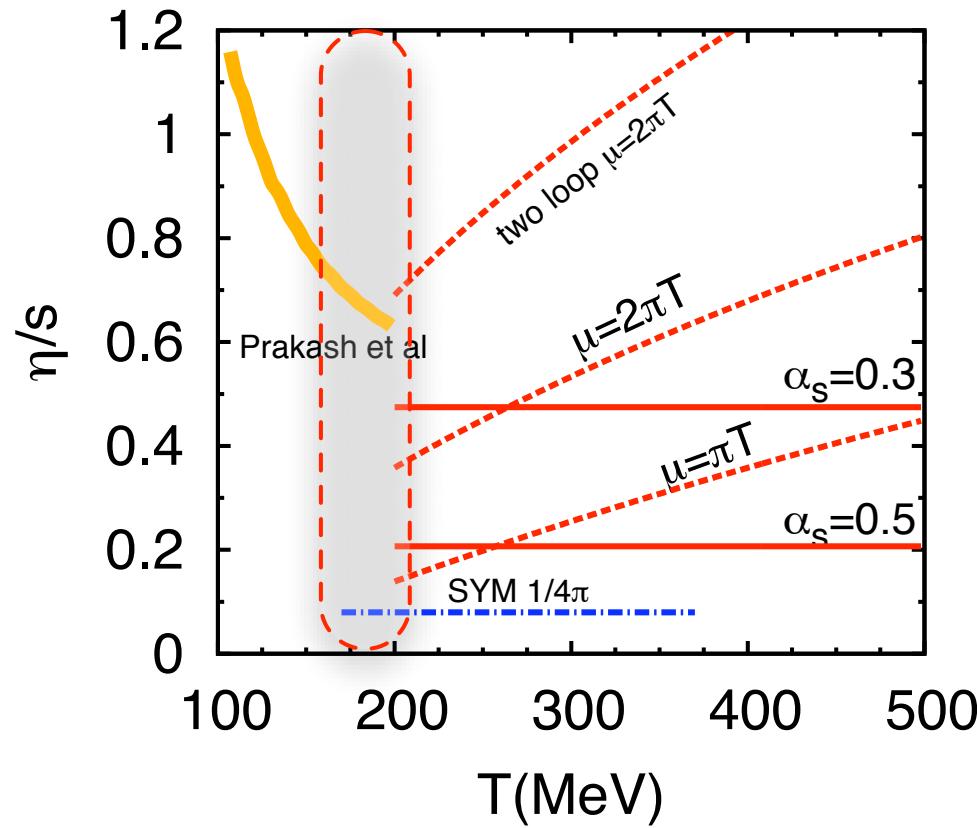


$$t_{\text{coll}} \sim \frac{1}{g^2 T}$$

So these two scattering events are phase coherent

- Need to solve a kind of Schrödinger Eqn to determine the emission rate

# Hadrons                    QGP



$$0.36 \left( \frac{\eta/s}{0.3} \right) \left( \frac{1 \text{ fm}}{\tau_o} \right) \left( \frac{300 \text{ MeV}}{T_o} \right) \ll 1$$